



Image: Kwon O Chul

YOUNG STAR SCIENCE WITH SKYMAPPER

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ASTROSMURPH

SIMONMURPHY.ME



UNIVERSITÄT
HEIDELBERG
ZUKUNFT
SEIT 1386

BACK TO THE FUTURE...

November 2007

PH.D. THESIS PROPOSAL

Towards a complete census of young stars in the solar neighbourhood with SkyMapper

Simon J. Murphy, RSAA

Supervisor

Prof. Michael Bessell, RSAA

Supervisory panel

Dr. Stefan Keller, RSAA

Dr. Agris Kalnajs, RSAA

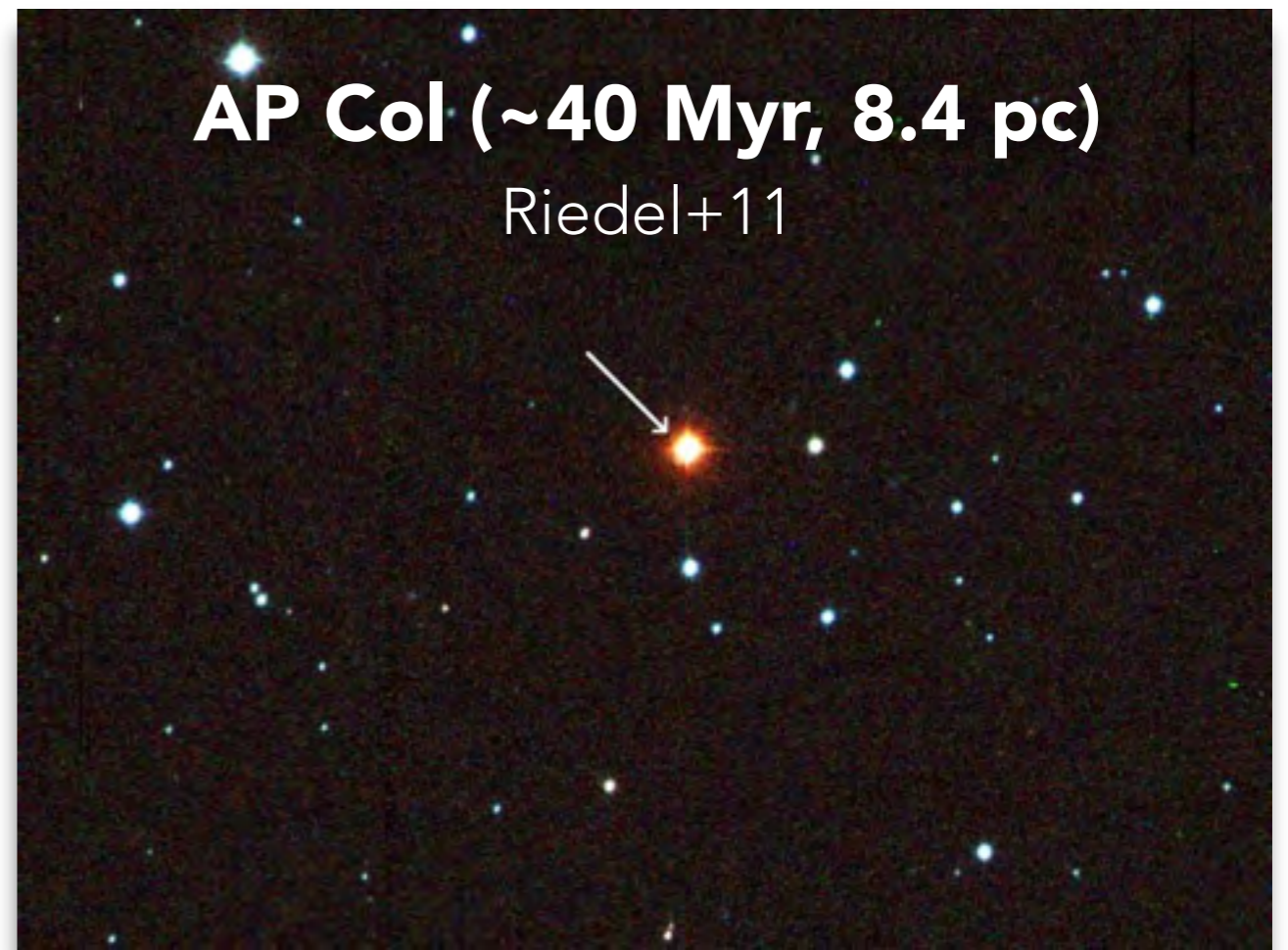
Assoc. Prof. Warrick Lawson, UNSW@ADFA

Dr. Inseok Song, IPAC, Caltech

