

International Centre for Radio Astronomy Research

The importance of robust identification of simulated galaxies

Mock Perth: Challenges for Simulations in the Era of SKA and Large IFU Surveys

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Curtin University

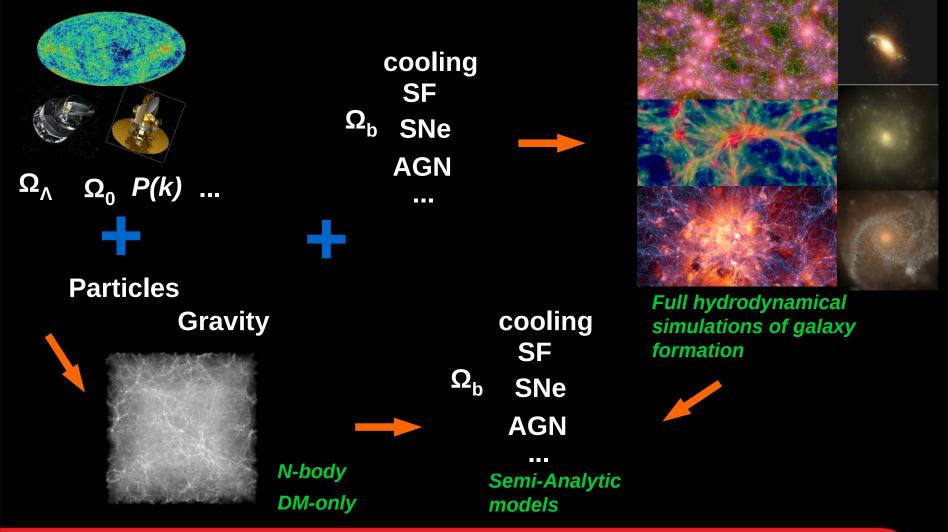


Government of Western Australia artment of the Premier and Cabinet



Small background

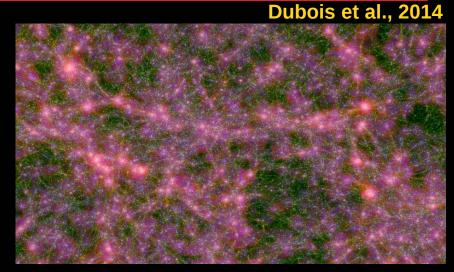
Numerical simulations of galaxy formation

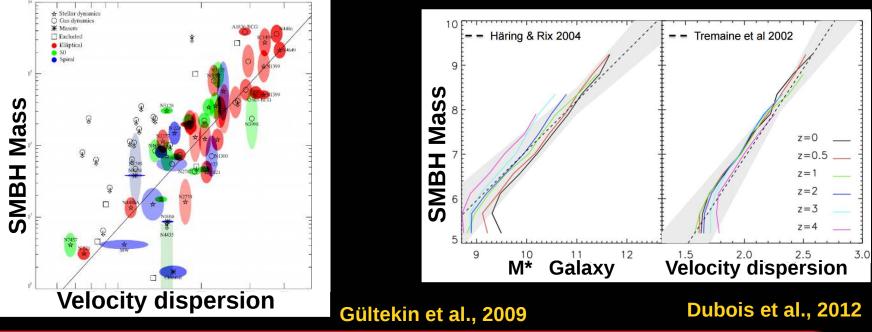




Horizon-AGN simulation

- State-of-the-art cosmological simulation of galaxy formation
- Two AGN feedback modes: *radio* and *quasar*
- Coupling between SMBH and host galaxy
- AGN feedback tuned to reproduce observed scaling relations



