

A movie poster for 'The Usual Suspects' featuring five men standing in front of a height chart. The height chart on the left shows measurements from 3'0" to 6'0" in increments of 6 inches. The men are dressed in various styles, including leather jackets, suits, and casual wear. A semi-transparent black banner with white text is overlaid across the middle of the image.

Uncovering key contributors to cosmic reionization

**Jeff Cooke
Emma Ryan-Weber
Thibault Garel
Gonzalo Díaz**

SWIN
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SWINBURNE
UNIVERSITY OF
TECHNOLOGY

NOT
The
Usual Suspects



the crime scene

the “crime”: *the universe was reionized*

- galaxies are largely responsible

$$n_{\text{ion}} = f_{\text{esc}} \xi_{\text{ion}} \rho_{\text{UV}}$$

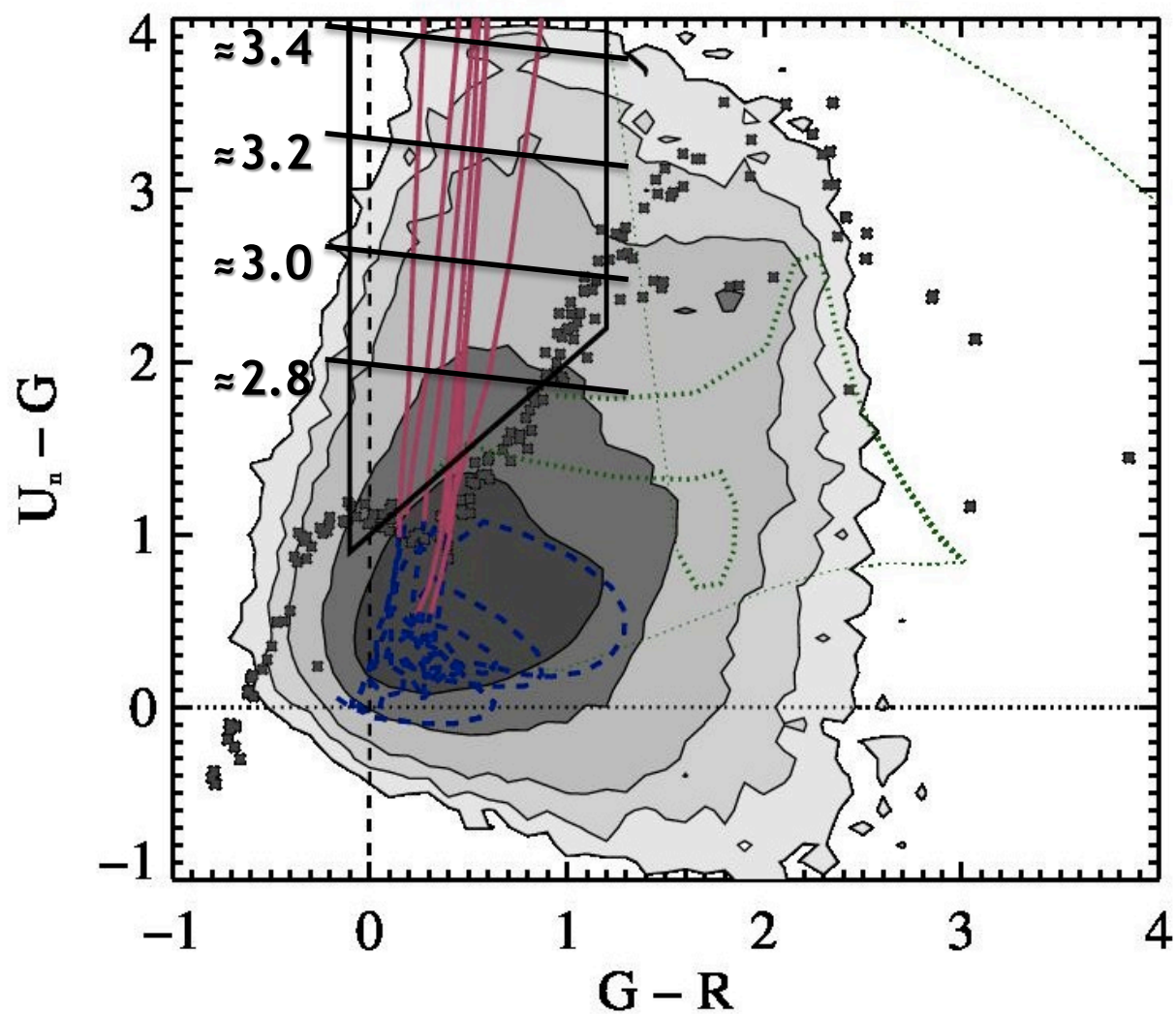
- need a full high redshift census
- need to determine the escape fraction, f_{esc}
- get at ξ_{ion} from the measured continuum slope B

- measuring f_{esc} of galaxies is challenging
 - observationally $f_{\text{esc}} = (f_{1500\text{\AA}}/f_{900\text{\AA}})$
 - expected observed $f_{900\text{\AA}} < 1 \times 10^{-30} \text{ ergs s}^{-1} \text{ cm}^{-2} \text{ Hz}^{-1}$ ($m > \approx 26$)
 - optical detectors are more efficient to the red, $> \sim 3500\text{\AA}$
 - continuum absorption by the Ly α forest increases with redshift
- $z \sim 3 - 4$ is the “sweet spot” for $< 912\text{\AA}$ measurements



rounding up the usual suspects

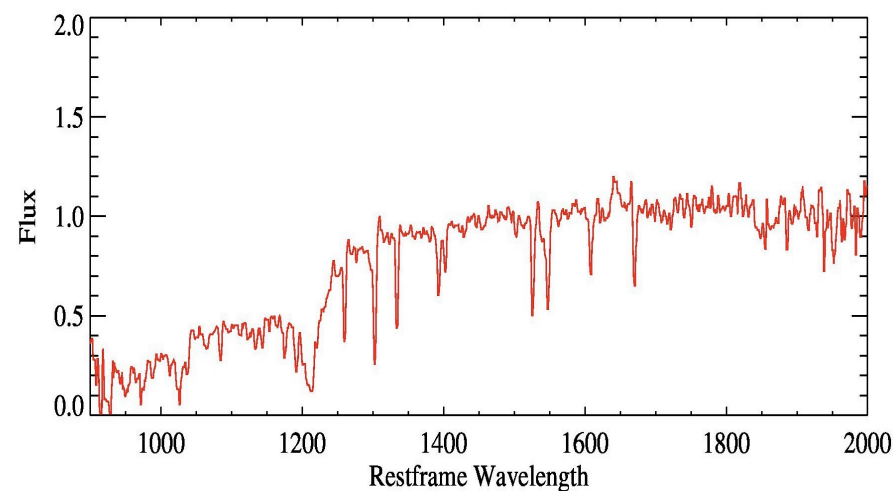
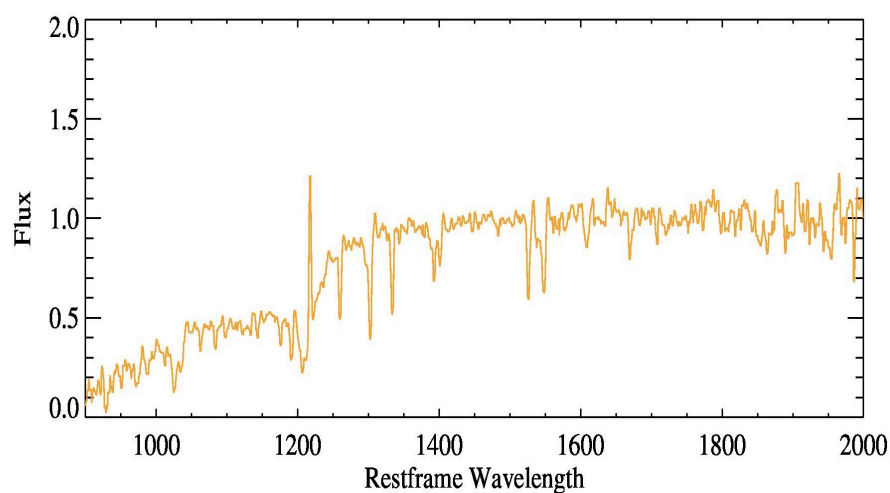
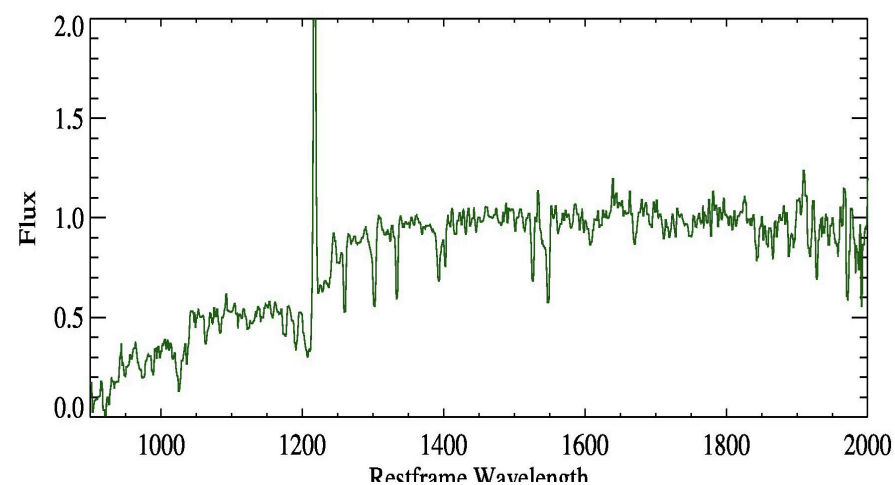
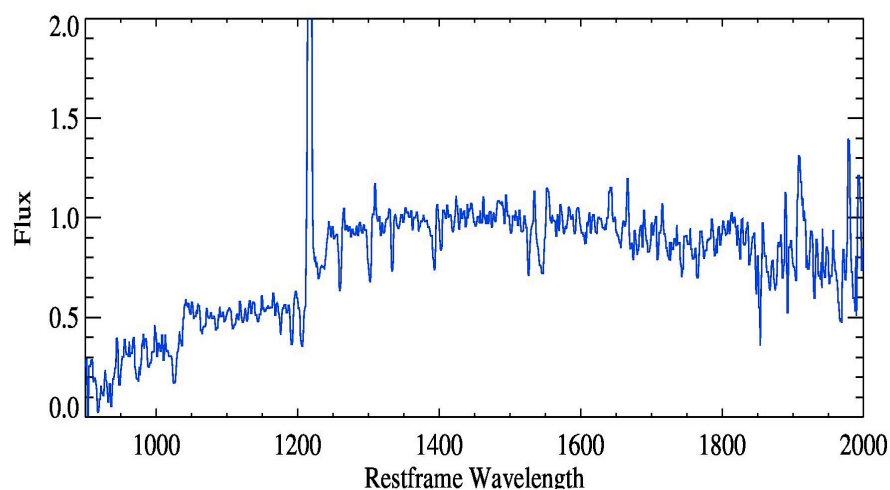
U_n GR-selected $z \sim 3$ LBGs (e.g., Steidel et al. 1996, 2003)





