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Coherent emission from CR air showers

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Summary of Talk

- ▶ Observations of EASs
- ▶ Cerenkov, geomagnetic & geo-synchrotron emission
- ▶ Codes for coherent radio emission in EASs
- ▶ Radio emission in air
- ▶ Implications for astrophysics?

Observations of EASs

- ▶ EASs from CR primaries $E \gtrsim 10^{15}$ eV
 $N \sim 10^6, 10^4 \text{ m}^2$ at $E \sim 10^{15}$ eV; $N \sim 10^{11}, 10^7 \text{ m}^2$ at $E \sim 10^{20}$ eV
- ▶ Detections of HECRs (e.g., Pierre Auger Observatory)
- ▶ Cascade includes e^\pm pairs (plus $\mu^\pm, \pi^{0,\pm}, \dots$)
- ▶ e^\pm exceed phase speed of light in air ($E \gtrsim 20 \text{ MeV}$)
- ▶ \Rightarrow emit Cerenkov light
- ▶ Detectable only on cloudless, moonless nights

Askaryan effect

- ▶ Possibility of Cerenkov radio emission from EASs (Jelley 1958)
- ▶ Coherent if $\lambda \gg$ thickness of EAS (ℓ)
- ▶ Askaryan recognized $N_{e+} \neq N_{e-}$ (Askaryan 1962, 1964)
allows coherent emission $\propto (N_{e-} - N_{e+})^2$
- ▶ Effect confirmed observationally (Jelley *et al.* 1965)

Geomagnetic & geo-synchrotron emission

Early theory (Kahn & Lerche 1966)

- ▶ $N_e = 10^6$, $\ell = 1$ m, $w = 30$ m, $E_e = 100$ MeV
- ▶ Individual e^\pm lifetime 10^{-6} s, continuously replenished
- ▶ Emission mechanisms: Cerenkov, geomagnetic, geo-synchrotron, creation/annihilation, ...

Geomagnetic emission

- ▶ Separation of e^+ and e^- by Earth's **B**
- ▶ \Rightarrow changing dipole moment \Rightarrow radiation

Geo-synchrotron emission

- ▶ Modified form of synchrotron emission due to Earth's **B**
 - ▶ $\tau \ll$ period of one gyration; breaks azimuthal symmetry
 - ▶ effect of refractive index of air
- ▶ Difficult to disentangle from geomagnetic emission

Codes for coherent radio emission in EASs

Monte Carlo simulation codes

- ▶ REAS1, AIRES, CoREAS, ZHAireS, ...
- ▶ Collection of individual charges: specified \mathbf{x}_0 , β , ...
- ▶ Include all relevant accelerations in $\dot{\beta}$
- ▶ Sum electric and magnetic fields,

$$\mathbf{E} = \frac{\pm e}{4\pi\epsilon_0} \left[\frac{\mathbf{n} - \beta}{\gamma^2(1 - \beta \cdot \mathbf{n})^3 R^2} + \frac{\mathbf{n} \times [(\mathbf{n} - \beta) \times \dot{\beta}]}{(1 - \beta \cdot \mathbf{n})^3 R c} \right]_{\text{ret}}, \quad \mathbf{B} = \frac{\mathbf{n} \times \mathbf{E}}{c},$$

$\mathbf{n} = \mathbf{k}c/\omega$, over all particles

- ▶ Calculate Poynting vector, $\mathbf{E} \times \mathbf{B}/\mu_0$

Features of models

- ▶ Consistent with emission $\propto N^2$
- ▶ Geomagnetic/geo-synchrotron emission dominates

Radio emission in air

Cerenkov condition: $n\beta \cos \theta = 1$

- ▶ $n = 1 + 1/2\gamma_0^2$, $\beta = 1 - 1/2\gamma^2$, $\cos \theta = 1 - \theta^2/2$
 $\Rightarrow \theta^2 = \theta_c^2 - 1/\gamma^2$, $\theta_c^2 = 1/\gamma_0^2$
- ▶ Cerenkov emission: on the Cerenkov cone, $\theta = \theta_c$

Geo-synchrotron emission

- ▶ Occurs both inside ($\theta < \theta_c$) & outside ($\theta > \theta_c$)
- ▶ Synchrotron emission suppressed in plasma (Razin effect)
 $n = 1 - \omega_p^2/\omega^2$, suppression $\omega < \omega_{RT} = 2\omega_p^2/3\omega_B \sin \theta$
- ▶ Synchrotron emission enhanced for $n > 1$; by how much?

Conventional model for synchrotron emission

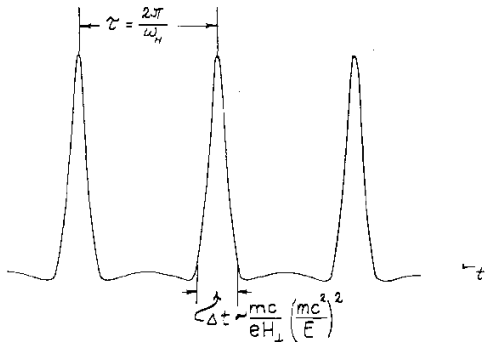


Fig. 4, Ginzburg & Syrovatskii 1965, ARAA

- ▶ One pulse seen by observer each gyration
pulse \Rightarrow frequency spectrum
- ▶ Emission confined to $(\theta - \alpha)^2 \lesssim 1/\gamma^2$
- ▶ $n > 1 \Rightarrow$ Cerenkov cone $(\theta - \alpha)^2 = \theta_c^2 \gg 1/\gamma^2$
- ▶ Broadens & adds structure to each pulse

Implications for astrophysics?

Emission by bunches

- ▶ Early astrophysics theories: coherence = emission by bunches
- ▶ Requires $\lambda \gg \ell$: valid for radio emission in EASs
- ▶ 'Bunch' part of initial conditions:
 - coherent emission is transitory phase
 - analogous to superradiance in quantum optics
- ▶ Emission by bunches unobservable in astrophysics

Role of refractive index $n > 1$

- ▶ Few examples where $n > 1$ important in astrophysics
 - possible exception: HECR passing through Moon
- ▶ Extension of synchrotron theory to $n > 1$
 - primarily of theoretical interest