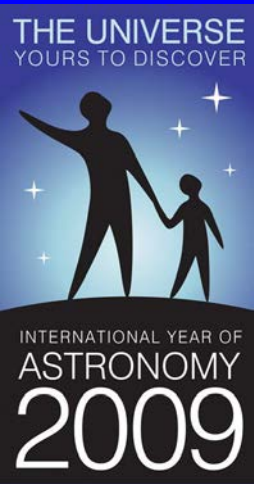


# The Hubble Space Telescope and the Hubble Constant



Jeremy Mould  
Centre for Astrophysics and Supercomputing  
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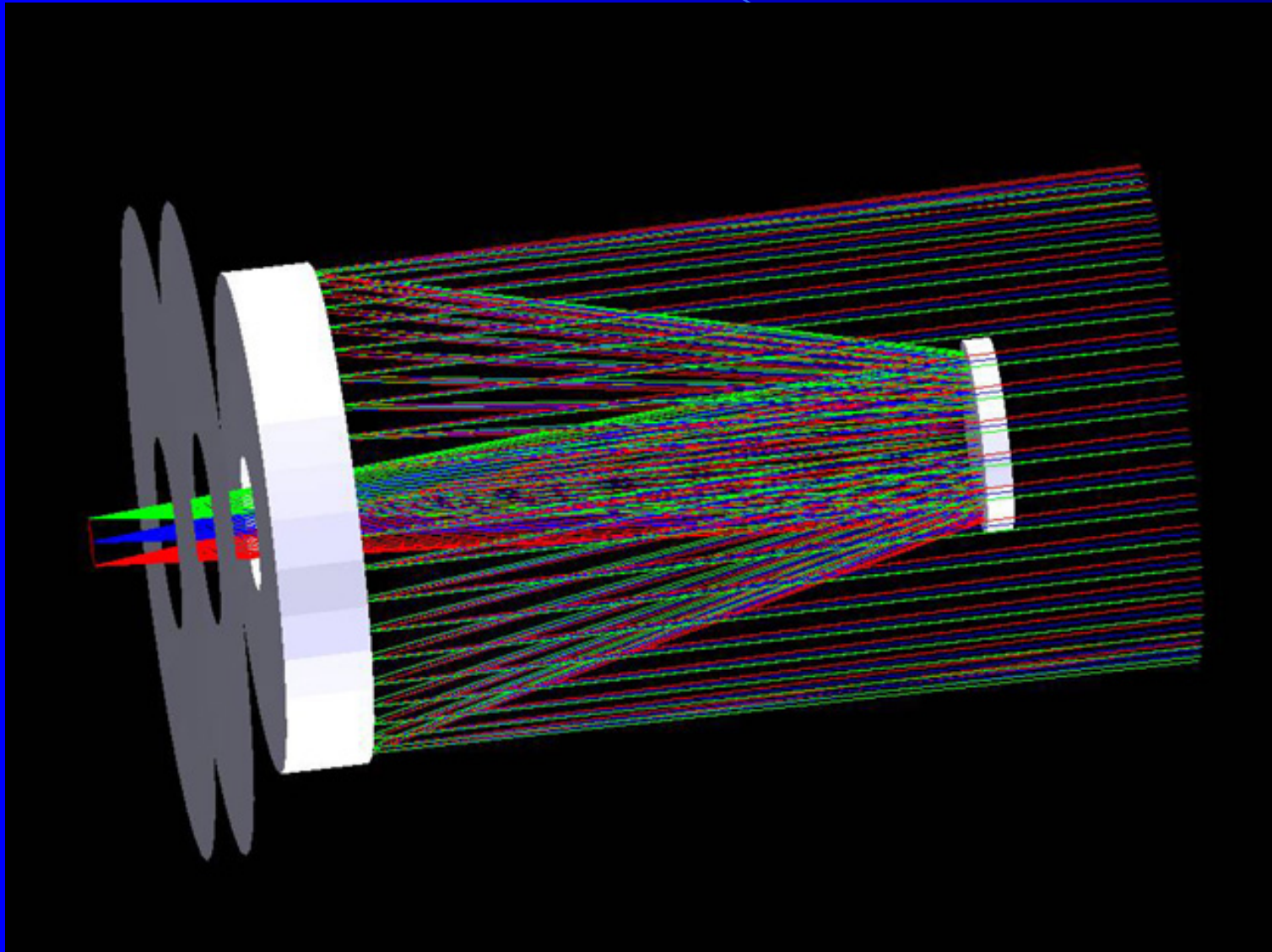


# Telescope resolution

Due to the wave nature of light (diffraction) the optical resolution of a telescope or camera is  $\lambda/D$

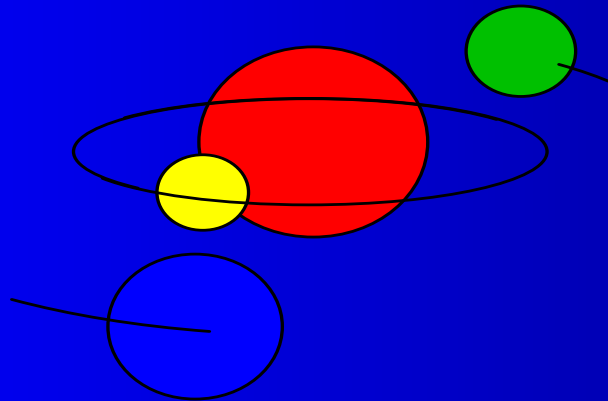
- The human eye: 20 arcsec
- Galileo's telescope: 3.8 arcsec
- Hubble Space Telescope 40 milliarcsecs

# Ritchey Chrétien optical design



# Introduction

- Subject today is the Expanding Universe
- The Expanding Universe model describes the motions of galaxies
- Just like the model of Copernicus describes the motions of the planets



# Outline

- Hubble's evidence for an expanding Universe
- How fast is the Universe expanding now ?  
(the Hubble Constant)
- Variation of the expansion rate over time
- How long is it since the Big Bang ?



Edwin  
Hubble

1889

-

1953



# Hubble's evidence

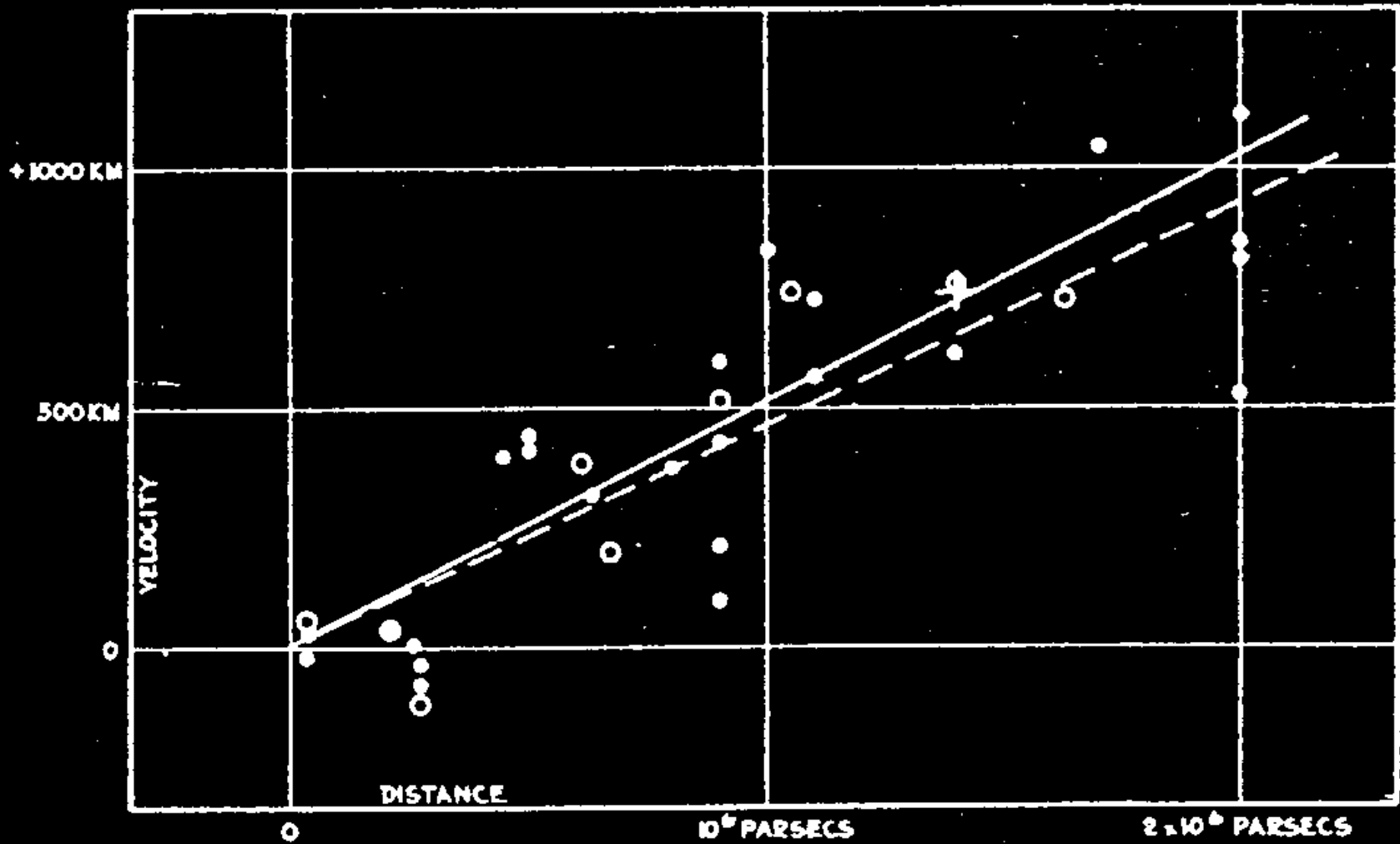
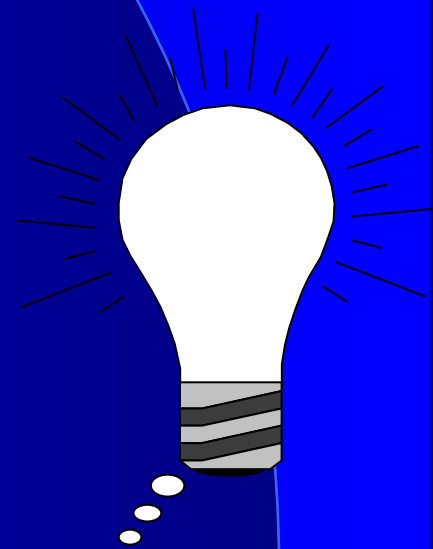