rounding up the usual suspects

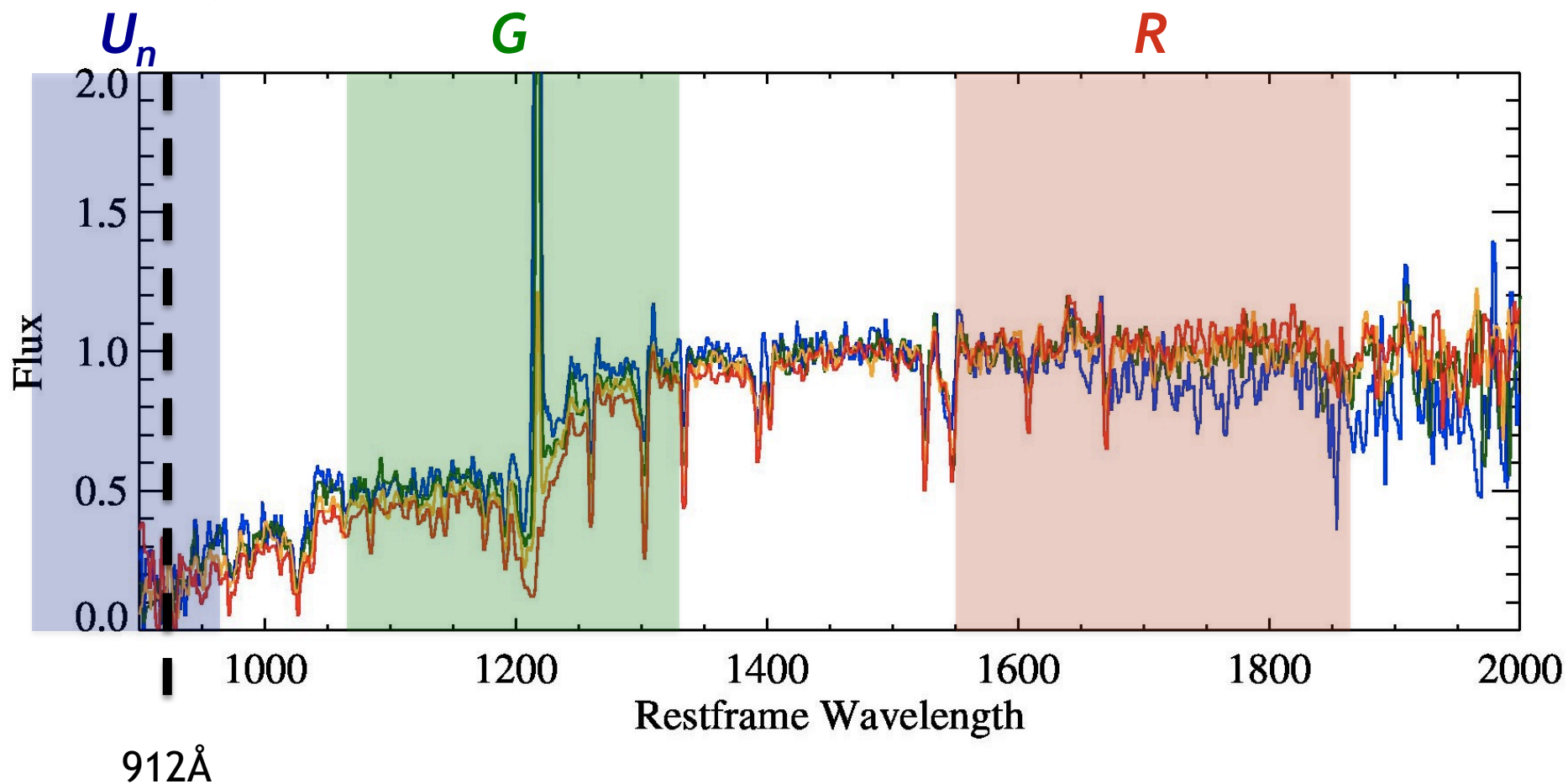
U_n GR-selected $z \sim 3$ LBGs composite spectra (Shapley et al. 2003)





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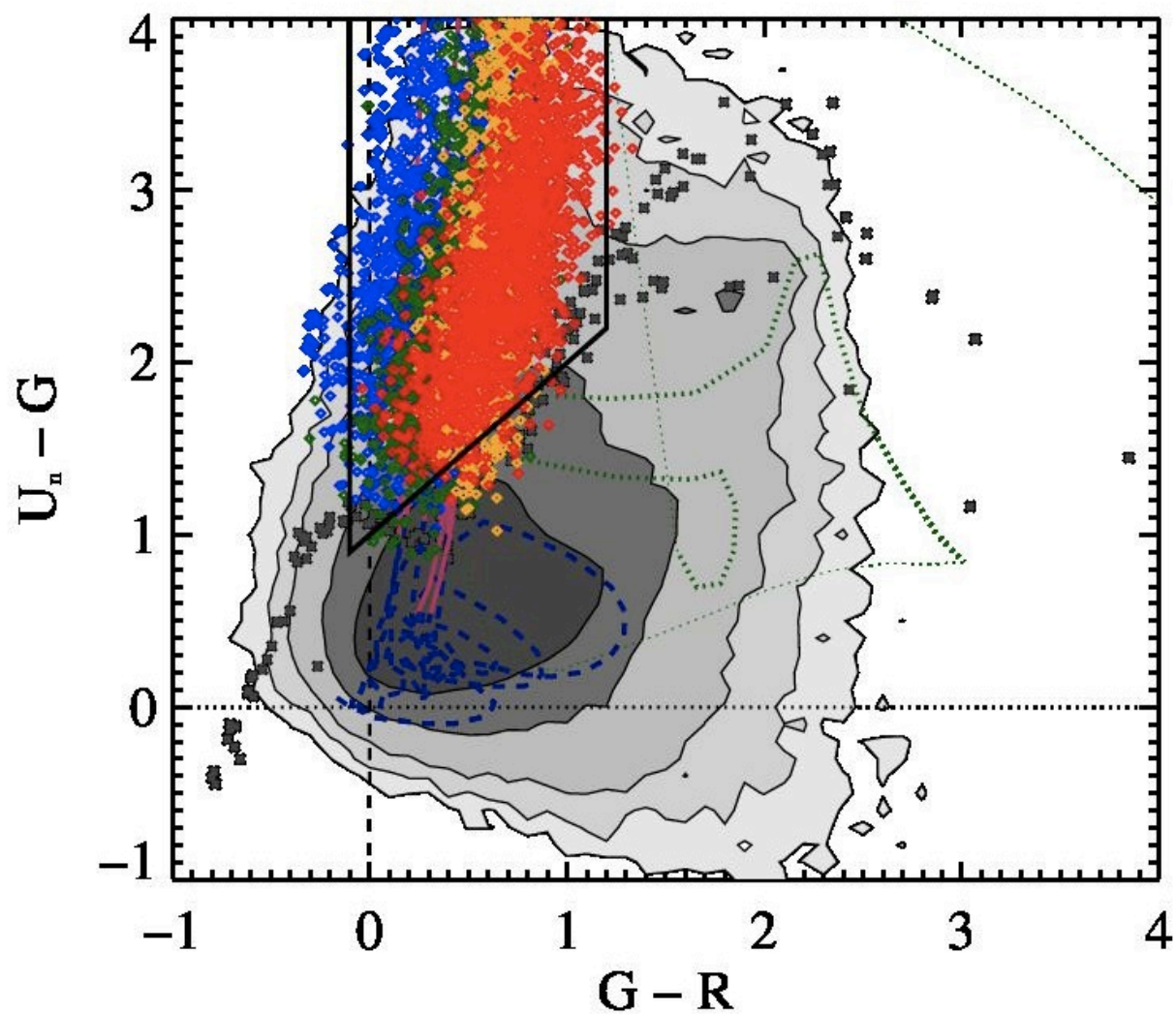
U_nGR-selected z~3 LBGs composite spectra (Shapley et al. 2003)