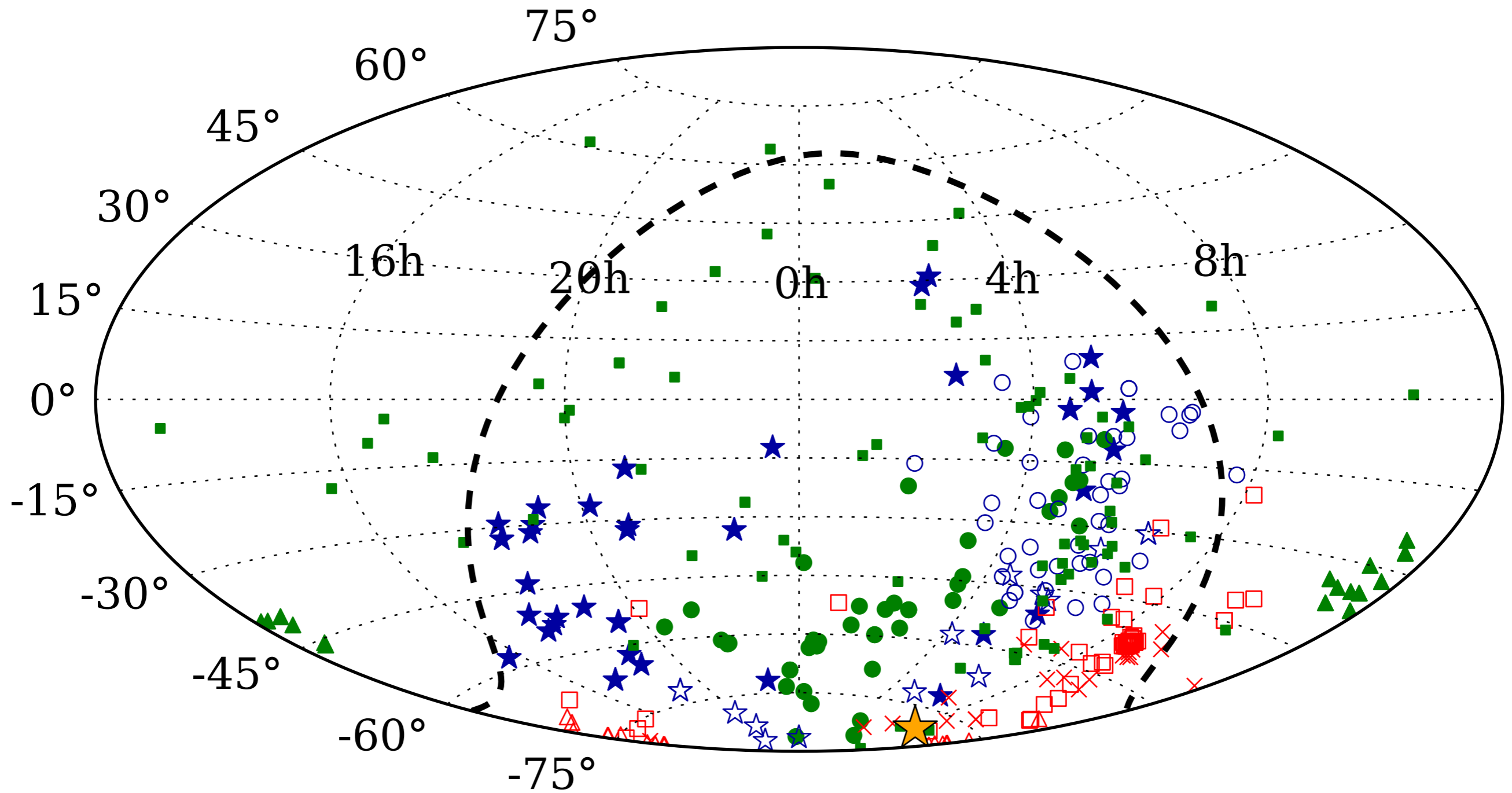
Prof. Ben Zuckerman, UCLA

IDENTIFYING YOUNG STARS

- Wide-field photometry
- Variability
- Astrometry
- **Ha photometry**



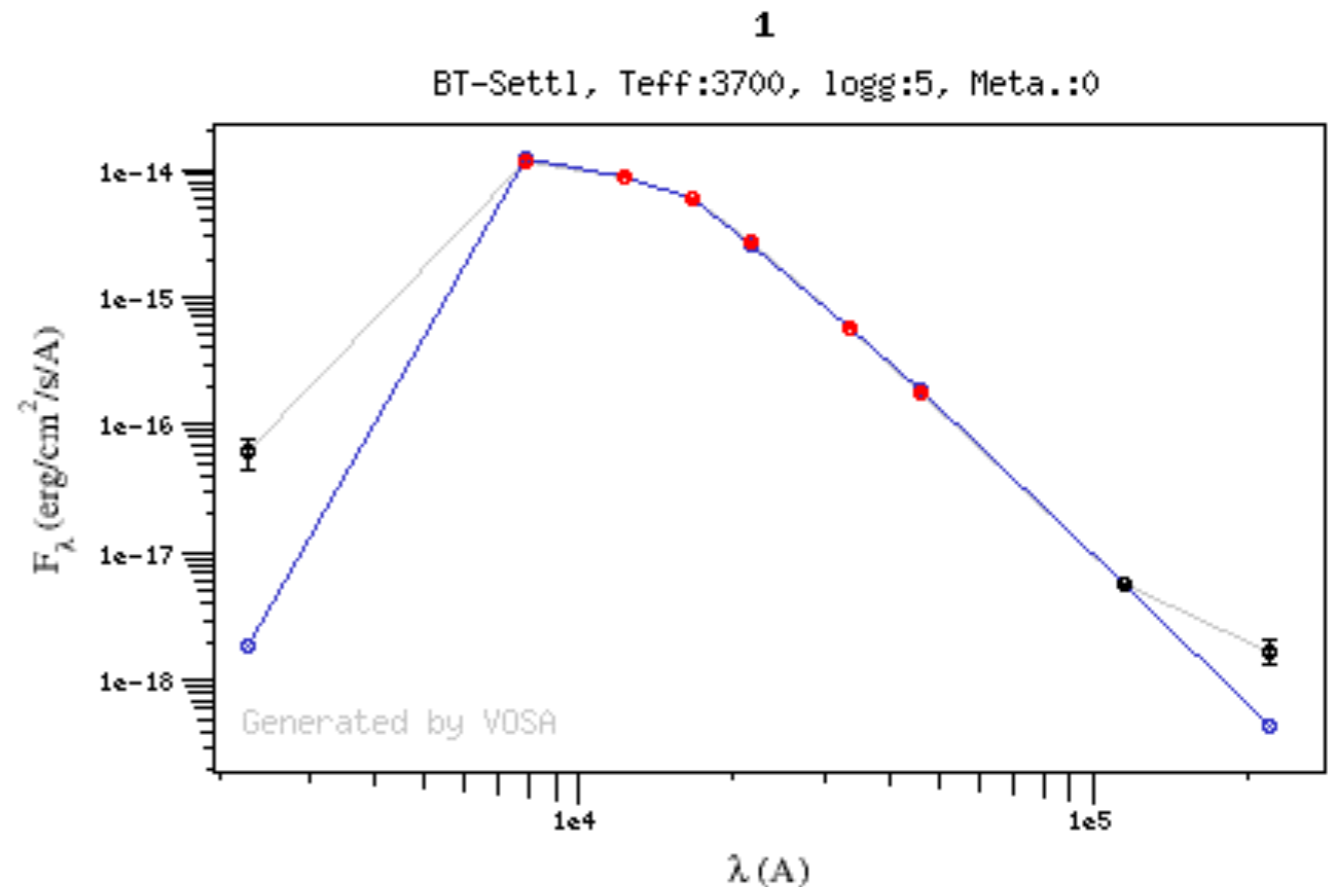
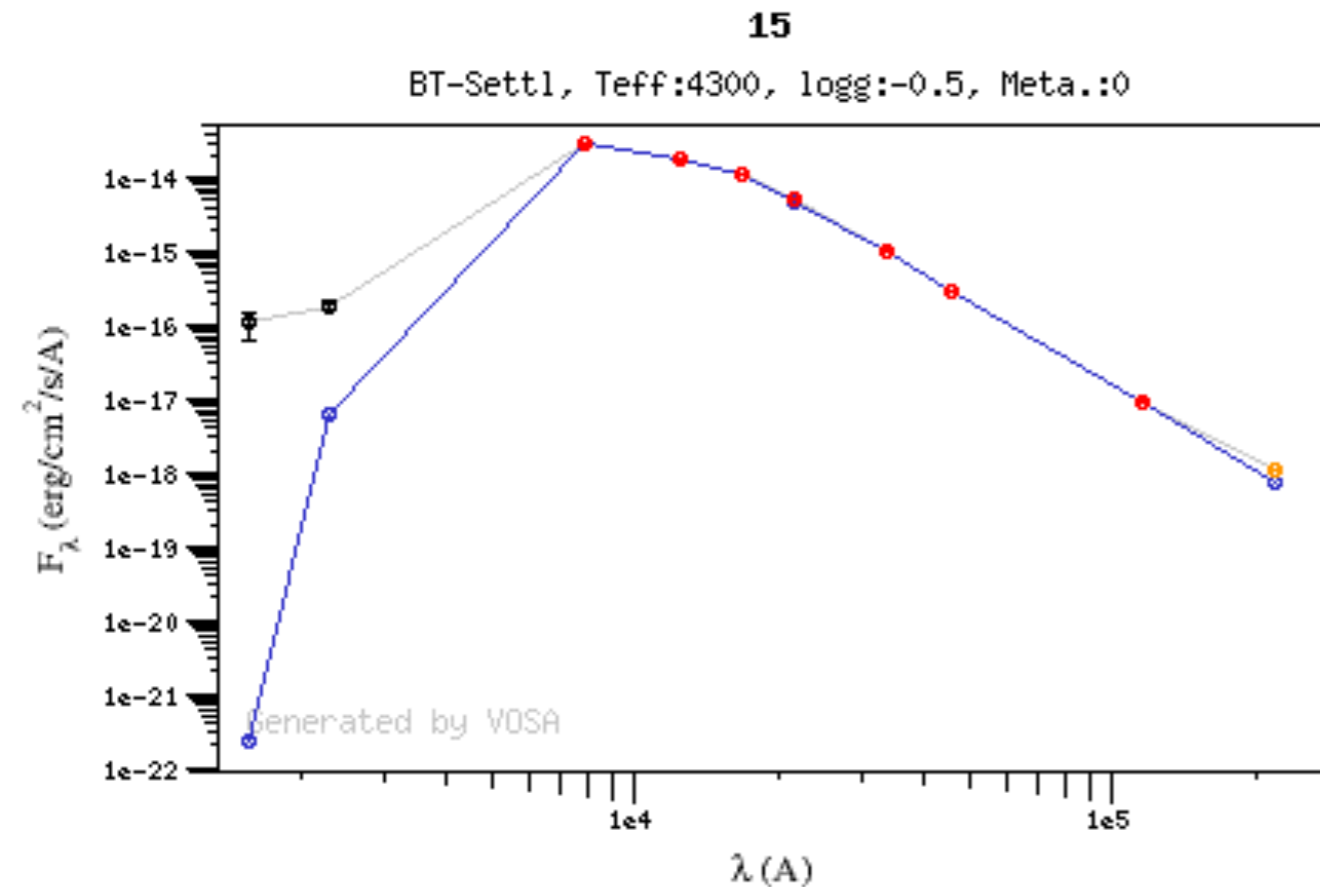
WIDE-FIELD PHOTOMETRY