FIGURE 1

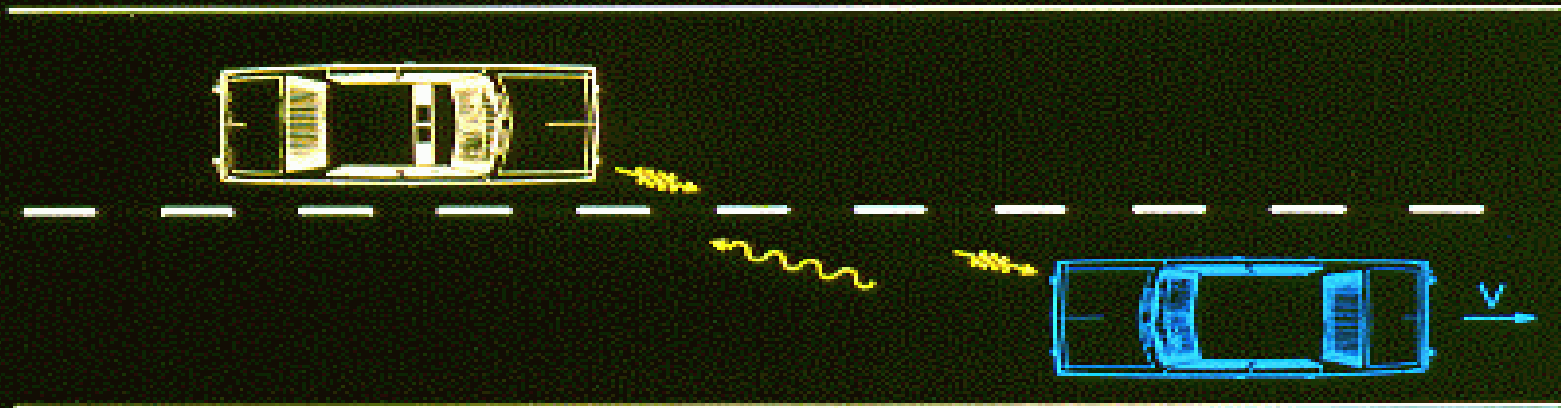
# Overview on the Hubble Constant

- what is **redshift** ?
- measuring distances geometrically
- we start with the Large Magellanic Cloud
- the Hubble Space Telescope Key Project
- Cepheids as standard candles
- Supernovae as standard candles





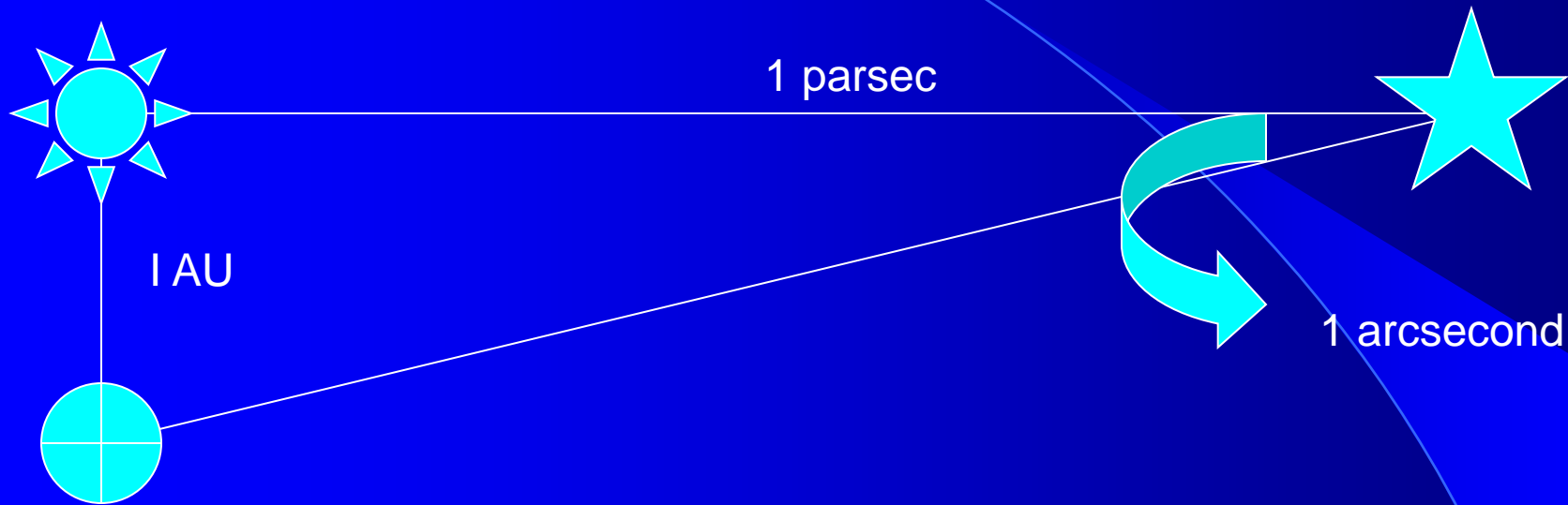
## REDSHIFT



$$\frac{\Delta\lambda}{\lambda} = \frac{v}{c}$$

$$c = 3 \times 10^5 \text{ km/sec}$$

# Parallax distance measurement

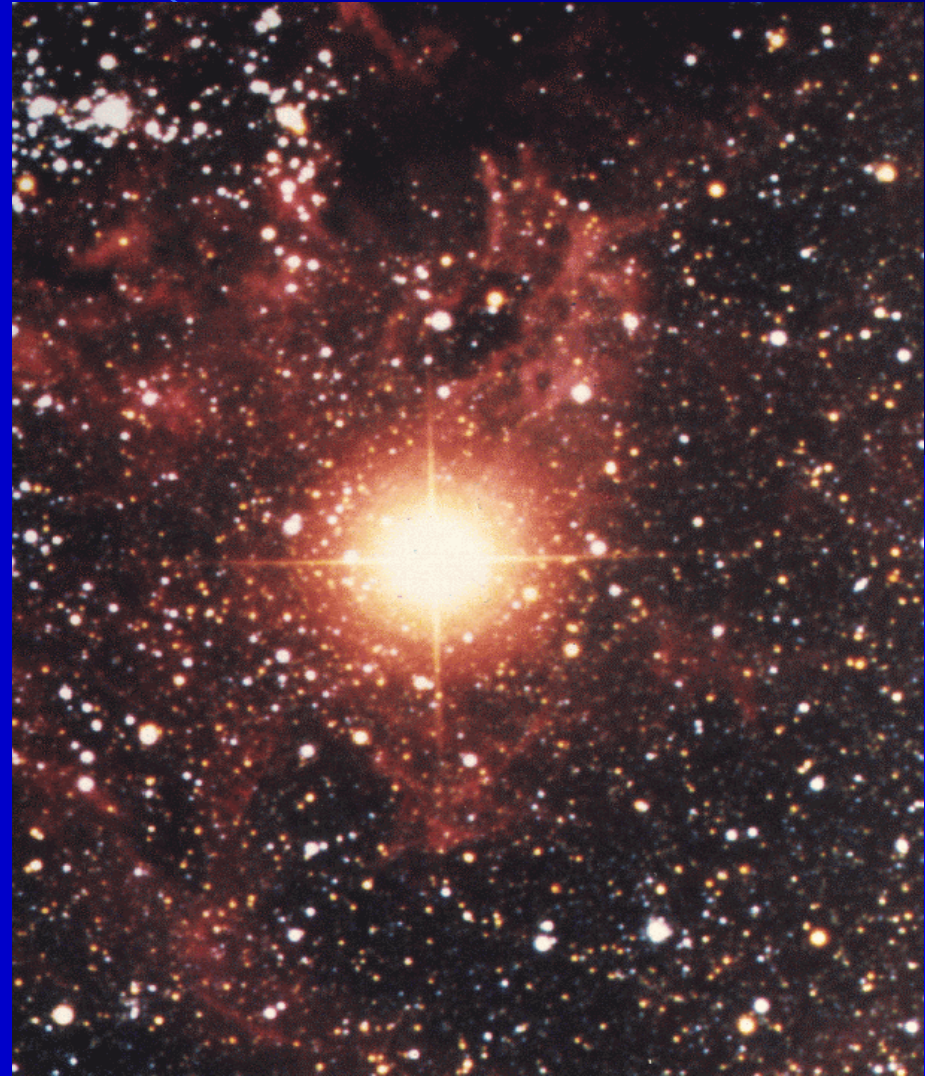


Also: 1 AU at 1 kiloparsec subtends 1 milliarcsec  
1 AU at 1 megaparsec subtends 1 microarcsec