rounding up the usual suspects

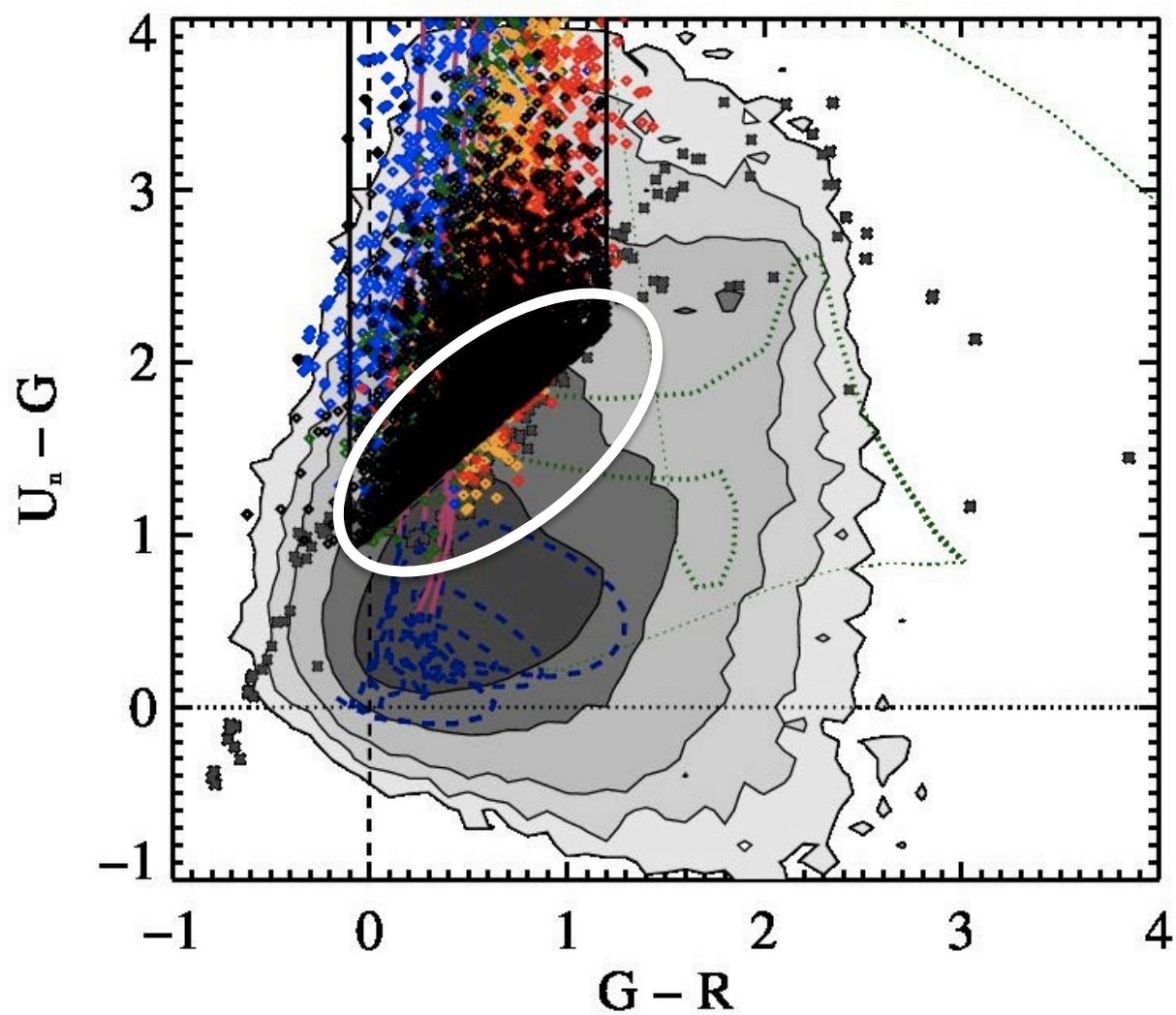
U_n GR-selected $z \sim 3$ LBGs (e.g., Steidel et al. 1996, 2003)





questioning the suspects

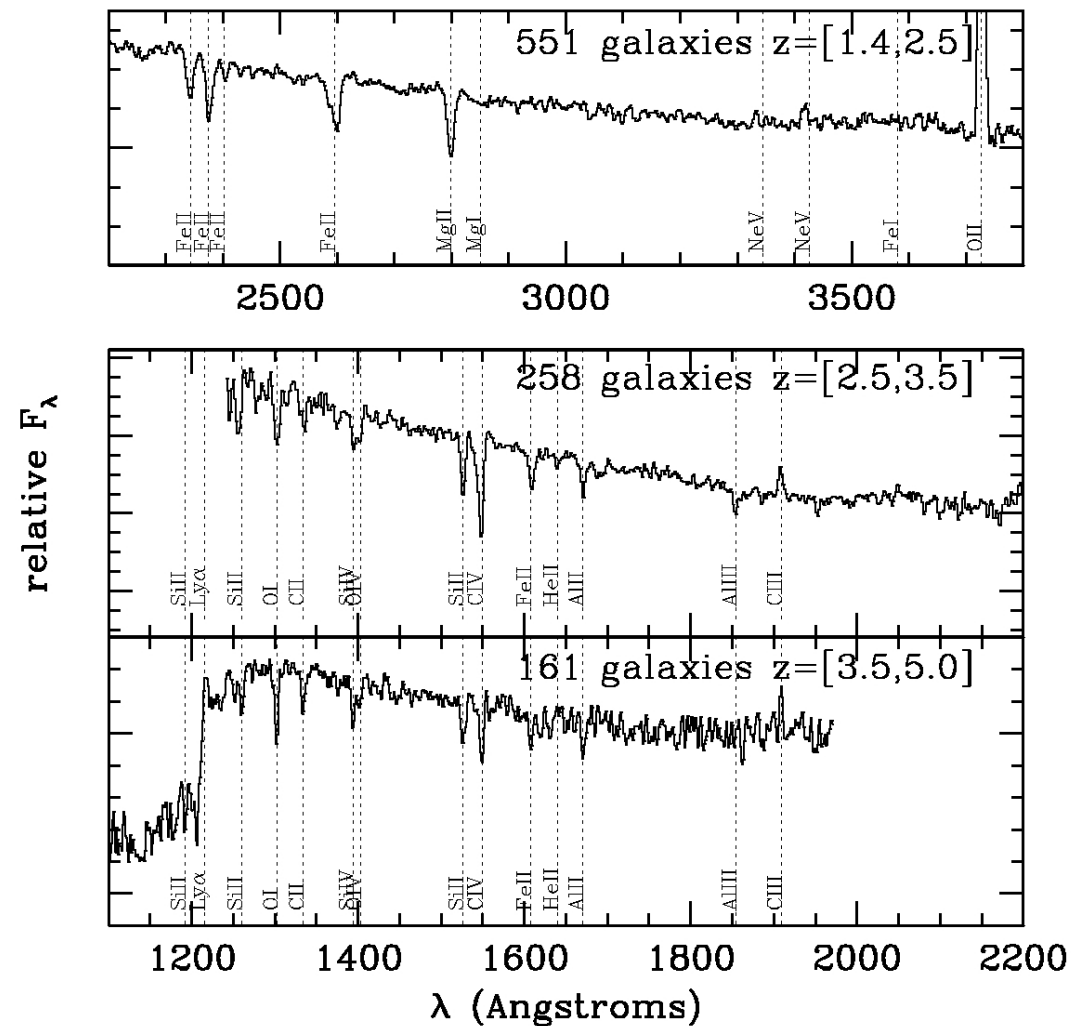
U_n GR-selected $z \sim 3$ LBGs (e.g., Steidel et al. 1996, 2003)