Torres+08 SACY sample

WIDE-FIELD PHOTOMETRY

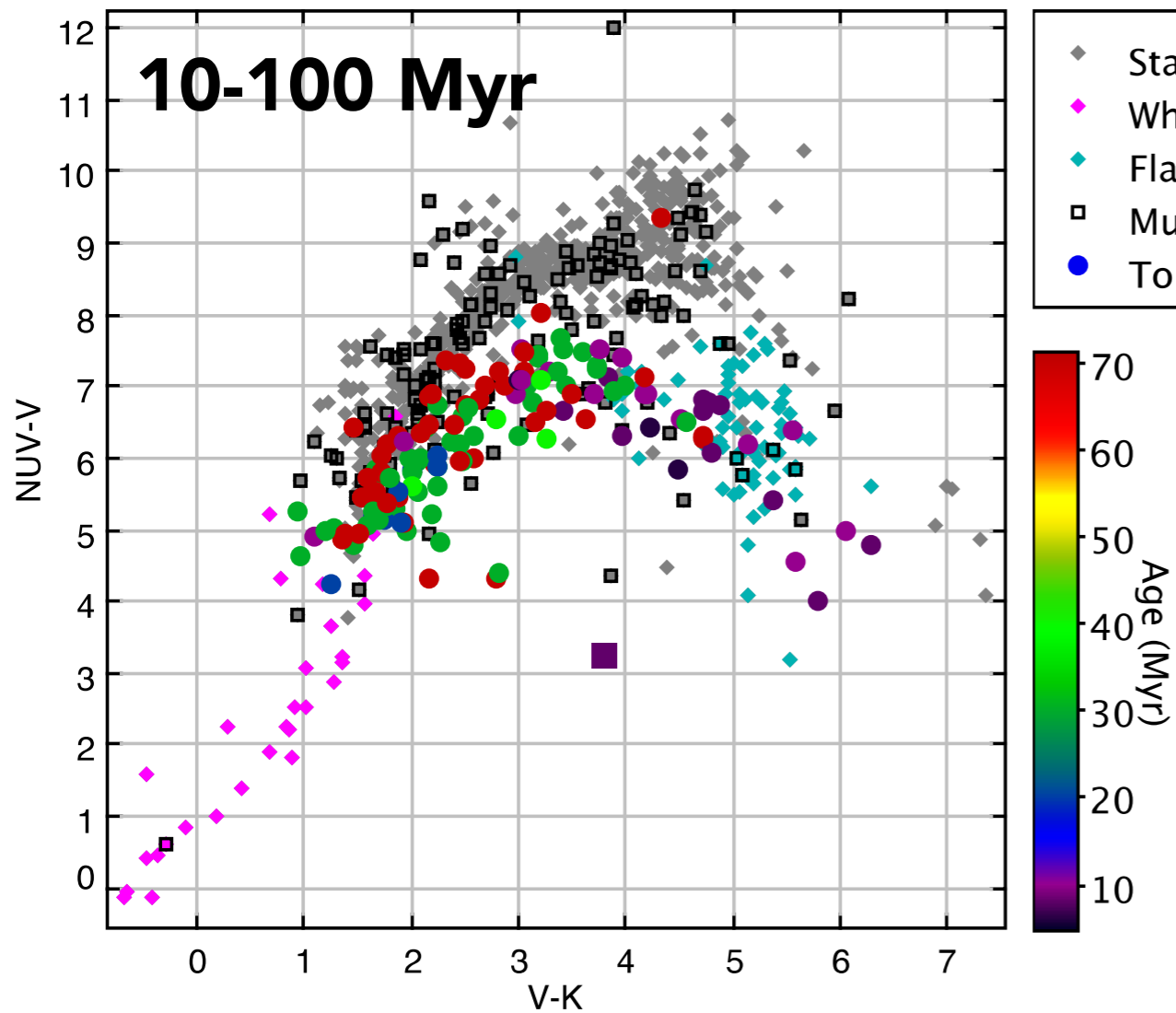
- Combine with GALEX, 2MASS, VHS, WISE, AKARI,....
 - SEDs from $0.3 < \lambda < 160 \mu\text{m}$
 - Disks, extinctions, temperatures, gravities, Fe/H?