1 parsec is  $3 \times 10^{13}$  km;  
if we can measure the angle, we can get the distance

# Supernova 1987A

- a massive star exploded in the LMC
- February 1987
- the LMC is our nearest neighbour galaxy
- in fact, it's a satellite



# SN1987A

- the ring lit up 250 days after the supernova
- radius known
- angle known
- $\Rightarrow$  distance

Supernova 1987A Rings

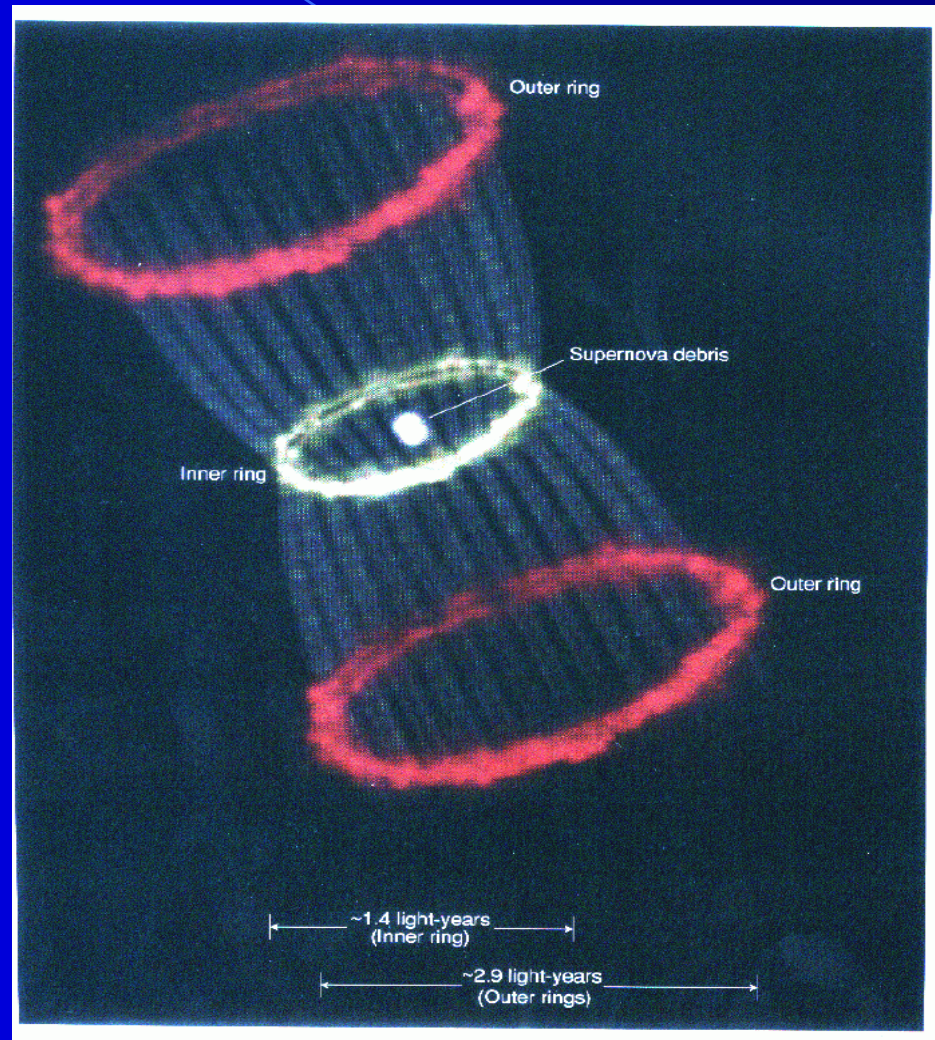


Hubble Space Telescope  
Wide Field Planetary Camera 2



# SN1987A

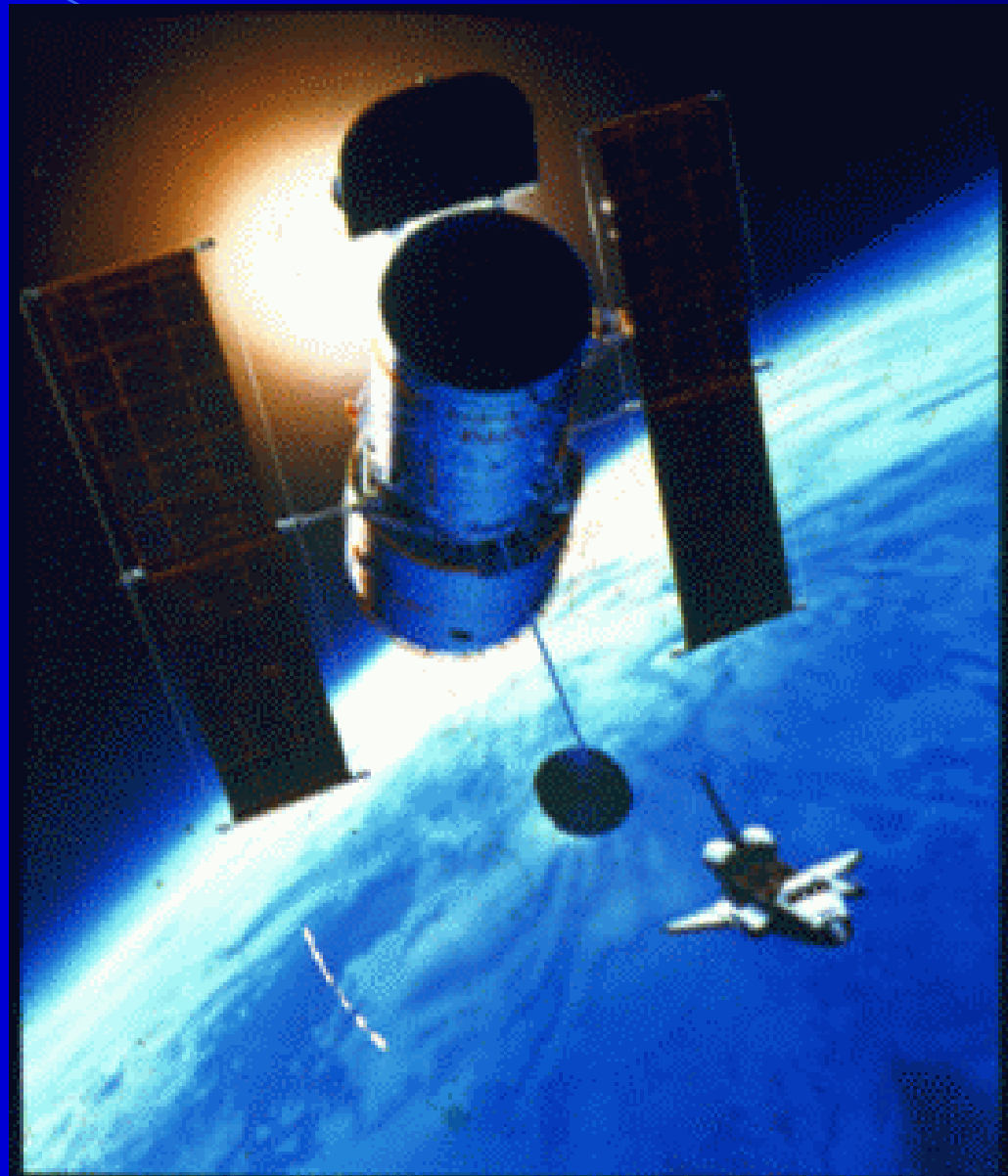
## Schematic of SN1987A



# The Hubble Constant Key Project

Goal:

Measure how fast the  
Universe is  
expanding  
to 10% accuracy



# The HST Key Project

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Shaun Hughes

Barry Madore

Nancy Silbermann

Shoko Sakai

Randy Phelps

Robert Hill

Abi Saha

Peter Stetson

Brad Gibson

Laura Ferrarese

Holland Ford

Garth Illingworth

Dan Kelson

John Graham

John Hoessel

Lucas Macri

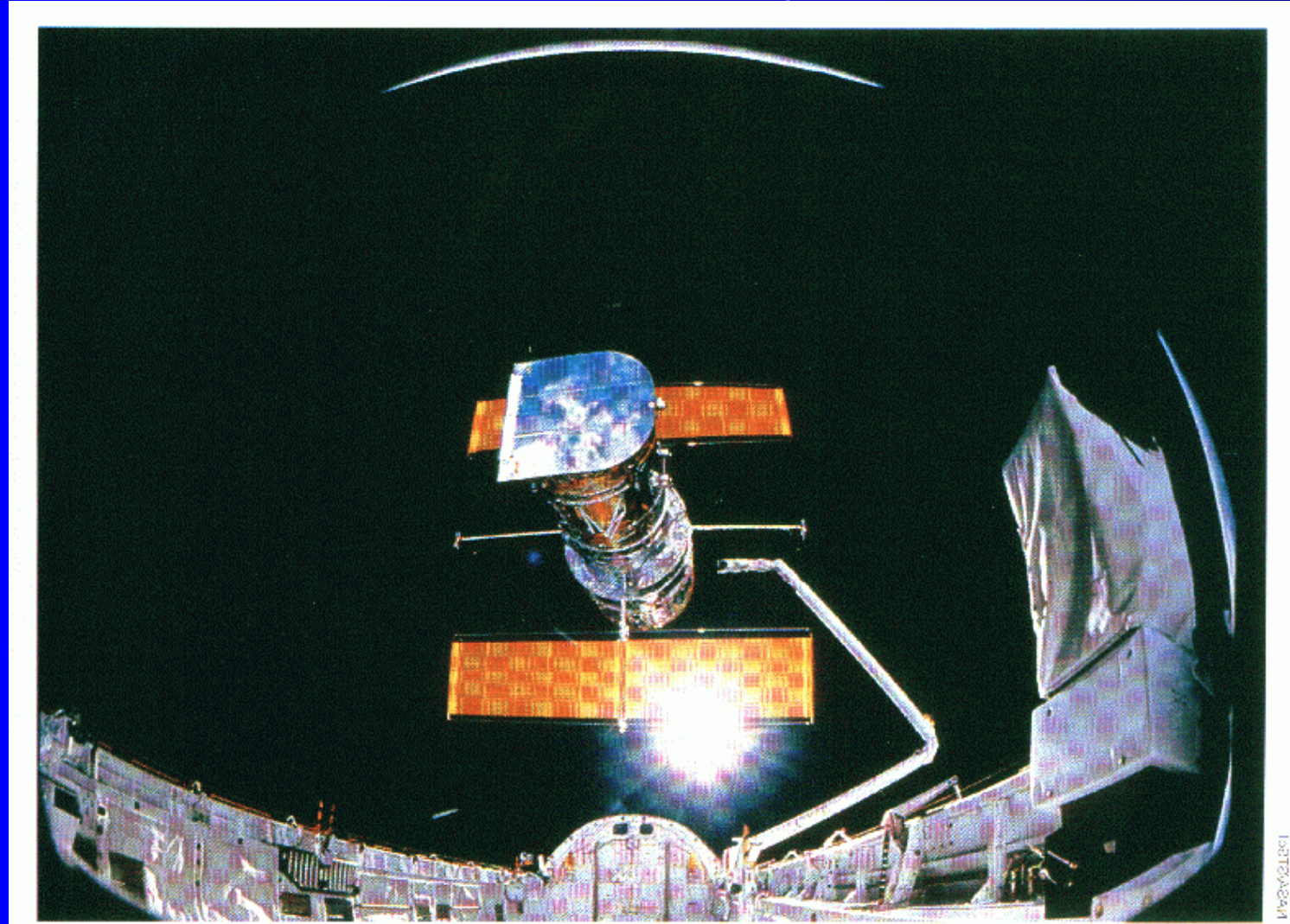
John Huchra

Anne Turner

Paul Harding

Fabio Bresolin

# servicing mission





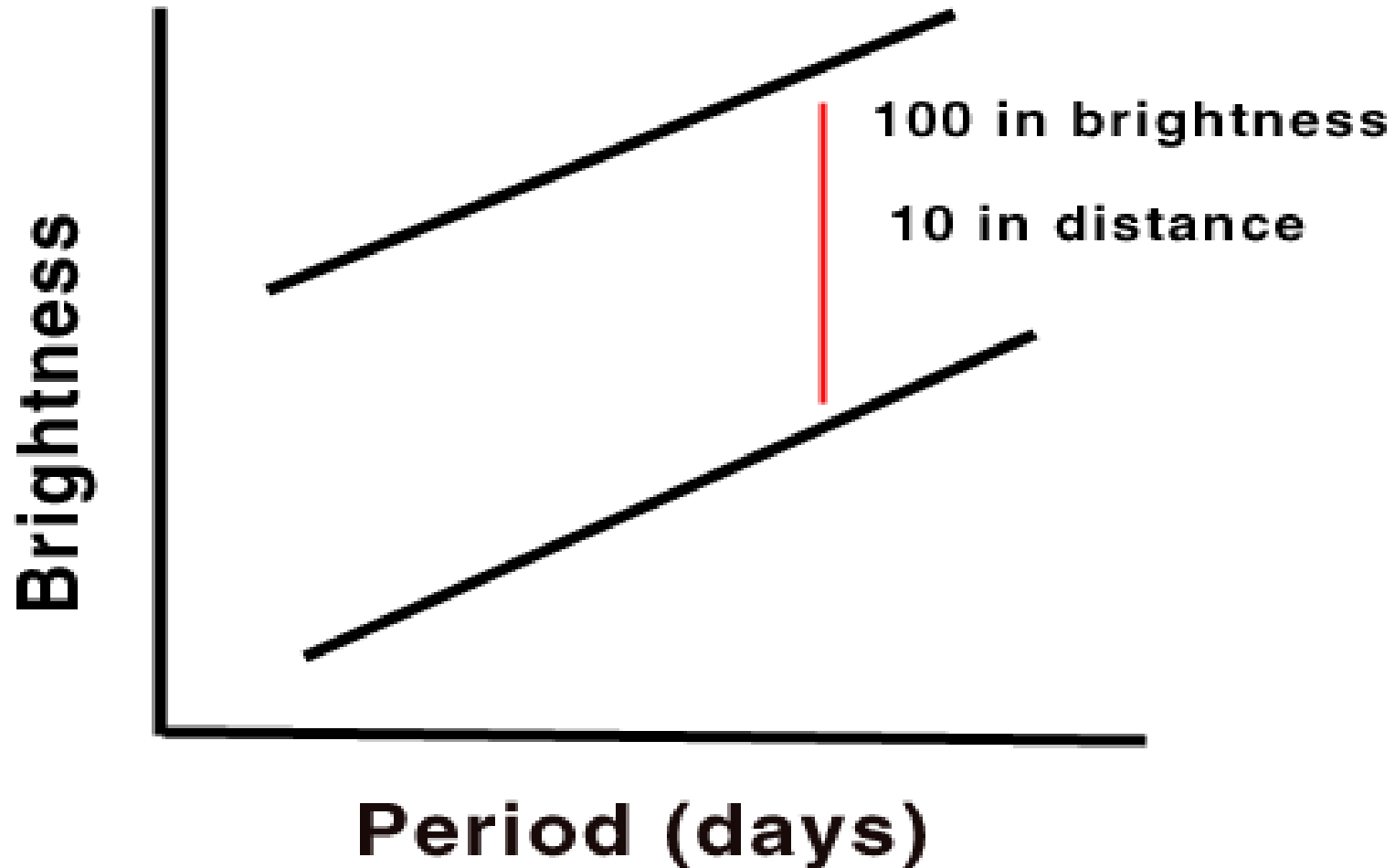
The first servicing mission replaced the WFPC with WFPC2 in December 1993, thus remedying spherical aberration



# The power of the Hubble Space Telescope

- From the ground we can resolve galaxies up to 2 or 3 Mpc away
- HST was designed to have ten times the resolution of ground based telescopes
- The project to find Cepheids up to 20 Mpc away was designated a Key Project for HST

# The Cepheid period luminosity relation



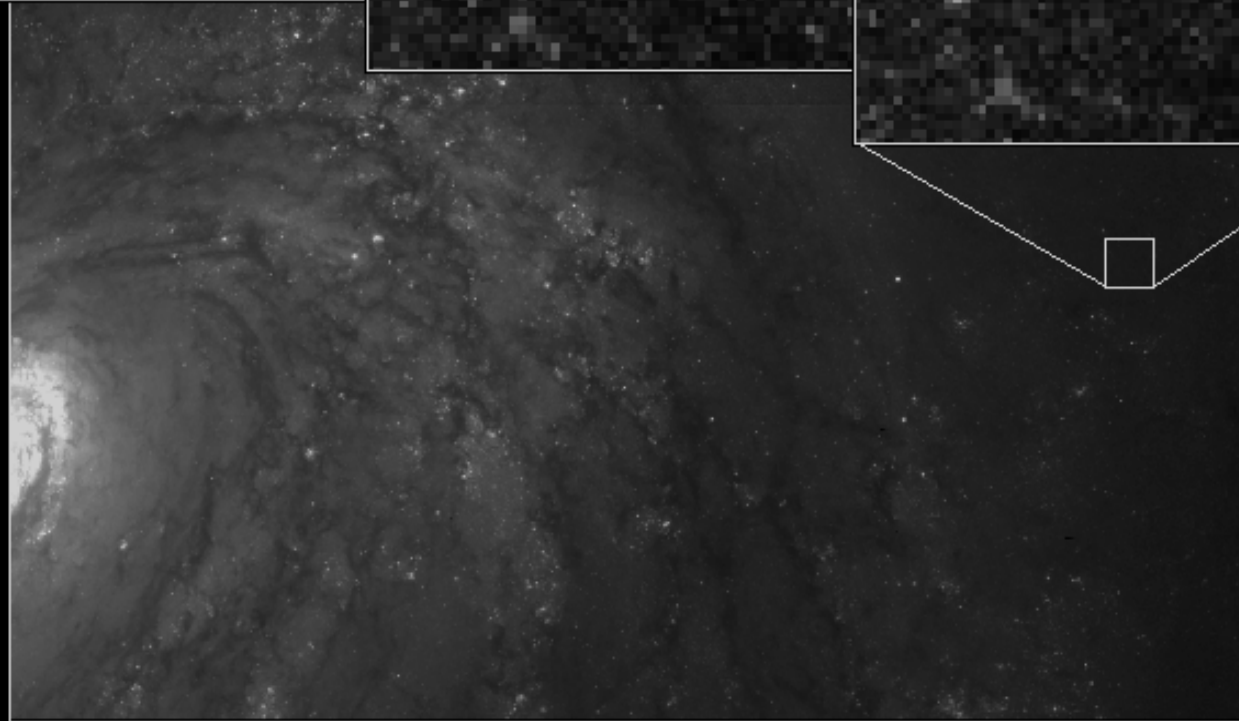
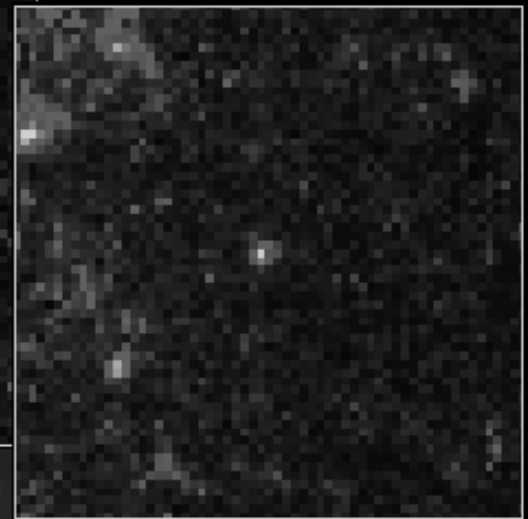
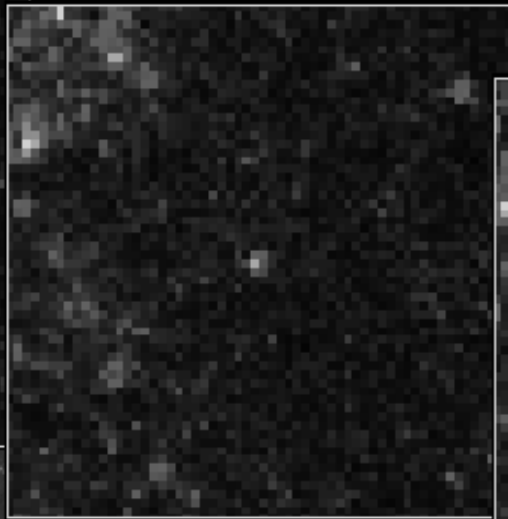
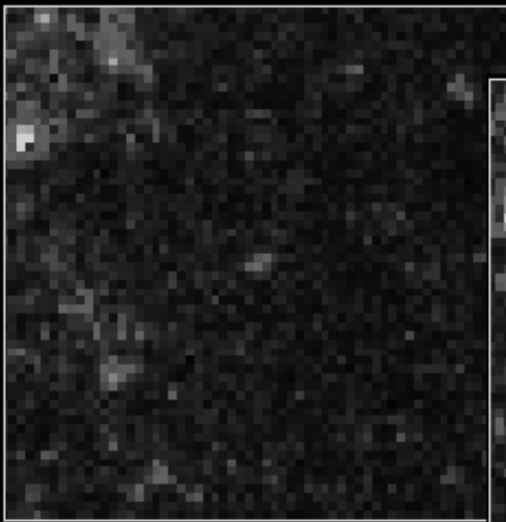
**M 100**