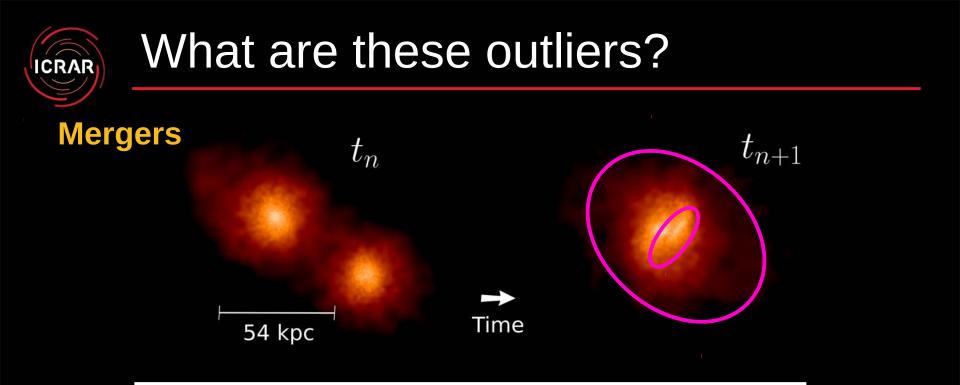


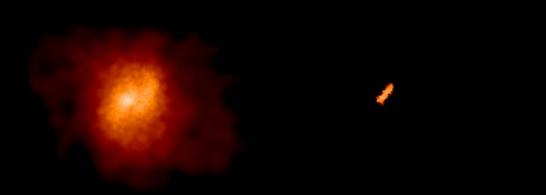


AGN-Galaxy co-evolution

Exploring Horizon-AGN

- Reproduce already published Mbh-Mgal* relation
 - 1e+10 1010 1e+09 109 **SMBH Mass SMBH Mass** 1e+08 108 1e+07 107 1e+06 106 1e+09 1e+11 1e+12 1e+10 1010 1012 109 1011 **Galaxy Stellar Mass Galaxy Stellar Mass** Volonteri et al., 2016
- More scatter than expected





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Galaxies in catalog (Halo-Maker)



What do we define as a galaxy?