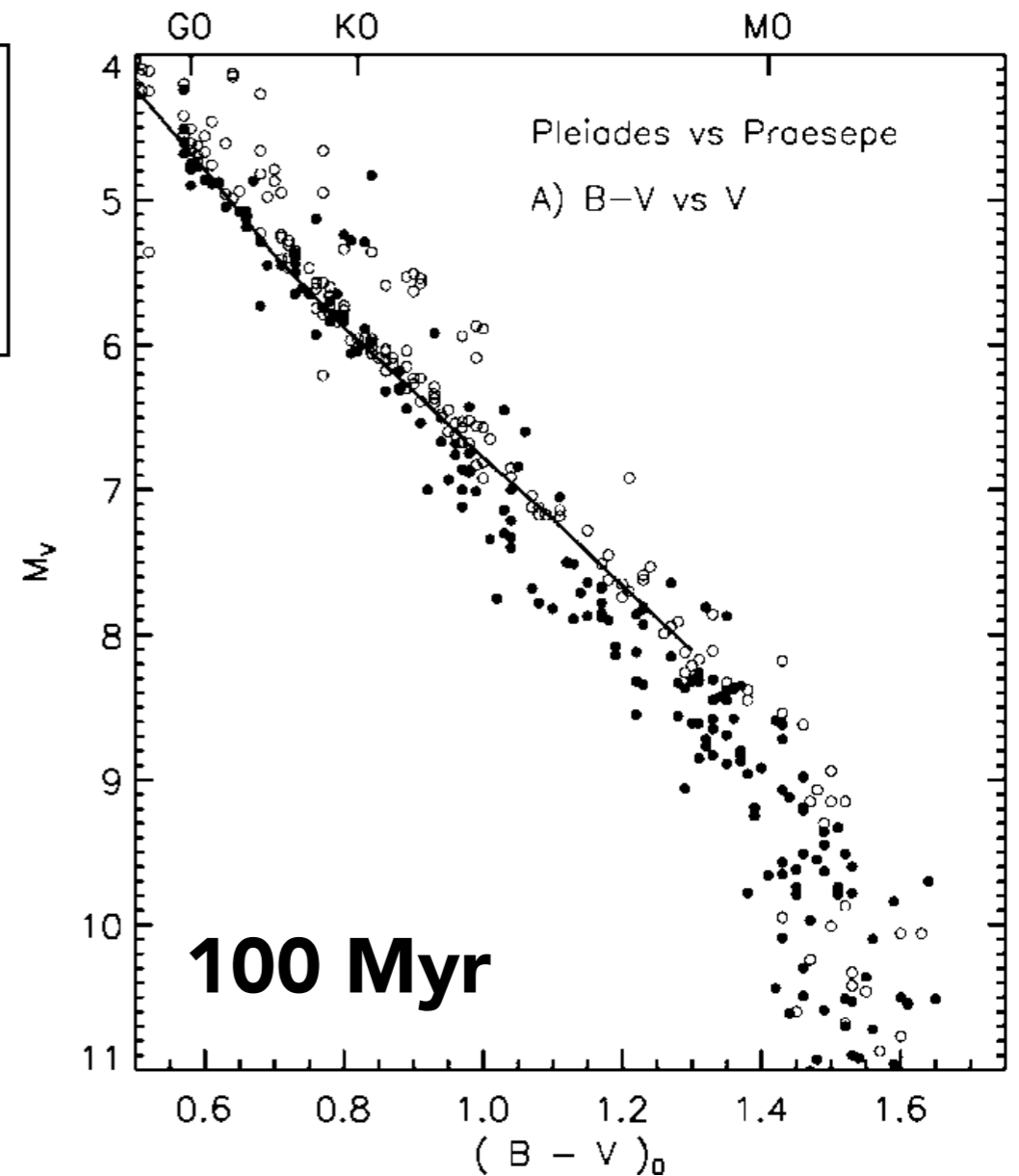
VO SED Analyser (VOSA, <http://svo2.cab.inta-csic.es/theory/vosa/>)

WIDE-FIELD PHOTOMETRY

- UV and blue excess emission (activity, spots, accretion)