HST-WFPC2

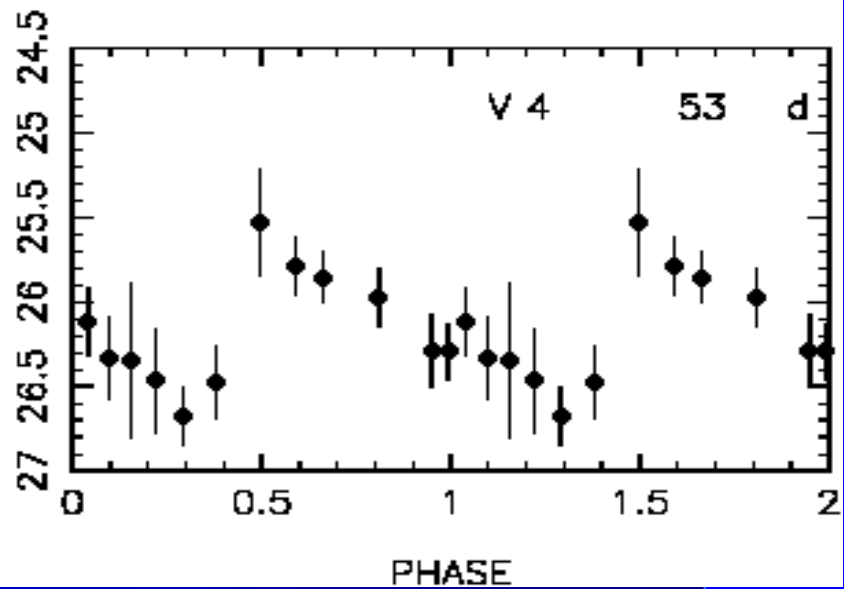
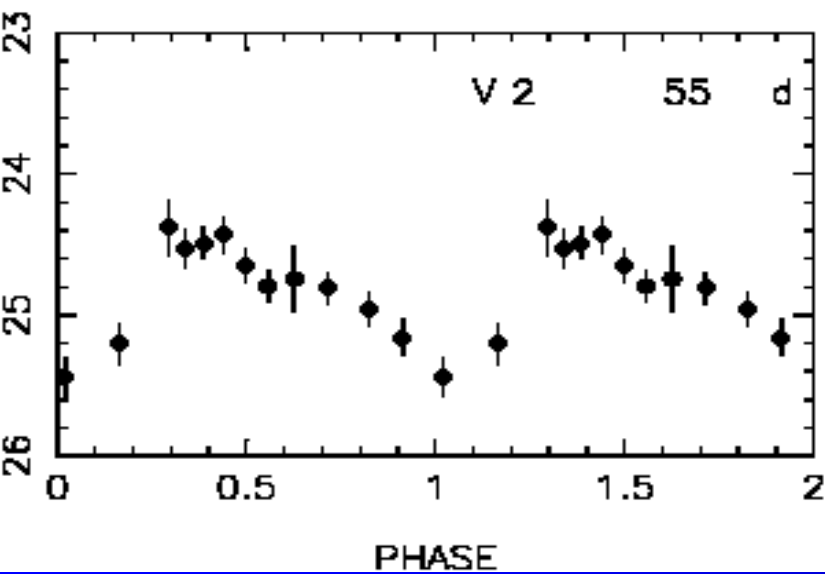
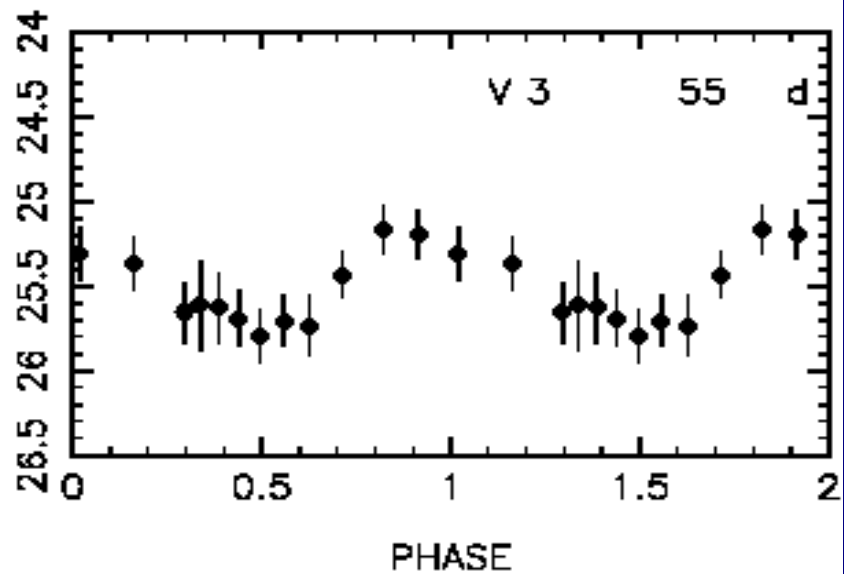
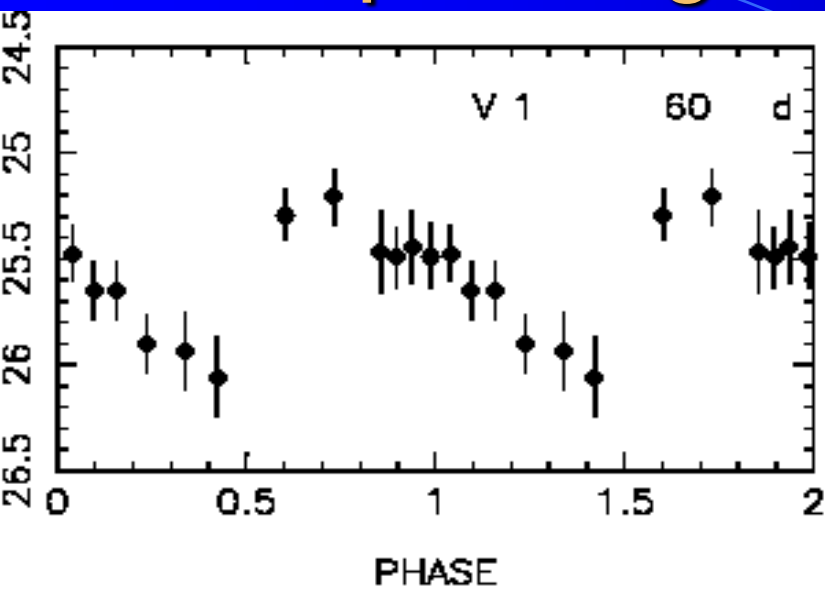