circumstantial evidence

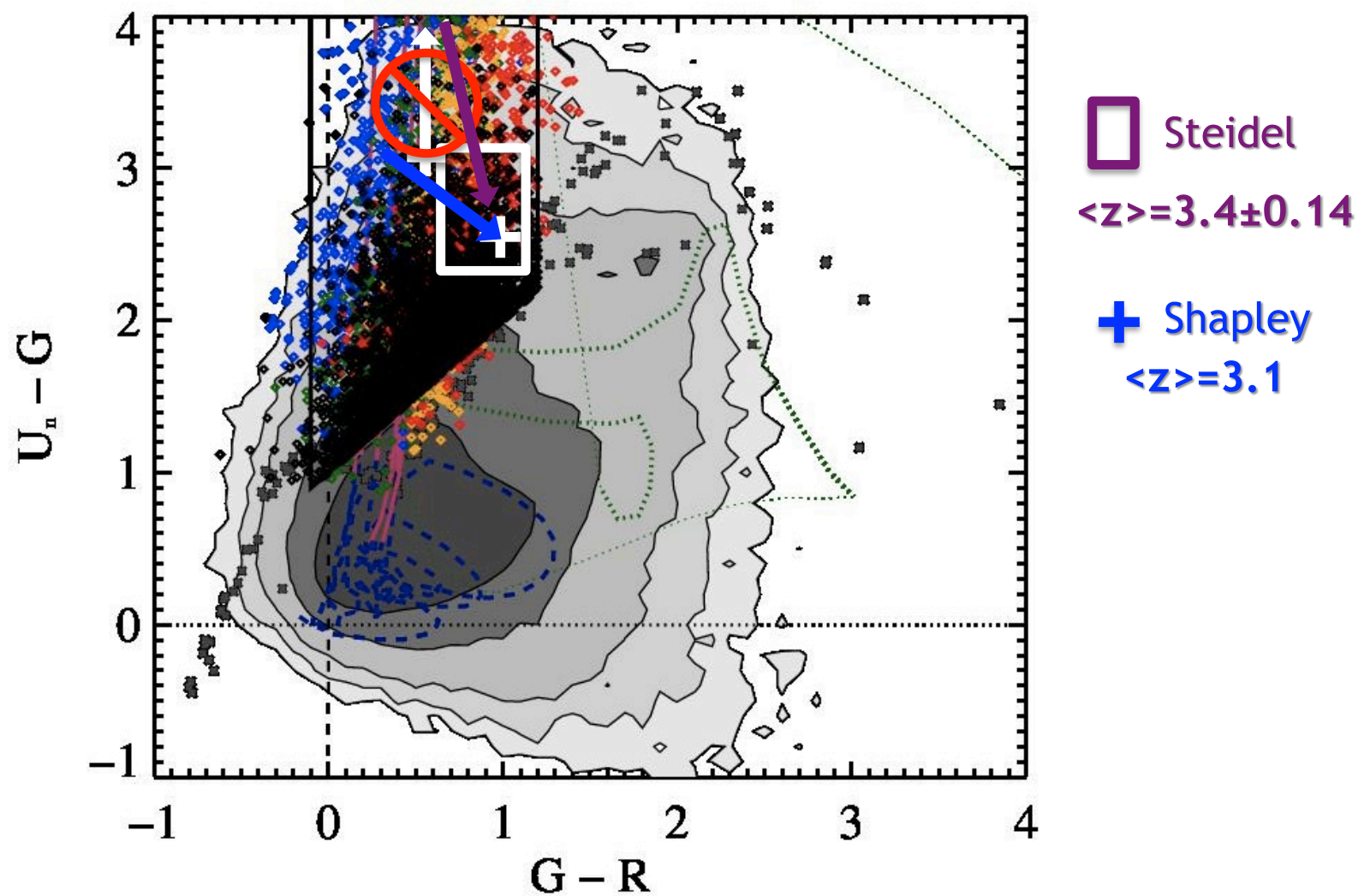
VVDS: $I_{AB} < 24$ selected unbiased spectroscopic survey (Le Fevre et al. 2005)

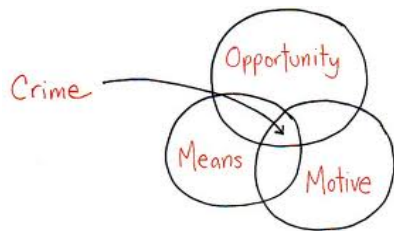




circumstantial evidence

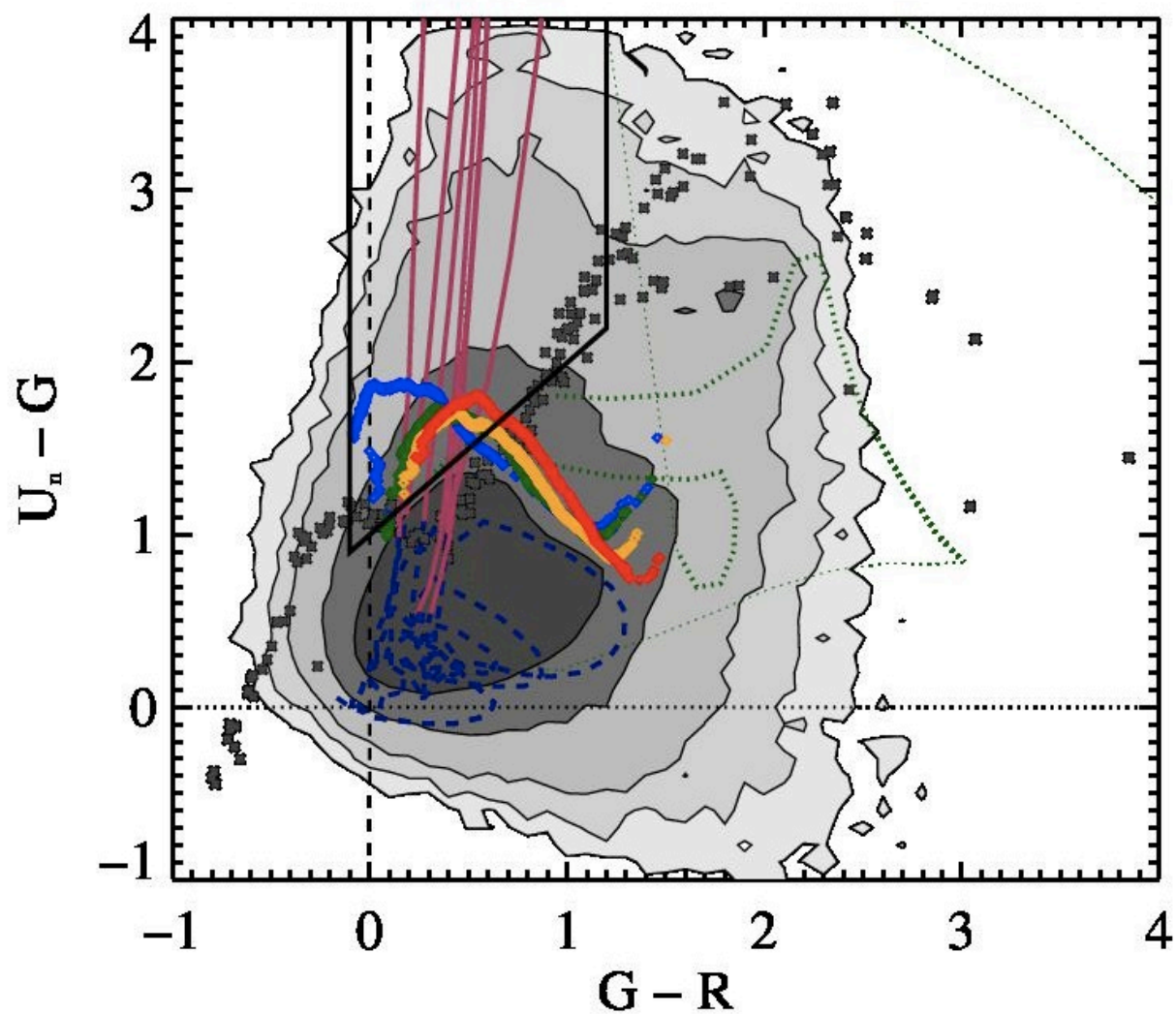
f_{esc} searches (Steidel et al. 2001, Shapley et al. 2006)