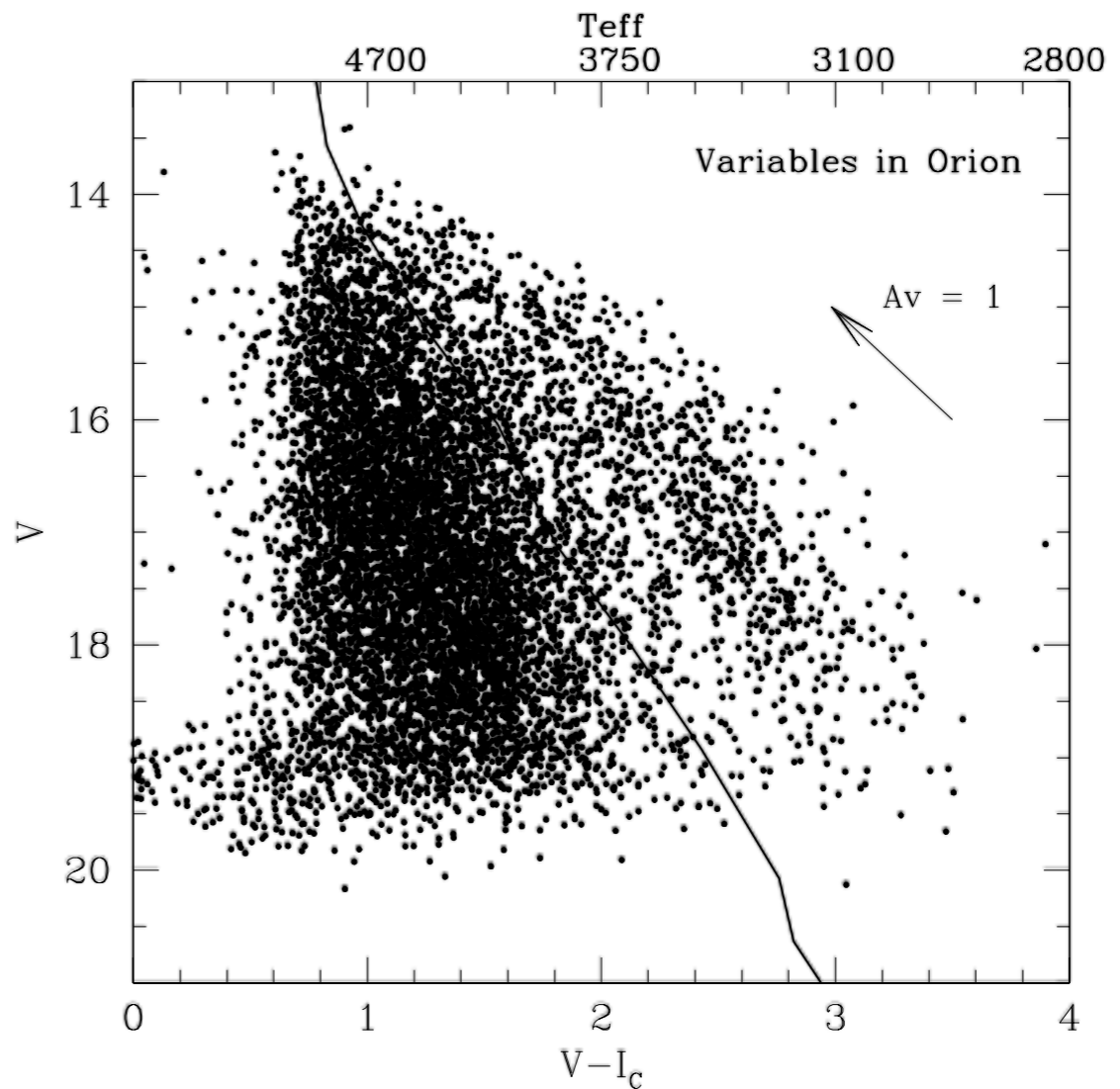