# Cepheid Variable in M100

## HST-WFPC2



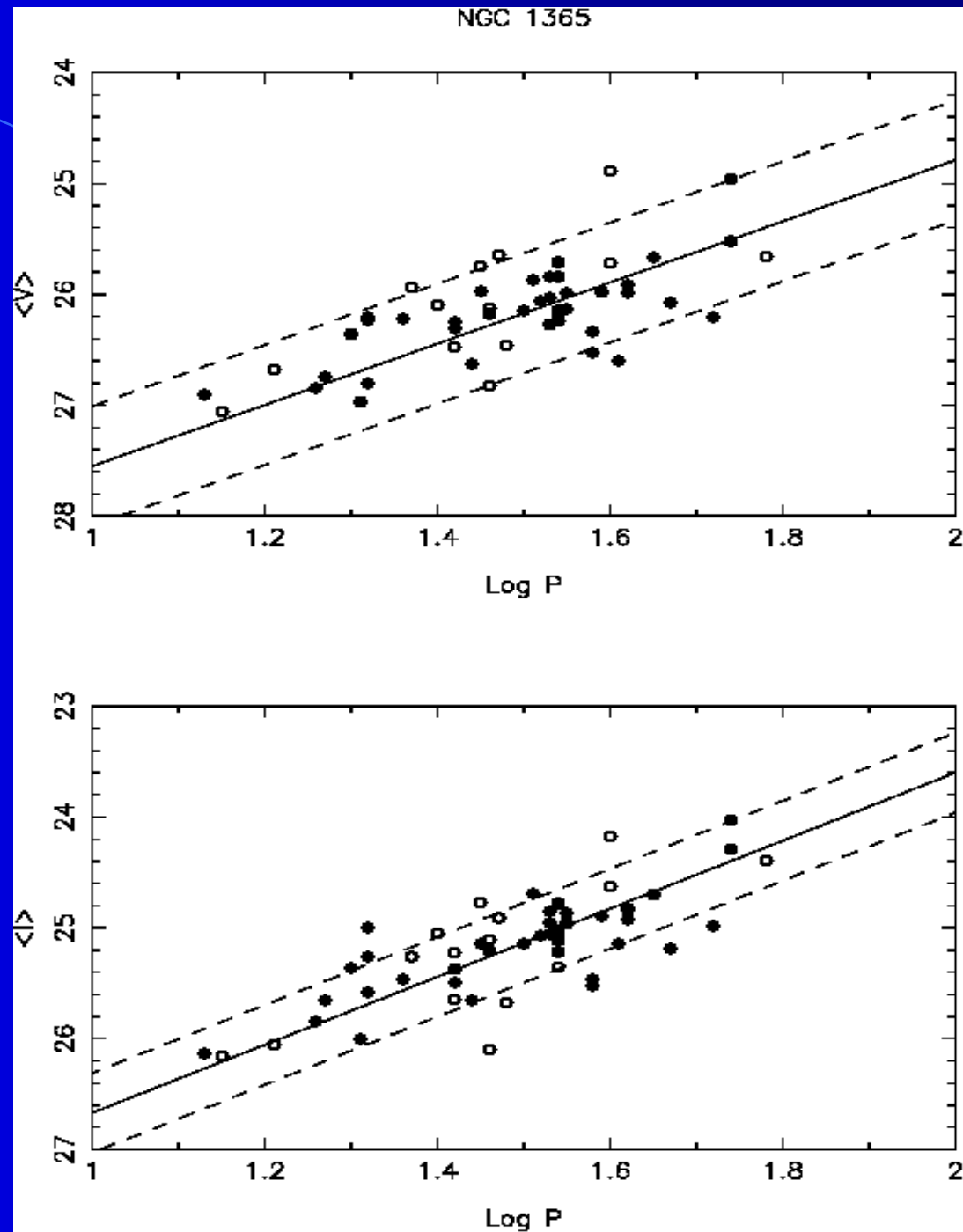
# Cepheid light curves in NGC 1365

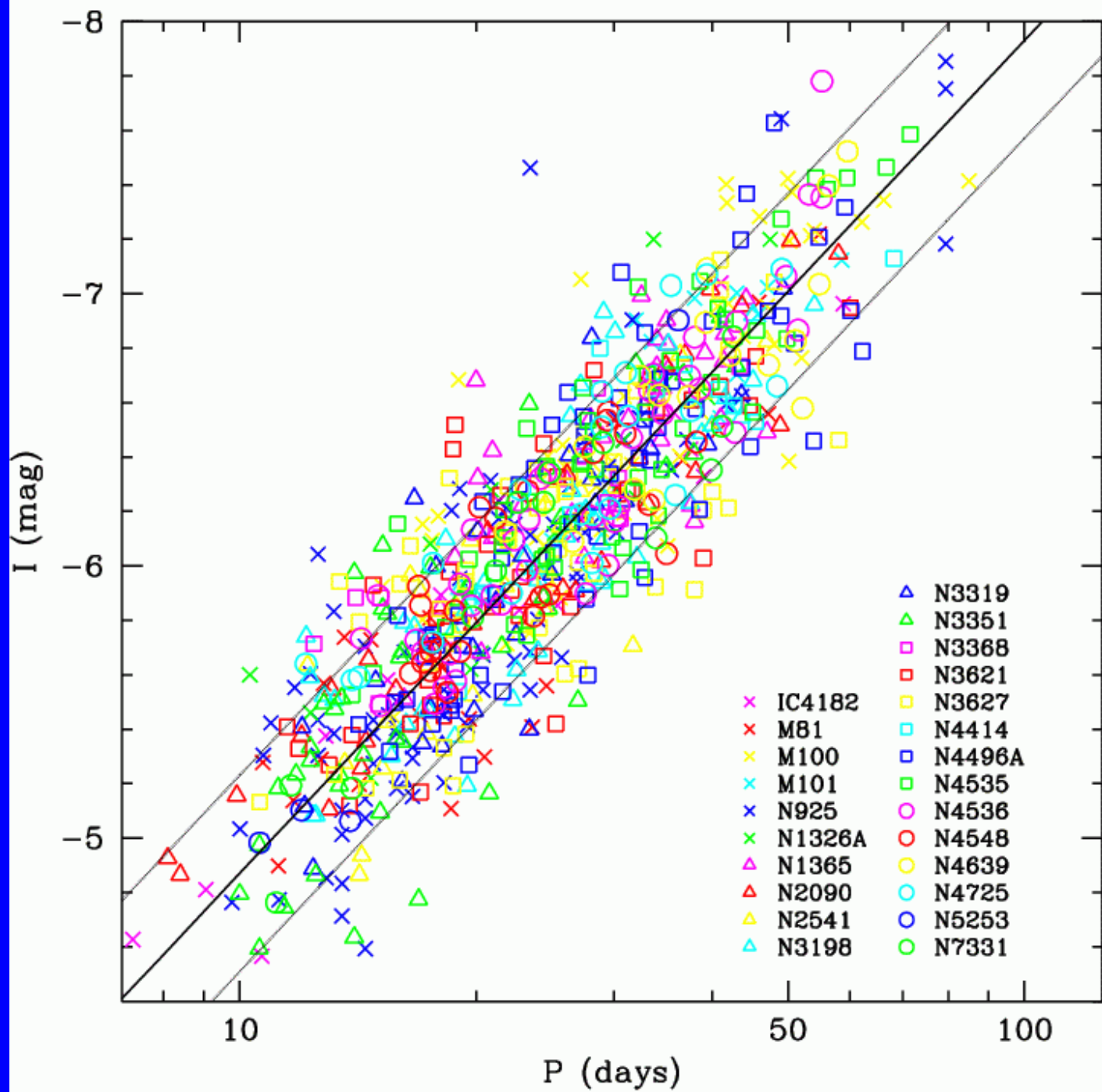


*top: visual*

## The PL relation in NGC 1365

*bottom: infrared*





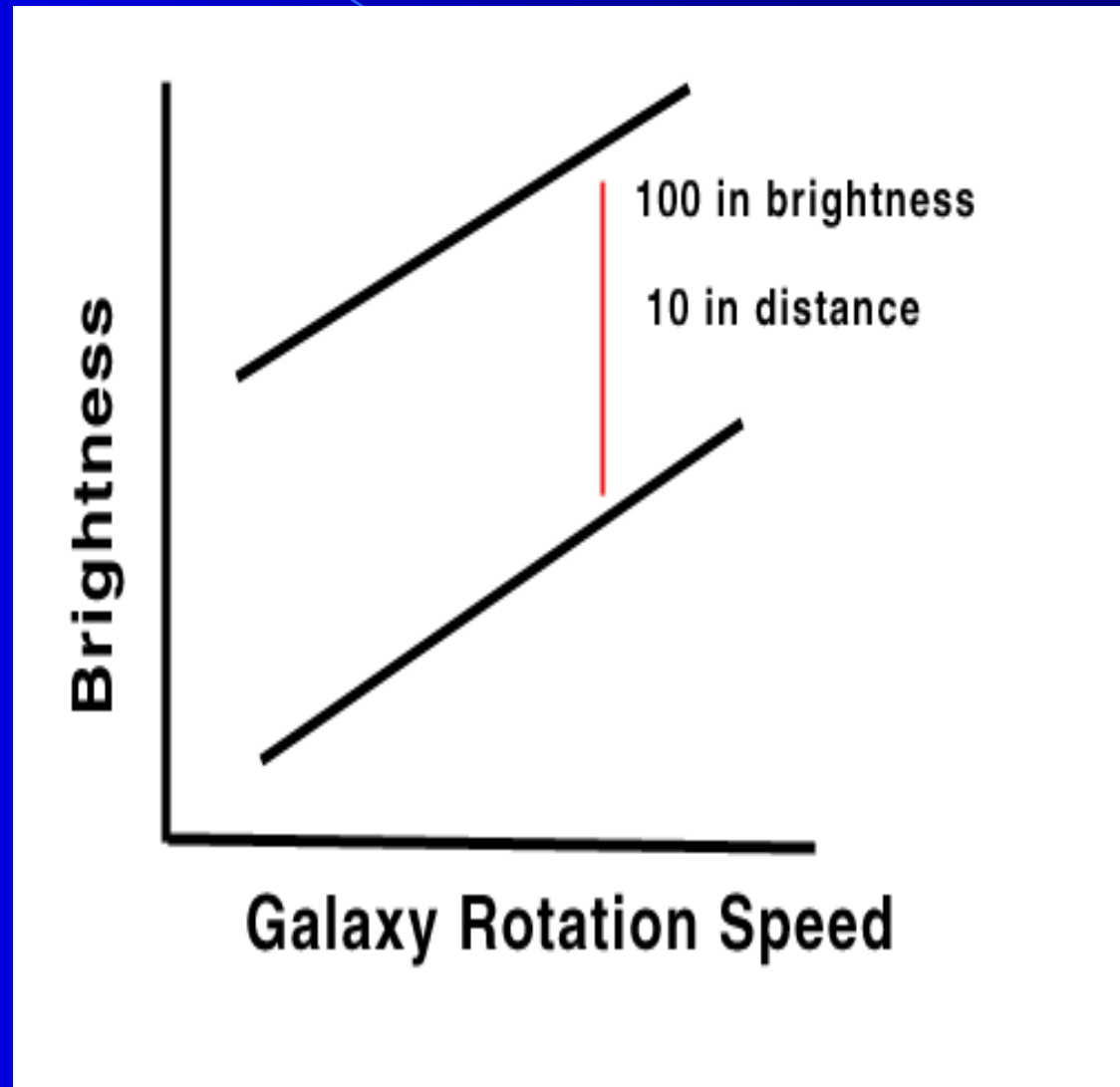


# Beyond the Cepheids

- Beyond 20 Mpc even HST has difficulty resolving Cepheids
- We use four other standard candles to measure distances a further factor of ten
- The Tully Fisher relation for spiral galaxies