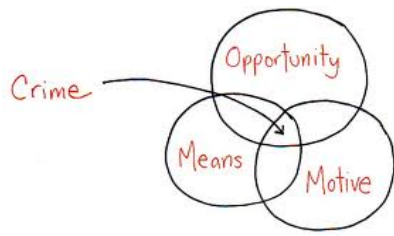




means, motive, and opportunity

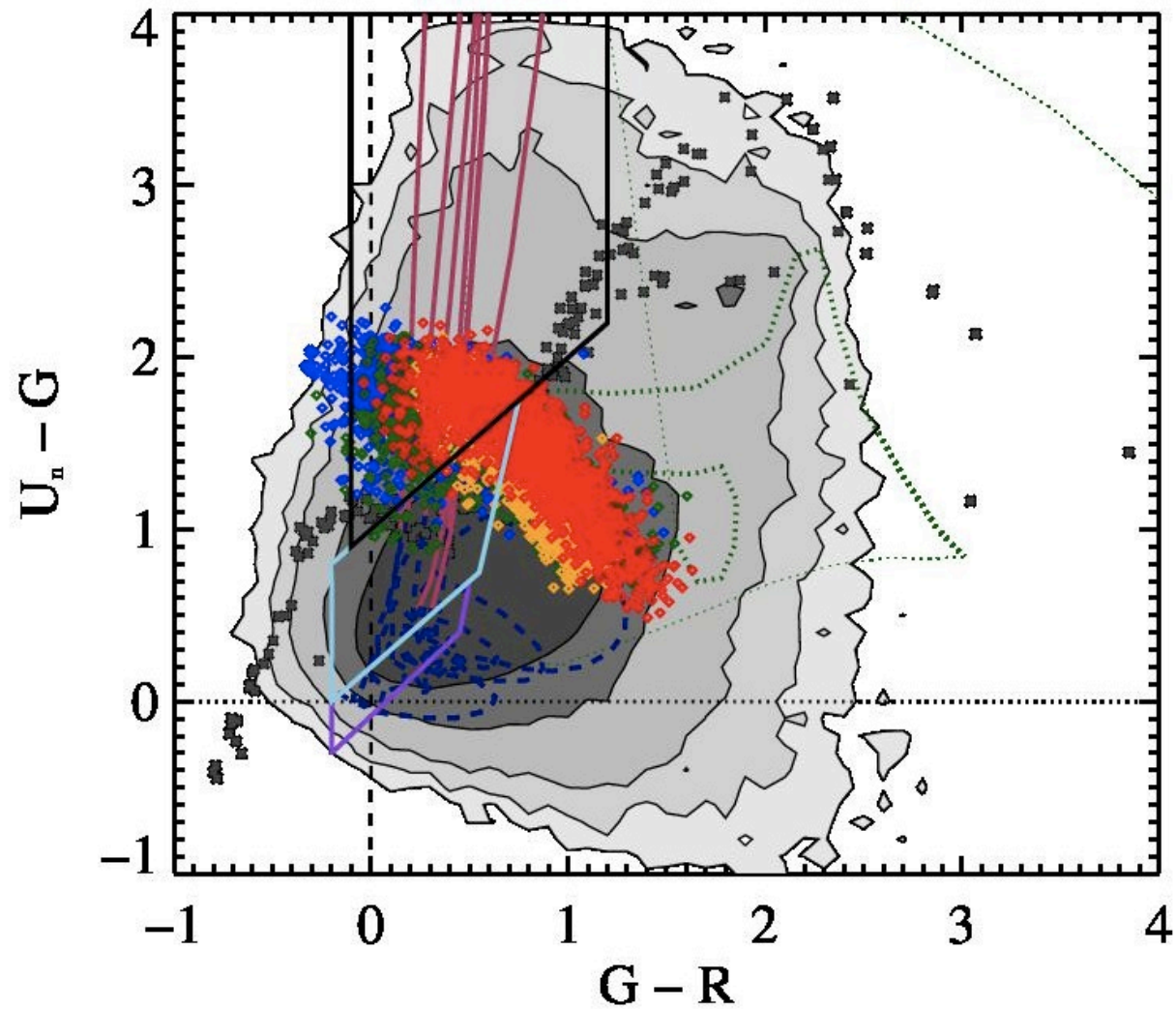
f_{esc} composite simulations (Cooke, et al, 2013, in prep)

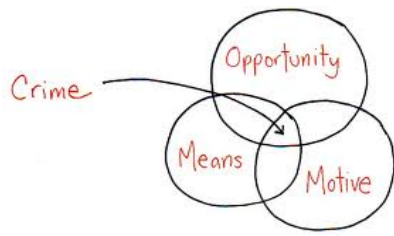




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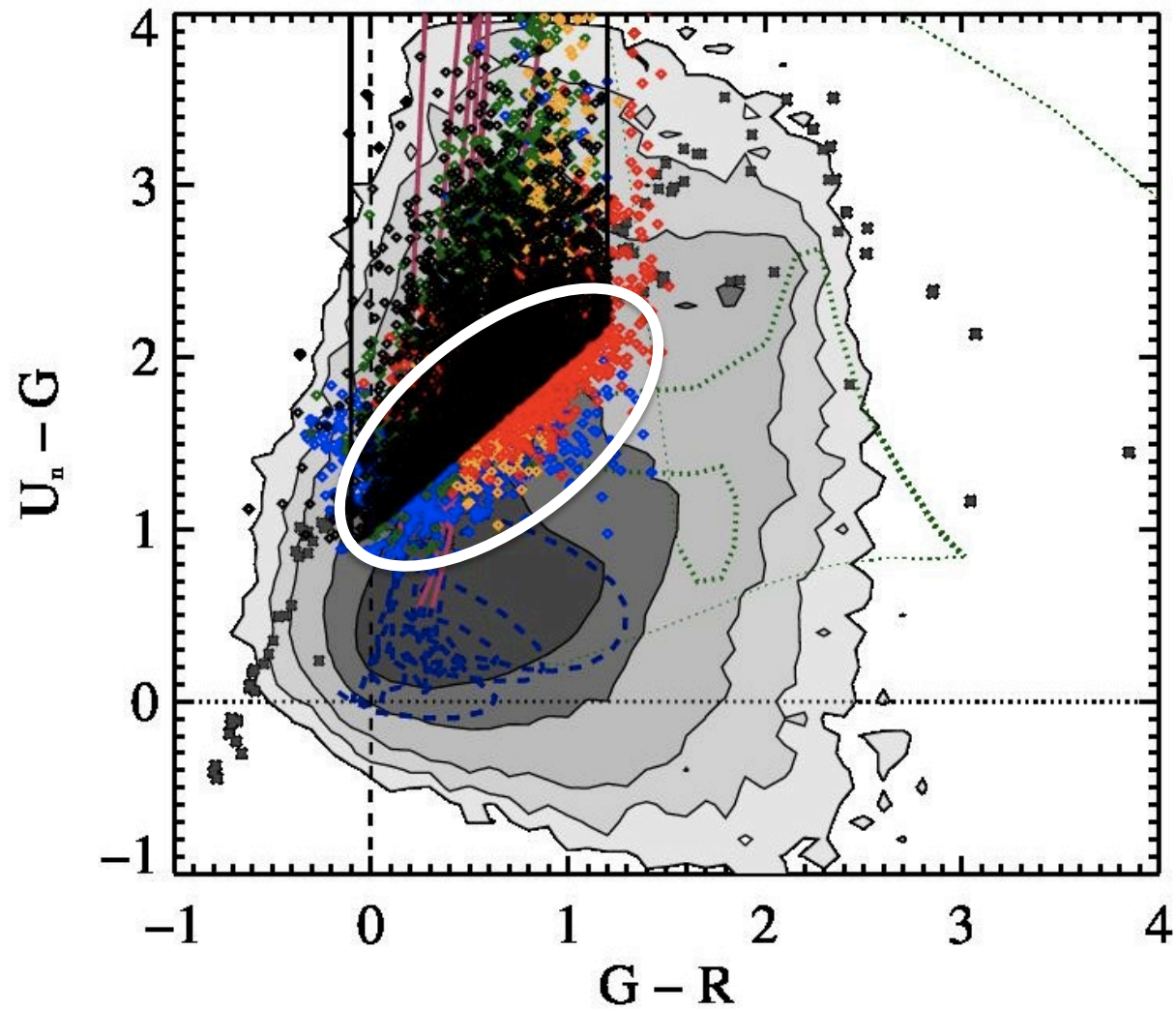
f_{esc} composite simulations (Cooke, et al, 2013, in prep)