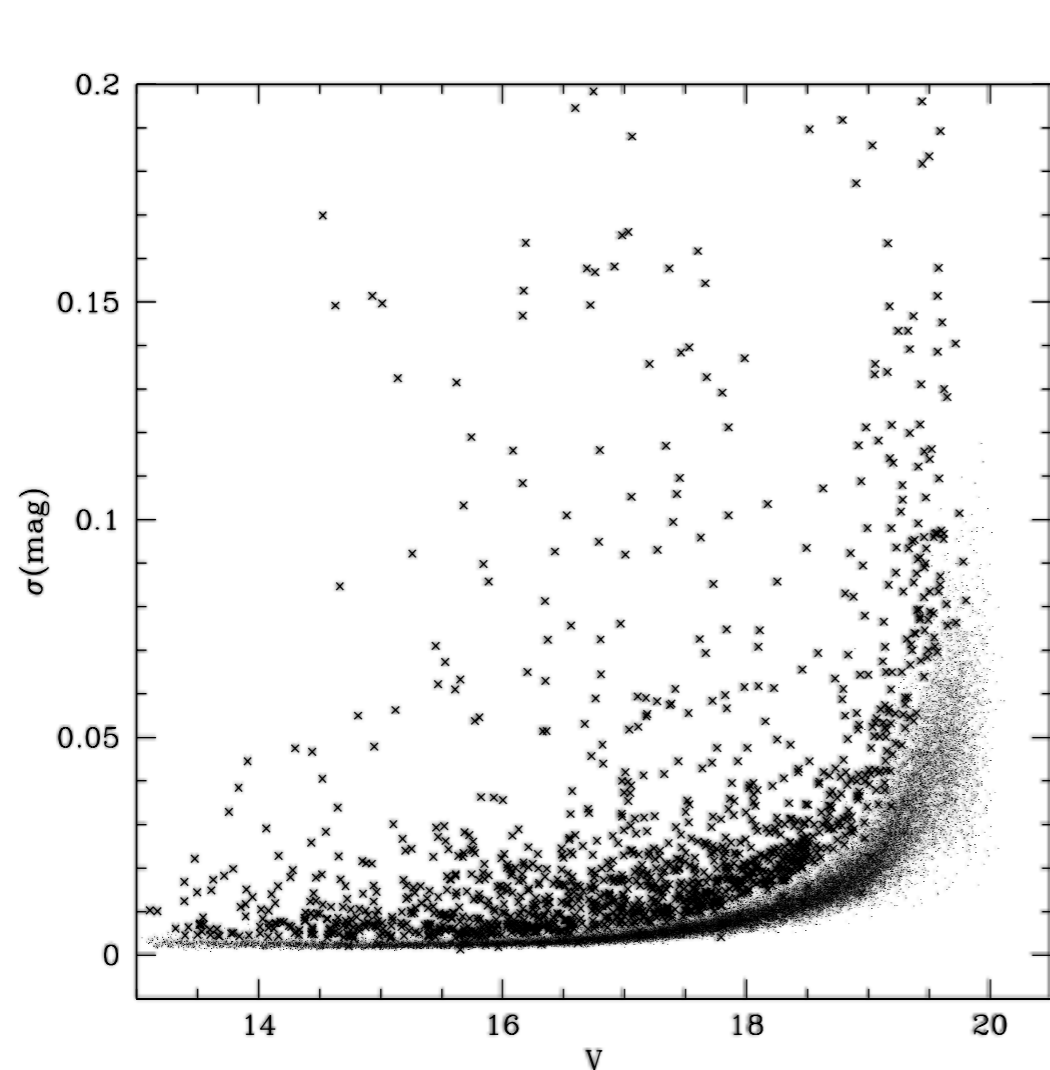
Rodriguez+11



