

# SKYMAPPER



and the state

### Australiar National University



the assessments





### It's a long way to the top of SSO 1 Jan 2003 - Our ARC DP grant starts.

E1 Title: The Southern Sky Survey

E2 Project Description, Aims and Background

**Project Description and Aims:** We propose to use the robotic Great Melbourne Telescope (GMT) to carry out an optical survey of the entire southern sky. The survey will be:

• Multi-Colour. We will observe at six wavelengths, from the near-UV to the near-IR

- 1. Photometric. We will measure the brightness of each object detected, at each wavelength, with systematic errors of less than 0.02 mag
- 2. Astrometric. We will determine the absolute positions of the objects we detect with an accuracy of better than 0.05 arcsec.
- 3. Sensitive to variability. Each par t of the sky will be observed multiple times, to look for tim e variability and movement.

The RSAA director has guaranteed that at least 80% of the Great Melbourne Telescope observing time will be devoted to this project, over 5 years. We will generate 25 terabytes of data and will detect more than 10<sup>9</sup> objects. All calibrated data will be made publicly available on-line via the ANU supercomputer facility.

We anticipate that the survey will be used for an enormous variety of scientific projects by astrophysicists worldwide for decades to come. The team members, however, are particularly interested in the following science goals:

- Mapping the distribution of dark matter in the outer regions of our own galaxy.
- Searching for high redshift QSOs to probe the reionisation of the universe.



### 20th January 2003







## 8 April 2003... Bushfire report to ARC

- Supervise construction of a wide field (approx 7sq degree field of view) 1.8 meter telescope at Siding Spring Observatory. (New Goal)
- 6. Supervise construction of a >100 million pixel CCD array for the Siding Spring Survey Telescope. (New Goal)

• 7. Start taking data for the Southern Sky



- Aug 2004
- Contract
  let for
  1.35m
  SkyMapper
  Telescope
  to EOS.









# What is SkyMapper?

- 1.35m telescope with a 5.7 sq. degree field of view
- Fully Autonomous observing
- To conduct the Southern Sky Survey:
  - Five year
  - Multi-colour (6 filters)
  - Multi-epoch (6 exposures, each filter)
  - 2π steradians
  - Limiting mag. g~23
- Aiming for regular operations this year
- Summary of program: Keller et al. 2007 PASA 24,1





Total Cost: Hardware \$15M Software: \$1.8M Dedicated Science/Operations: ~ \$0.6M/yr



#### **Telescope Optics**



- •The Mirror Crack'd 2005
  - •The sequel
  - The Mirror Crack'd Part III
  - •The Mirror Crack'd The Final Chapter
  - New Mirror supplier



#### 2006 - Russian Deliver more or less on time









#### • Bonn-Shutter - 2006 - arrives on time











#### \$2.5M of CCDs from E2V - arrive in 2006, on time







**SKYMAPPER** 



\$40,000 of filter glass from Russia...

Complete with with Cheese & Pepperoni



University





### Filters: u,r,z still complete

we await i filter to be finished in New Mexico Gabe is investigating issues in our v,g glass.









 The whole thing put together mid 2007, in Tucson









- First
  light,
  Tucson,
- July 2007

























#### **Telescope to SSO Sep 2008**





### Telescope handed over to ANU March 2009



# CCD Array -2011

- CCD Array has developed problems
- Saturation problems largely resolved - caused by short on pin
- Two chips failed caused by short inside dewar between two adjacent pins.
- Entire side has now failed due to power supply prob.
- RN in some channels has increased over time



# **CCD** Array

CCD Array delivered to telescope in 2010

Taken until 2013 to get all most of the issues sorted with it including

lower noise and fast cycle times a fault inside the dewar







Looks scary, but seems to be a very small ripple on the primary from milling which does not affect overall imaging of telescope. (pattern goes away at different focus values)



#### **AN OUT OF FOCUS IMAGE**

# Optics

Images cleaned up considerably by masking edge of primary Scattered light lowered through careful covering up of exposed white areas



# Optical Alignment

high-order pattern seems to be a nuisance, rather than a problem

Need more time to get optics aligned, but looks promising.

Central images limited by dome Temp



# Dome Temp

http://www.mso.anu.edu.au/metdata/

Plumbing issues preventing aircon from working perfectly. Longer term fix

We are currently generating 2kW in the dome because chilling system supplies too cold water to electronics and frosting/dew will occur. Short-term fix with second unit, will be addressed by longer term fix.

Dome leaks hot air around vents. Investigated next week with thermal camera.

Probably need addition Aircon unit.



File Edit View Frame Bin Zoom Scale Color Hegion WCS Analysis









## iversity 🛞







### SkyMapper Optimised for Stellar Astrophysics

- Encoded in the spectrum of each star
- Using filters we can isolate portions of the spectrum
- In designing our survey we sought to optimise our ability to determine the three important stellar parameters (T,log(g), Z)
- so SkyMapper not only compliments survey efforts in the northern hemisphere but enables us to tackle important astrophysics in an exciting new way.







### SkyMapper Filter Set is sensitive to Stellar Parameters







## **SkyMapper Filter Set**

Filter+CCD +Atmosphere







# The Southern Sky Survey

# (>=75% of all available time)

- 2π coverage: 5-sec survey. Photometric and astrometric Calibration of the Southern Sky in 6-filters (3-epochs per colour)
- 2π coverage: ~4000 fields observed in six filters, six times per filter to a (13-22 mag)
- Cadence: hours, days, weeks, months, years
- Photometry to 0.03(0.01 aspiration) mag globally (g>18.5)
- astrometry to 50 (15) mas
  - 36 images of each object over 5 years
  - → proper motions to ±5(2) mas/yr. (i.e. σvtan=25km/s at 2.5(6) kpc)
  - ⇒ parallax ±5(2) mas (i.e. 20pc (50)  $\sigma$ d=10%)
- survey complete in 5 years using 75% of telescope time





## Data Release

Deliverables to the Outside User:

-Data (epoch, RA, DEC, mags, galaxy shape info,...) to be available through a web-served interface which provides catalogs over a user defined area

–Images to be available through a web-served interface which provides images over a user defined area (1 degree max, cut out of a 2-degree TAN projection across the sky), or individual frames.

Data release will occur after extensive data validation:

-Five-Second data after closure in RA and trial application to concurrently obtained main-survey data

- -FDR Main Survey 3 epochs all filters
- -SDR Main Survey 6 epochs all filters



#### SKYMAPPER The Rapid Imaging Survey Era-May 2010

Name	Aperture (m)	FOV (sq deg)	Filter Set	Areal Coverage	Hemi sphere	First Light
SDSS	2.5	Drift scan	ugriz	π of 3/2π	Ν	Operating
CFHT MegaCam	3.6	1	ugriz	<1000	Ν	Operating
SkyMapper	1.35	5.7	uvgriz	2 π	S	2007
PanStarrs	1.8 (+3x)	7	grizY	3 π	Ν	2008
VISTA	4	1.65	zYJHK	2 π	S	2008
VST	2.6	1	ugriz	~5000	S	2008
Discovery Chn	4	2	?	?	N	2009?
Dark Energy	4	2	?	5000	S	2009?
LSST	8.4	10	ugrizY	3 π	S	2013







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