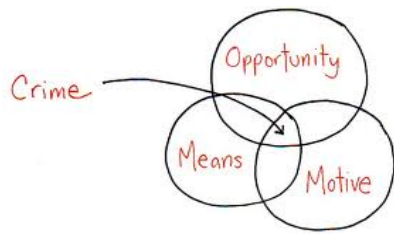




means, motive, and opportunity

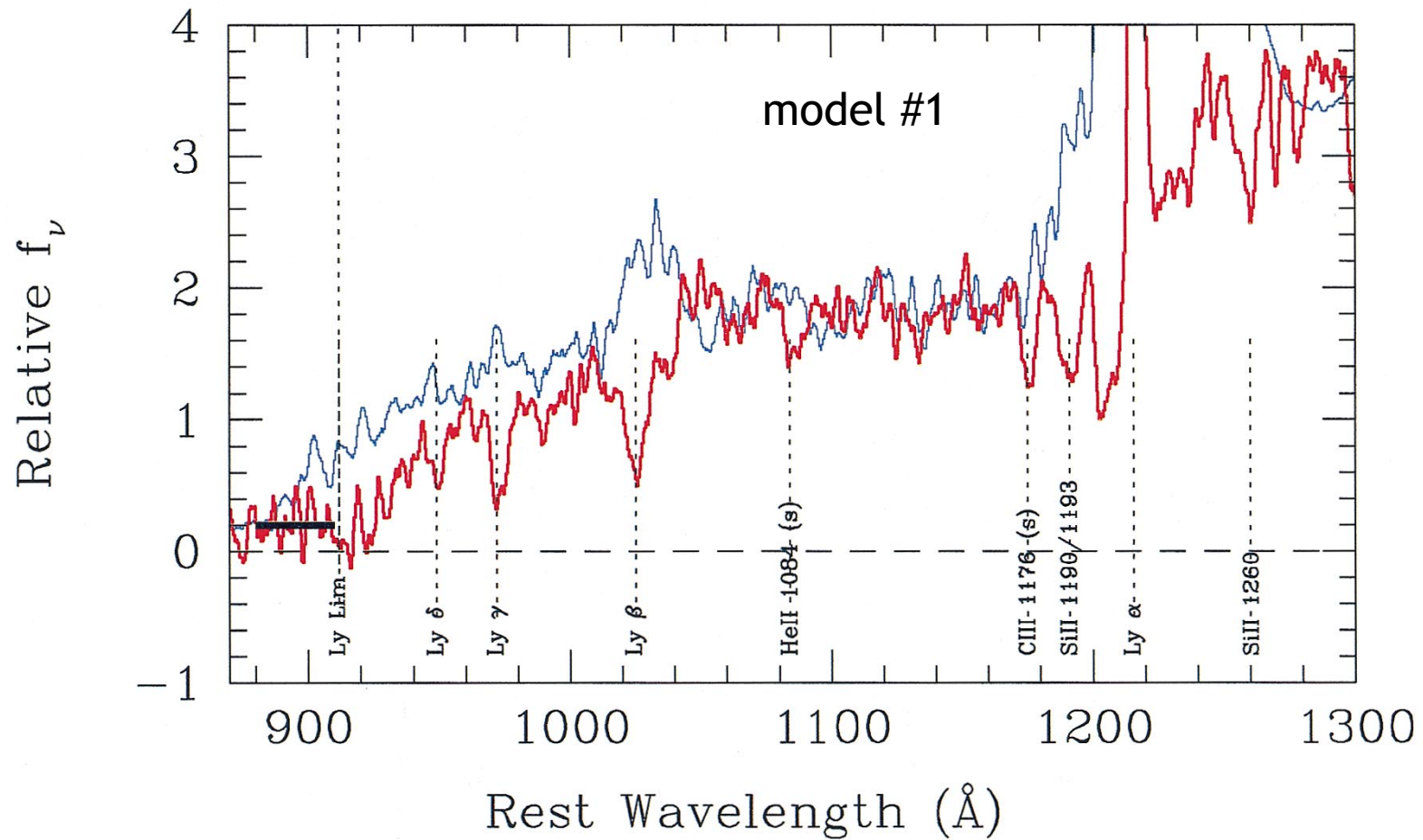
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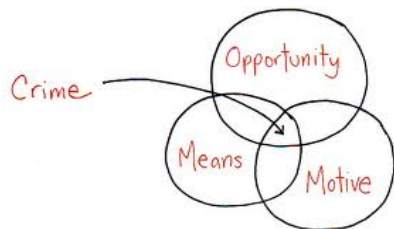




means, motive, and opportunity

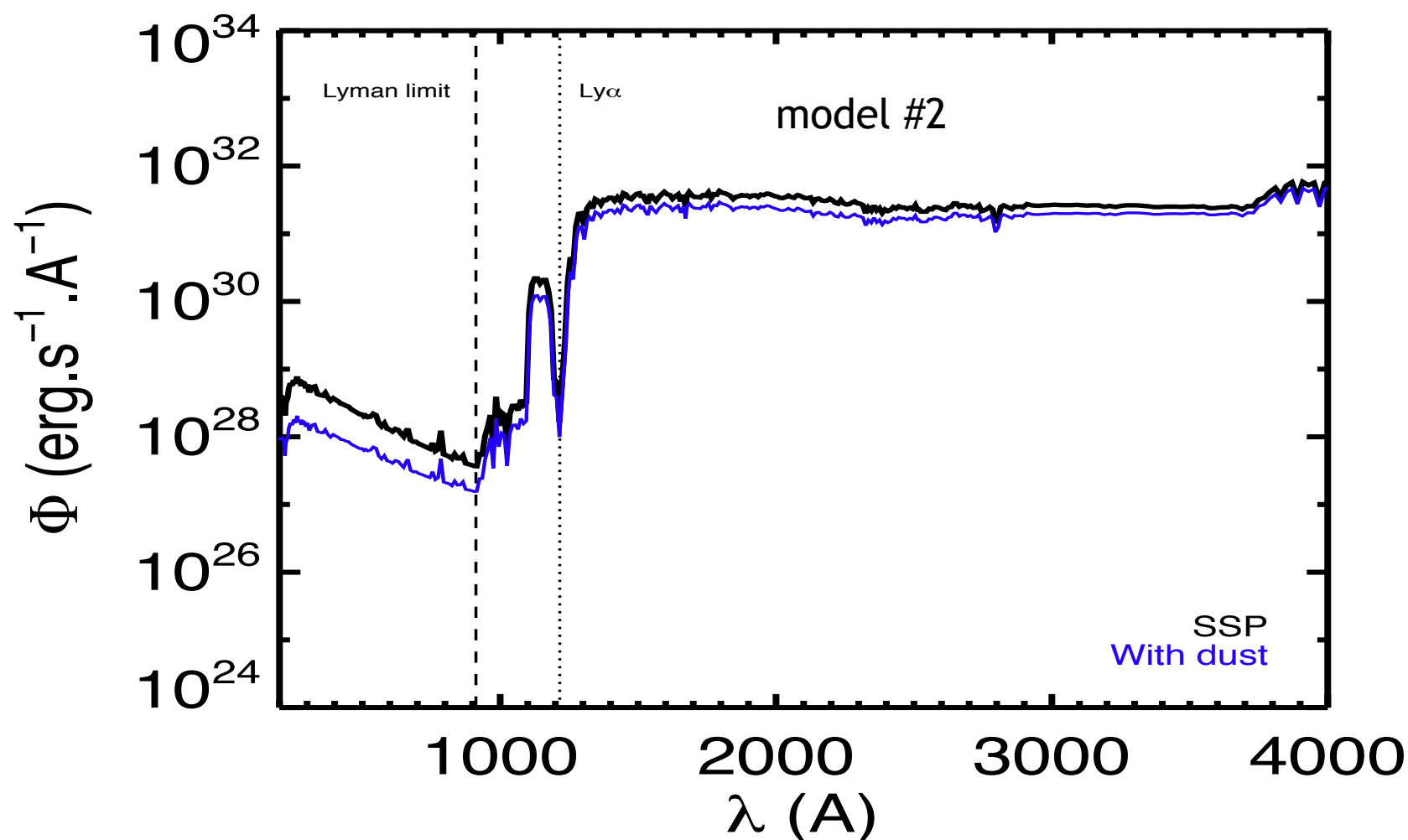
f_{esc} definitions (Steidel et al. 2001, Shapley et al. 2006)