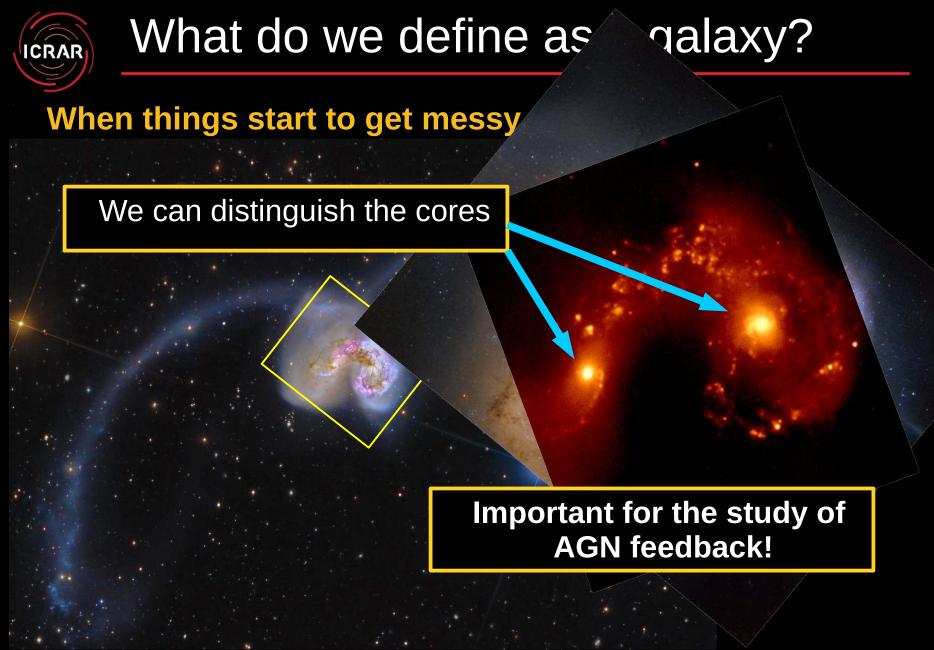
"Well behaved" Mergers





Arp 273

Arp 271



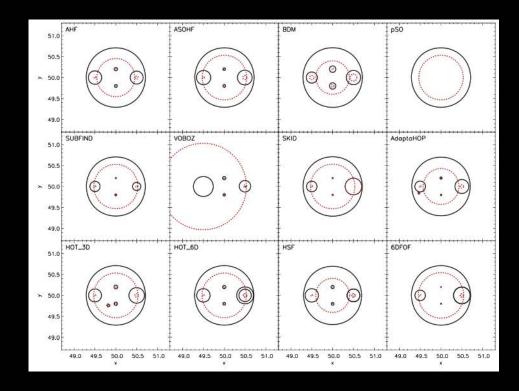
Antennae Galaxies - NGC 4038, NGC 4039



Galaxy finding

"Crazy" Comparison projects

- Haloes gone MAD (Knebe et al. 2011)
- Subhaloes gone Notts (Onions et al. 2013; Pujol et al. 2014; Hoffmann et al. 2014)
- Galaxies going MAD (Knebe et al. 2013)
- Streams going Notts (Elahi et al. 2013)



Knebe et al., 2011

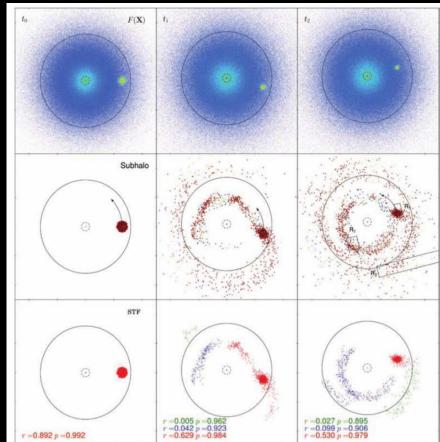


How do we solve this?



VELOCIraptor a.k.a. STF (Elahi et al., 2011)

- FOF algorithm in 6D to find structures in simulations
 - Halos
 - Galaxies
 - Streams
 - Shells



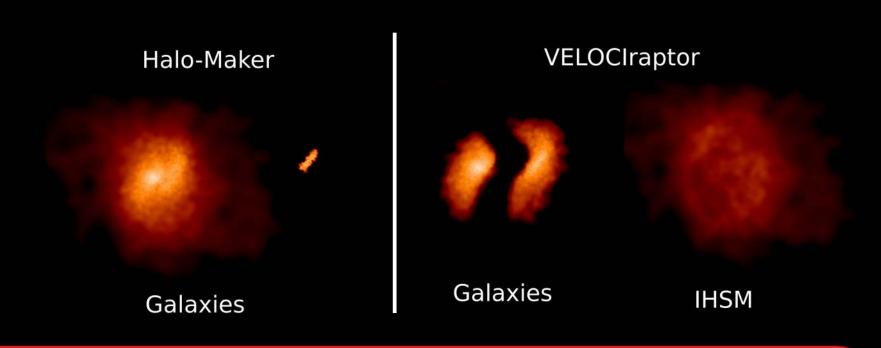
Elahi et al., 2011



Identification of galaxies

Modifications to VELOCIraptor

- Adaptive linking lengths based on the distribution of the system (independent of resolution and environment)
- We can keep track of the Intra-Halo Stellar Mass (IHSM) at no additional computational cost

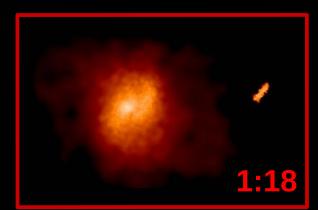


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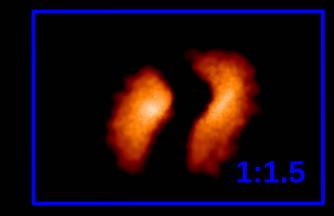


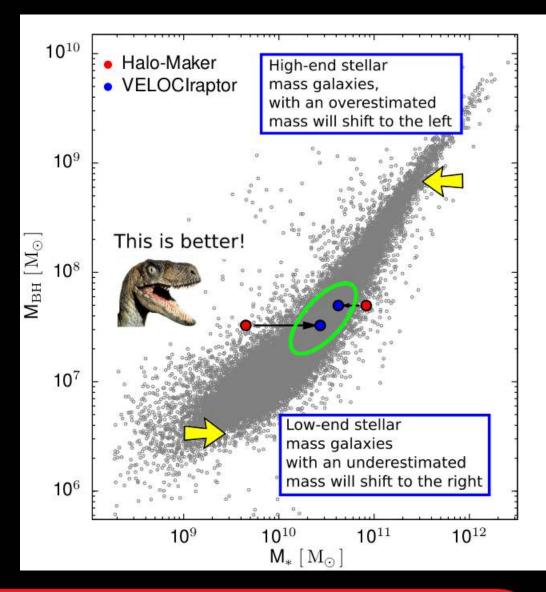
Why is this important?