- Supernovae of type Ia
- Surface brightness fluctuations = resolvability
- The fundamental plane for elliptical galaxies

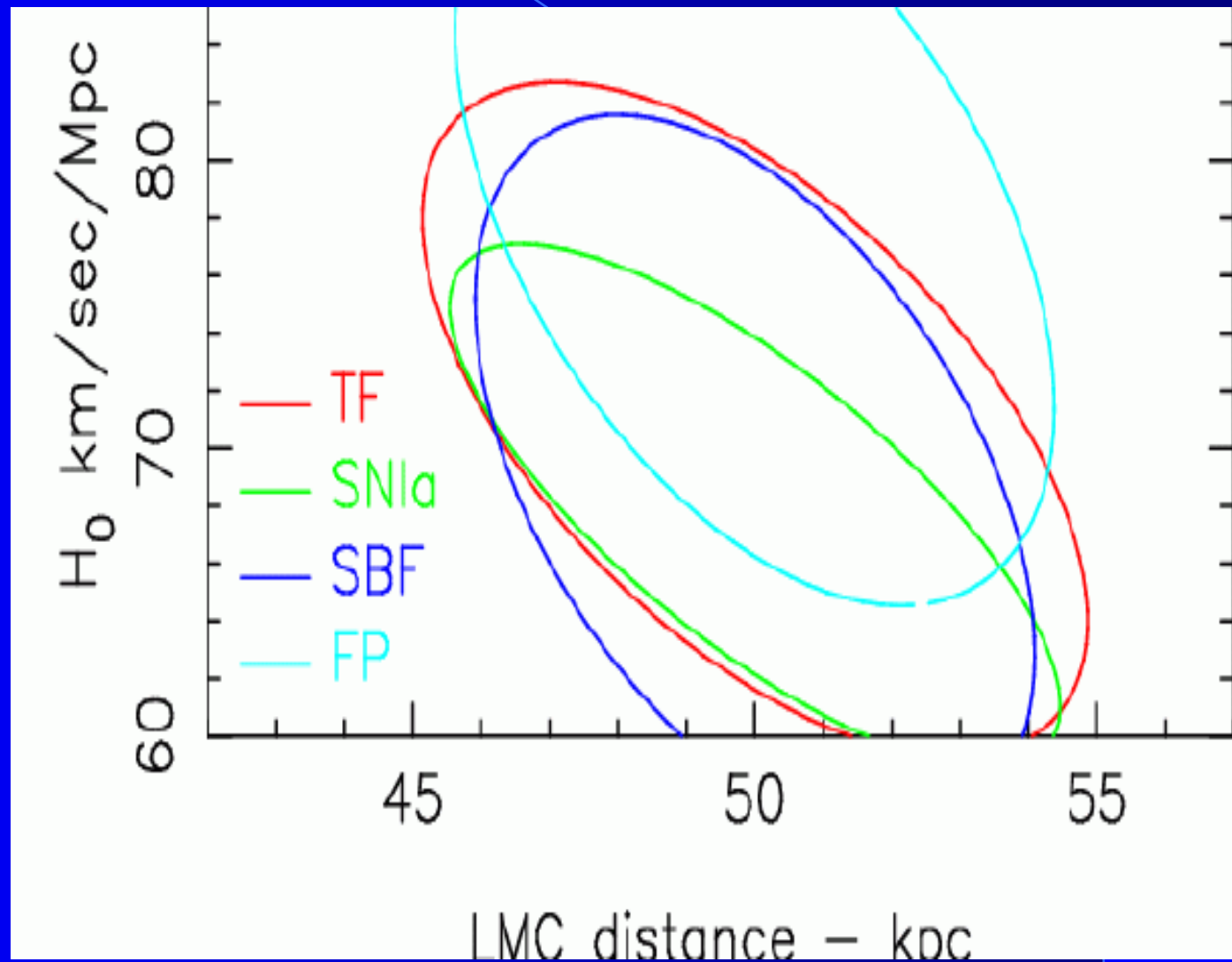
# Tully Fisher relation

- big galaxies rotate faster
- a galaxy with a given rotation speed is a standard candle



# The Hubble Constant

- All 4 standard candles agree
- $H_0$  lies in the range 65 to 77 km/s/Mpc



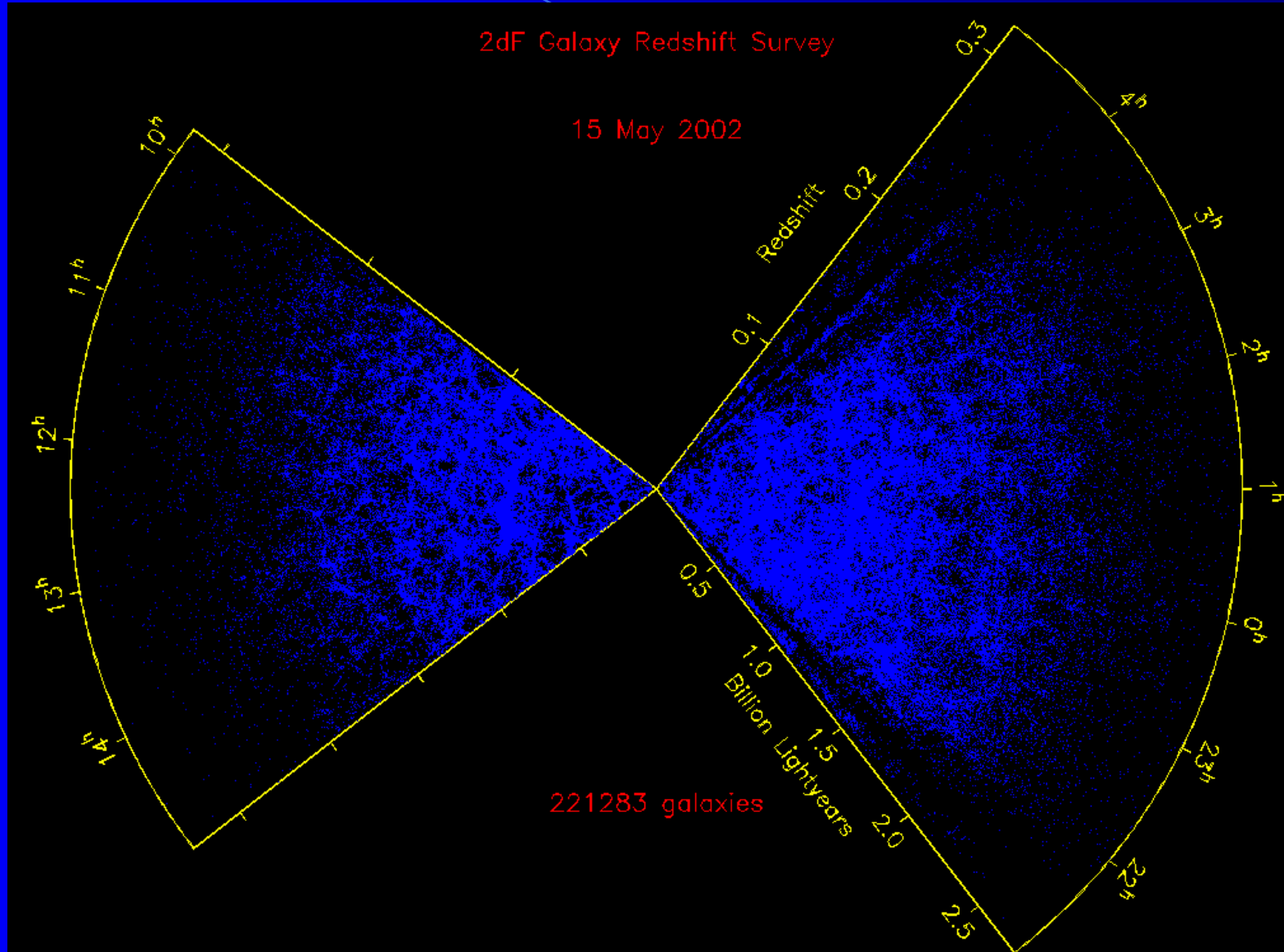
# How standard are standard candles ?