means, motive, and opportunity

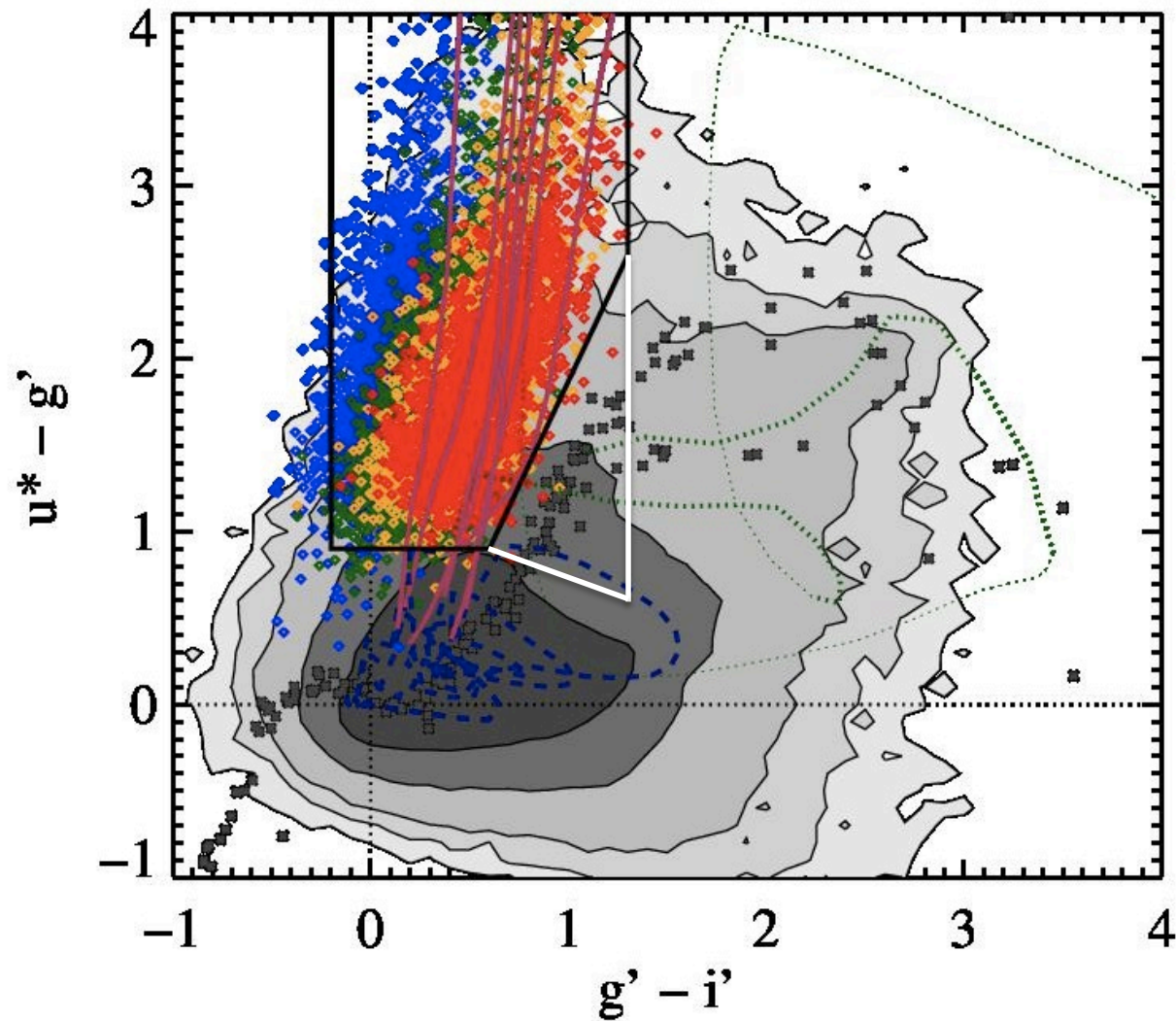
f_{esc} definitions





catching the real culprit

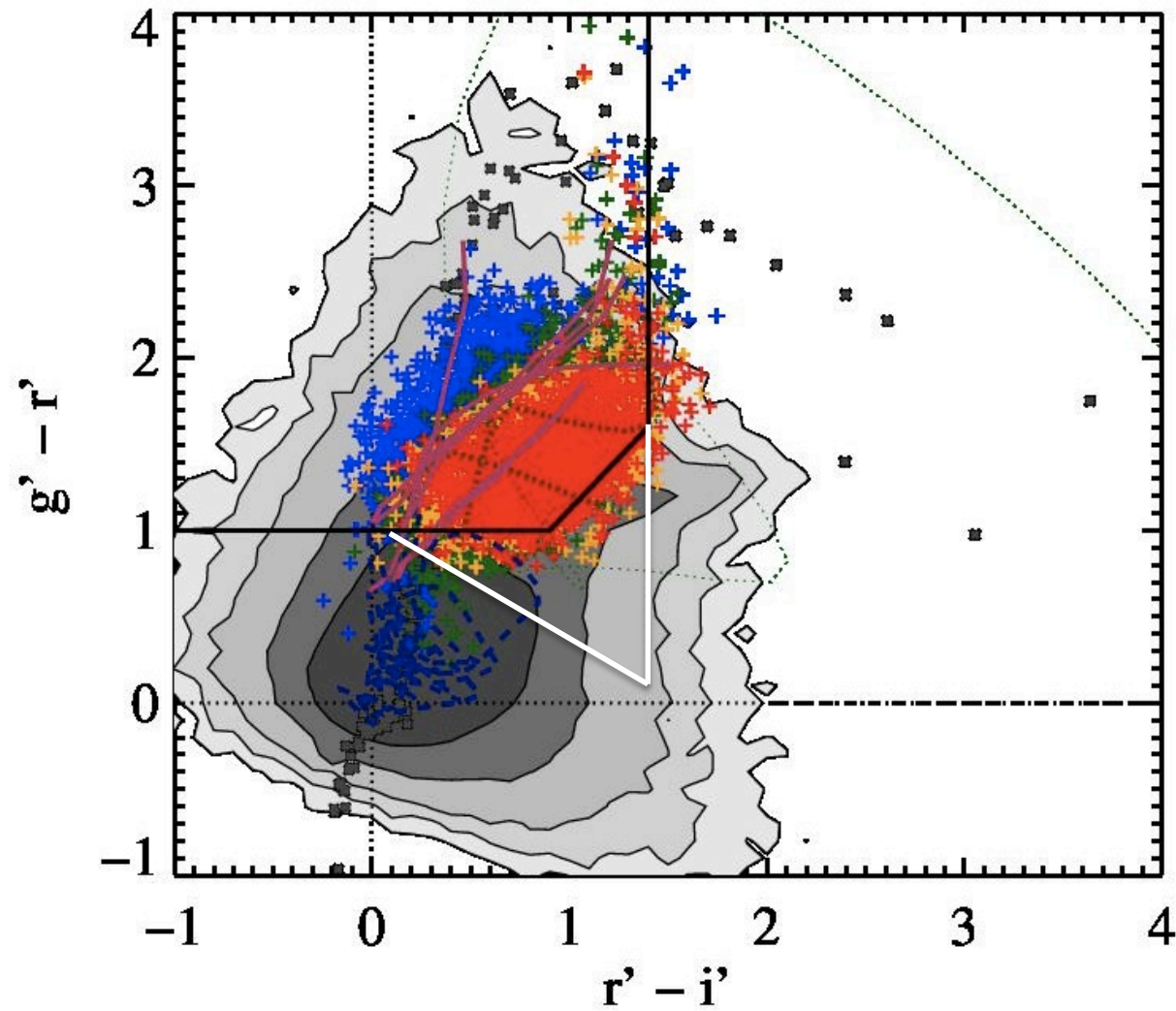
$z \sim 3$ f_{esc} search region





catching the real culprit

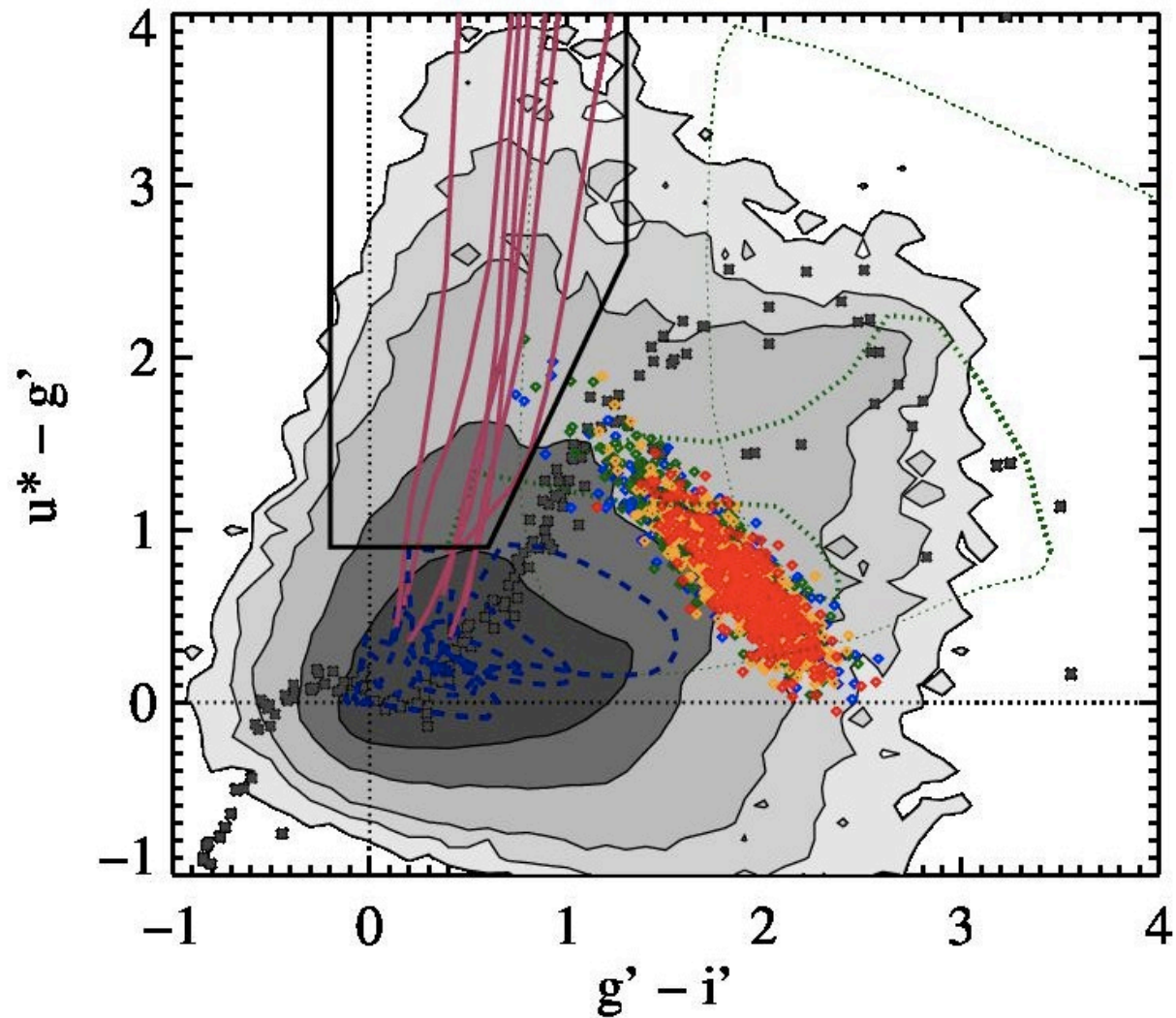
$z \sim 4$ f_{esc} search region





catching the real culprit

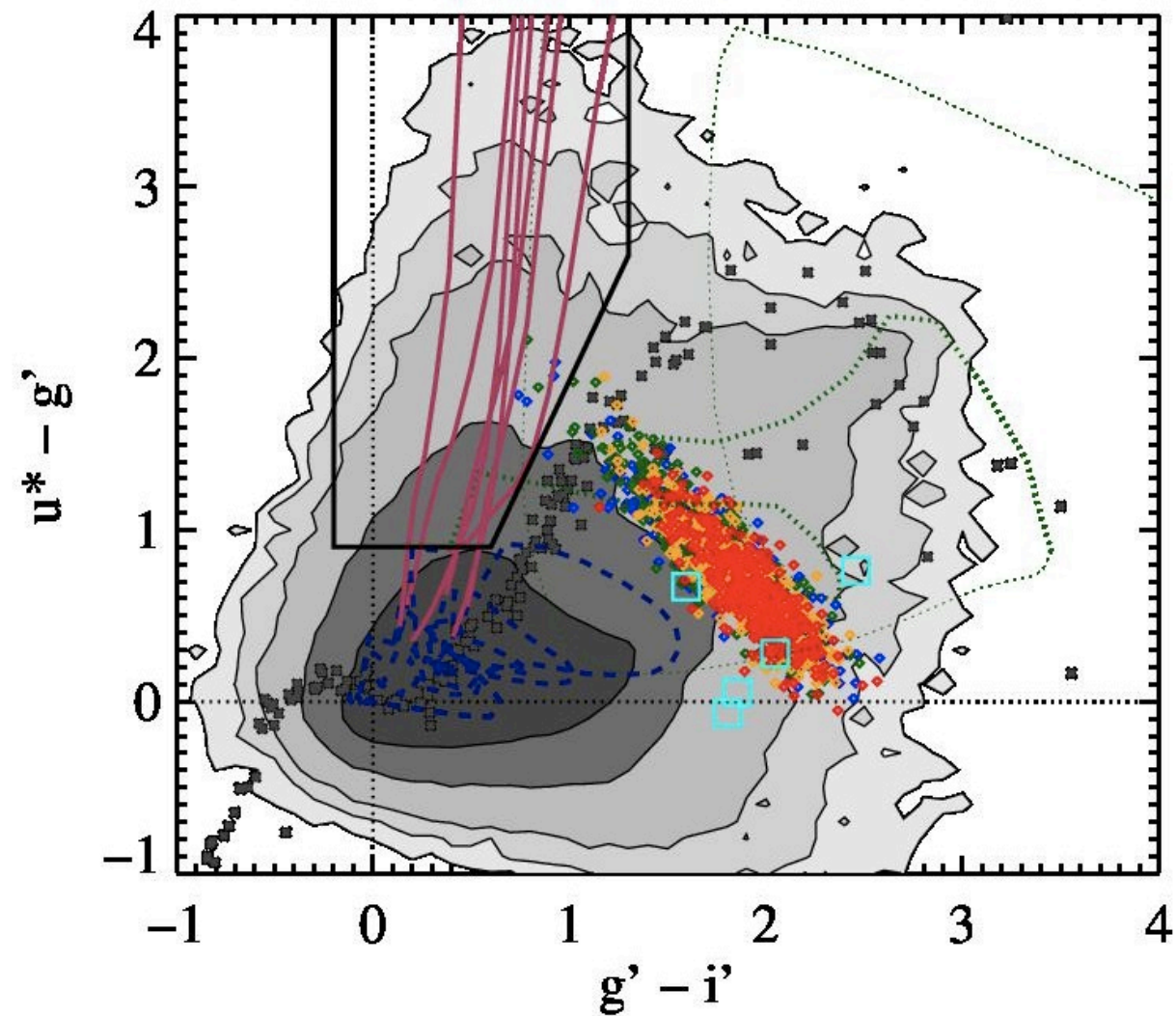
$z \sim 4$ f_{esc} in the $(z \sim 3)$ u^*-g' vs. $g'-i'$ color-color plane





standing trial

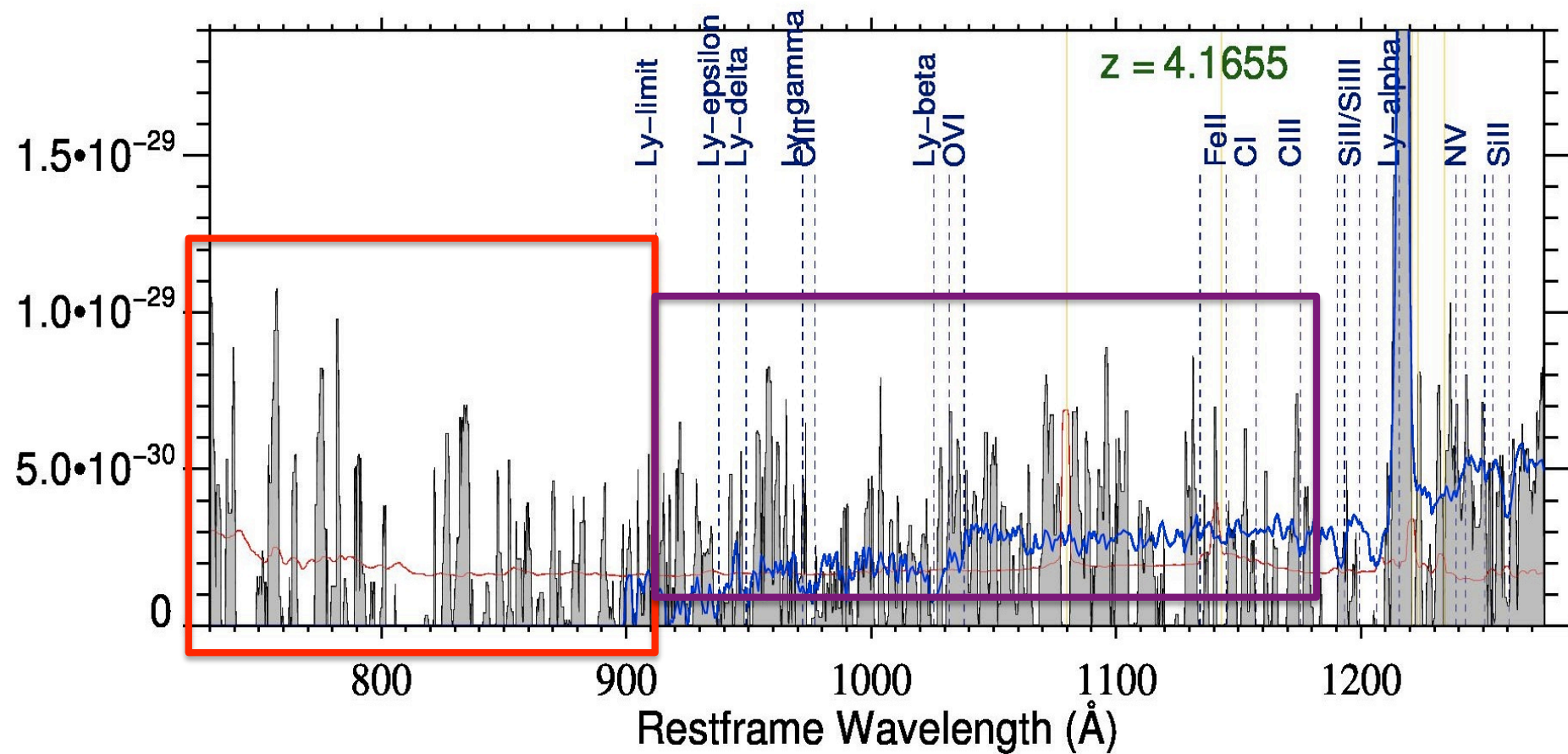
$z \sim 4$ f_{esc} candidate population confirmation





standing trial

$z \sim 4$ f_{esc} candidate population confirmation





the verdict

Result 1.) The distribution of conventionally-selected Lyman break galaxies indicates that a significant fraction of high- z galaxies are missed

Result 2.) By definition, the Lyman break technique selects *against* galaxies with high escape fractions (*u-band bright*)

Result 3) A significant population of galaxies in the “missed” region outside conventional color criteria has been spectroscopically identified

Result 4) The *only* explanation for the galaxies found outside the criteria is stronger 600-1050Å flux, hence Lyman continuum photons

Result 5) Search for these high- f_{esc} galaxies is crucial for the full census of high- z galaxies and the key contributors to reionization