Halo-Maker



VELOCIraptor



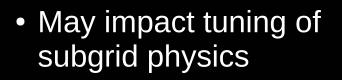


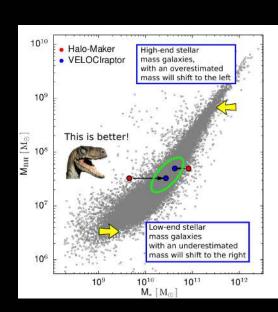


Why is this important?

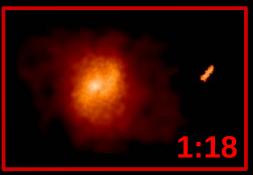
Robust identification of galaxies

- Affects the scatter of scaling relations
- Mass functions
- Estimation of the number of major and minor mergers

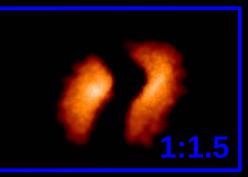




Halo-Maker



VELOCIraptor



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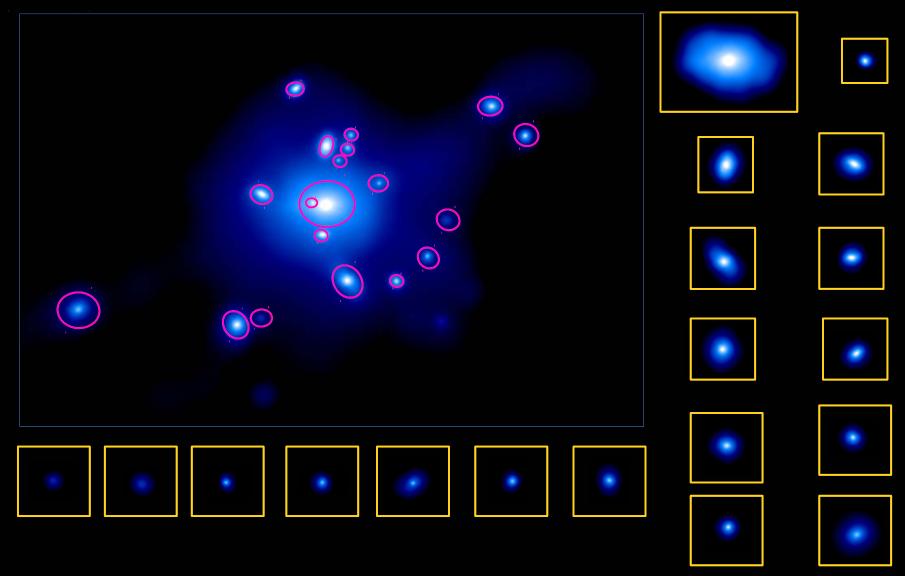


Groups of Galaxies





Groups of Galaxies





Groups of Galaxies

Intra-Halo Stellar Mass (IHSM)