- most stars have a chemical composition like that of the sun
- but there are some variations....
- Cepheids with different chemistry pulsate differently
- accounting for this changes our distances a few percent

# What is the density of the Universe ?

2dF  
map of  
the 3d  
density  
of the  
local  
few  
billion  
light  
years

Colless  
and  
team



# Rediscovery of $\Lambda$



See  
[www.mso.anu.edu.au/~brian](http://www.mso.anu.edu.au/~brian)

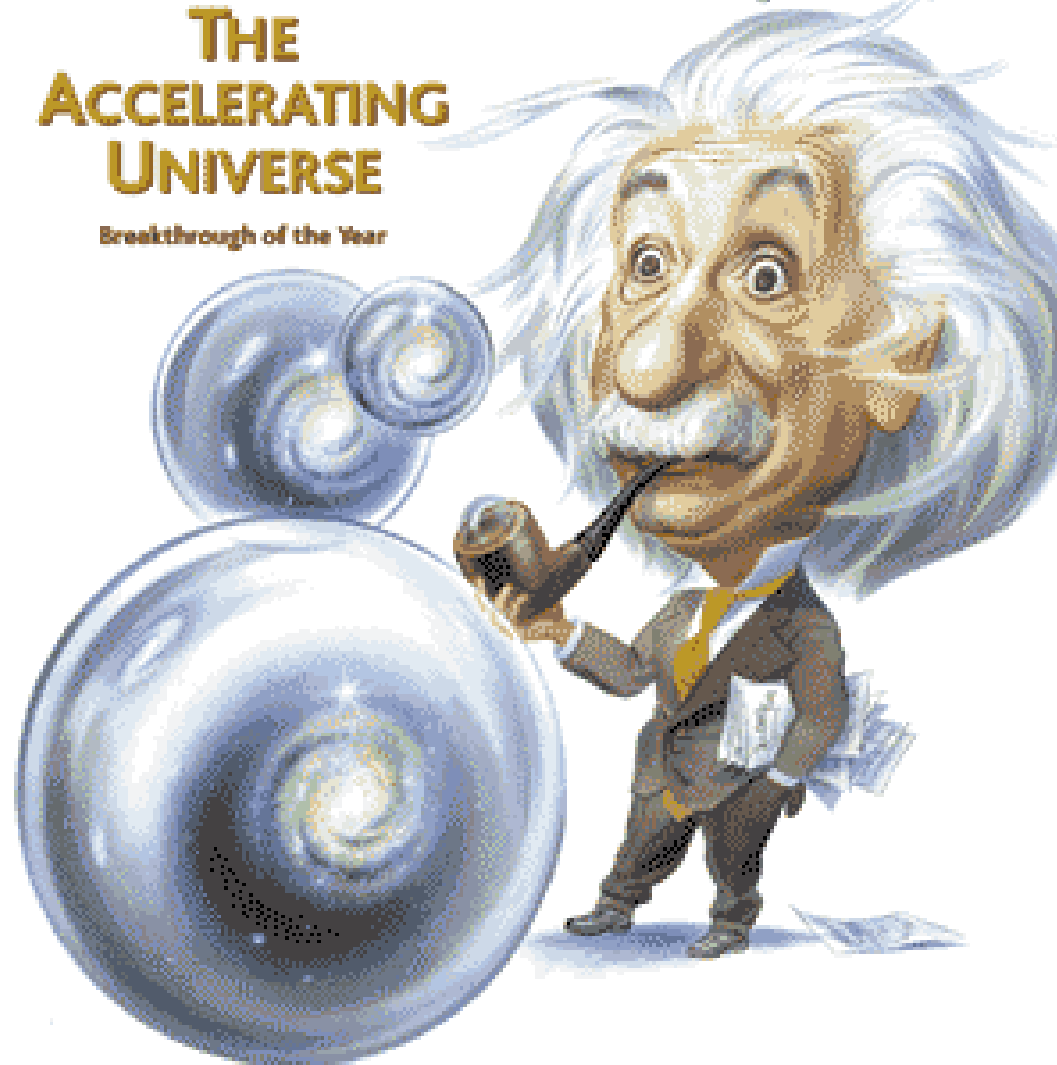
# Science

18 December 1998

Vol. 282 No. 5397  
Pages 2141-2336 \$7

## THE ACCELERATING UNIVERSE

Breakthrough of the Year

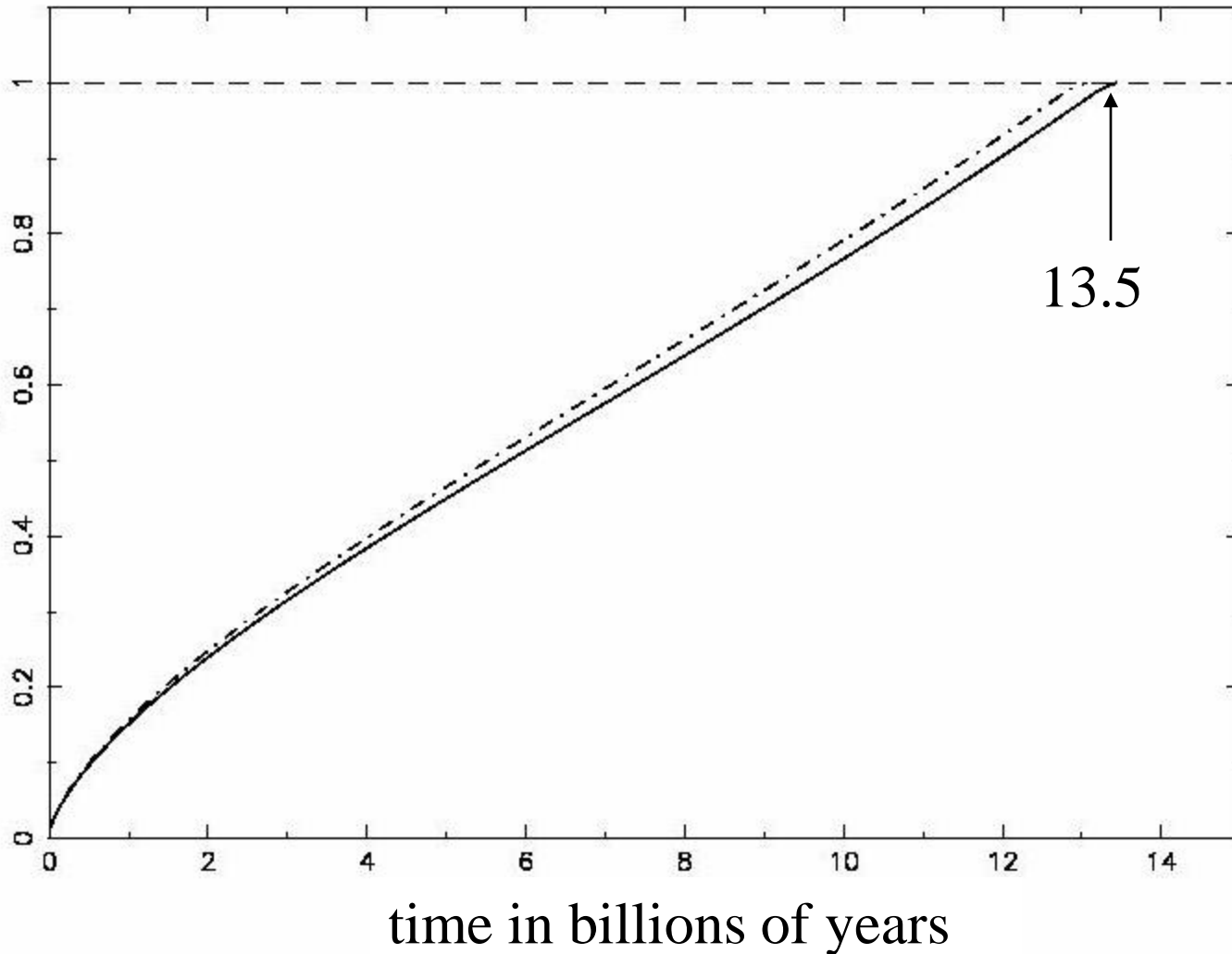


AMERICAN ASSOCIATION FOR THE ADVANCEMENT OF SCIENCE

# Measuring the age

Scale  
factor

a



← Size now

# Summary

- we start with the distance of the LMC
- a Cepheid of period  $P$  is a standard candle of luminosity  $L$
- HST maps the Cepheids out to 20 Mpc
- Four other standard candles map the expansion out to 200 Mpc
- $H_0 = 72 \pm 7$  km/sec/Mpc
- Universe is  $13.5 \pm 1.5$  Gyrs old



# Oldest stars

- globular star clusters
- parallaxes with SIM or GAIA
- measure ages to half a billion years



# Where to get more information

- ‘Measuring the Universe’ by Stephen Webb
- [www.stsci.edu](http://www.stsci.edu) (Hubble Space Telescope)  
<http://oposite.stsci.edu/pubinfo/1999.html>