ICRAR

Groups of Galaxies

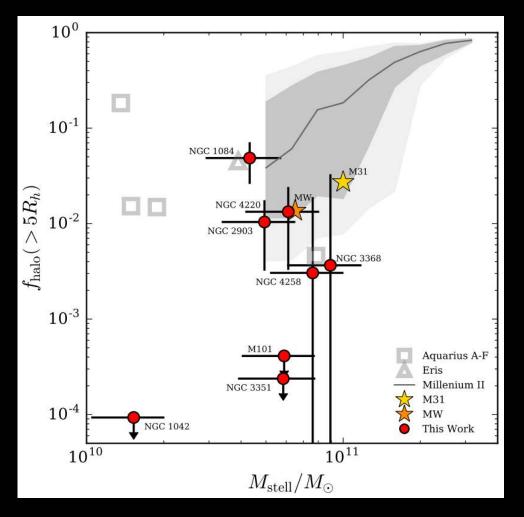
Galaxies

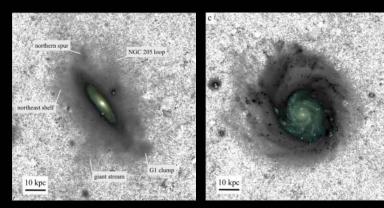
Intra-Halo Stellar Mass (IHSM)

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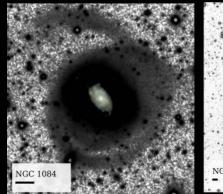


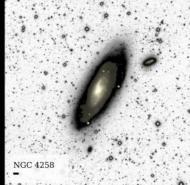
Diffuse stellar component





van Dokkum et al., 2014





Merritt et al., 2016

Merritt et al., 2016

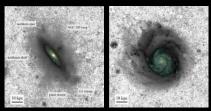


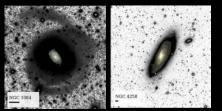
Intra-Halo Stellar Mass

- Stellar halos (e.g. Merritt et al. 2016)
- Intra-Cluster Light (e.g. Morishita et al. 2016) van Dokkum et al., 2014
- Predictions for future observational projects and surveys (e.g WAVES, JWST, WFIRST)

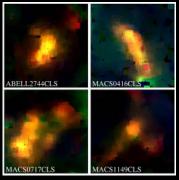
Galaxy identification

- Use improved VELOCIraptor to build new galaxy catalogs
 - Horizon-AGN
 - EAGLE
 - Illustris





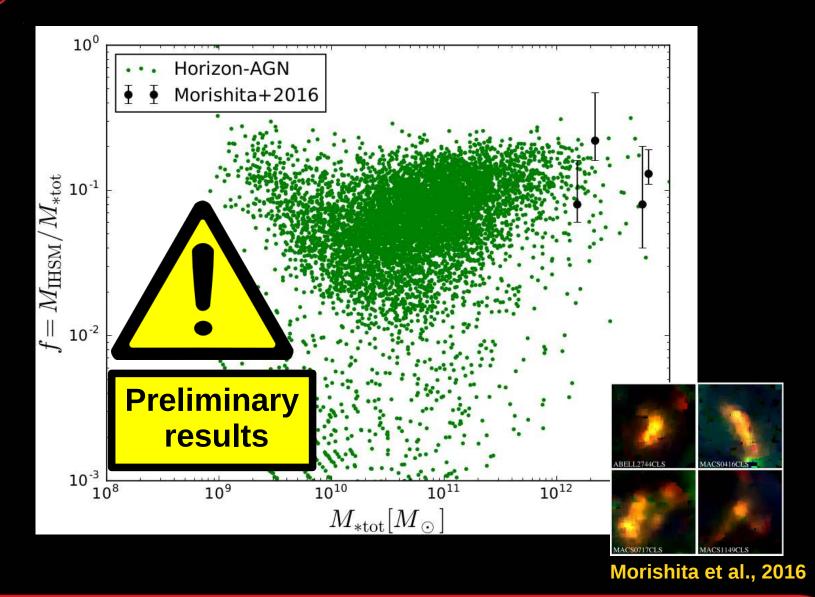
Merritt et al., 2016



Morishita et al., 2016

Diffuse stellar component

ICRAR





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References

- Dubois Y., et al., 2012, MNRAS, 420, 2662
- Dubois Y., et al., 2014, MNRAS, 444, 1453
- Elahi P. J., et al., 2011, MNRAS, 418, 320
- Gultekin K. et al., 2009, ApJ, 698, 198
- Merritt, A., et al., 2016, ApJ, 830, 62
- Morishita, T., et al., 2016, arXiv:1610.08503