Stauffer+03

VARIABILITY

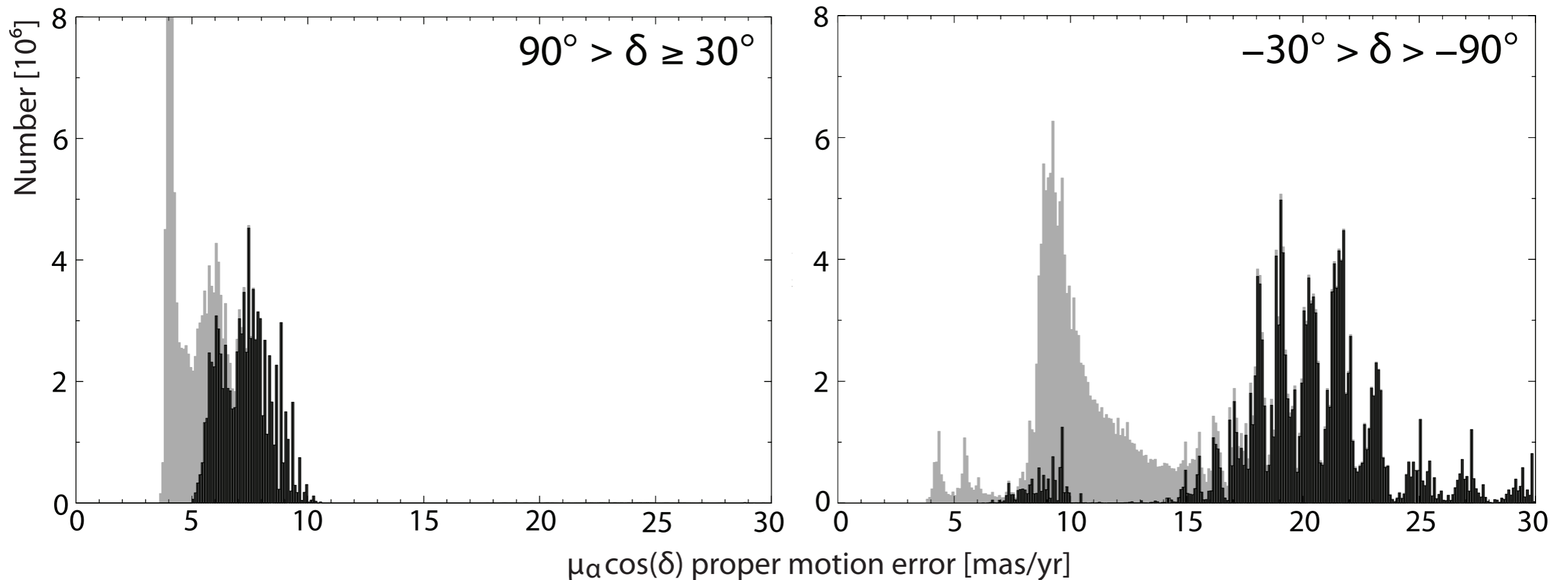
- Irregular (flares, accretion) and periodic (spots)
- Hour-week timescales - cadence problematic?



Orion OB1 (Briceño+05)

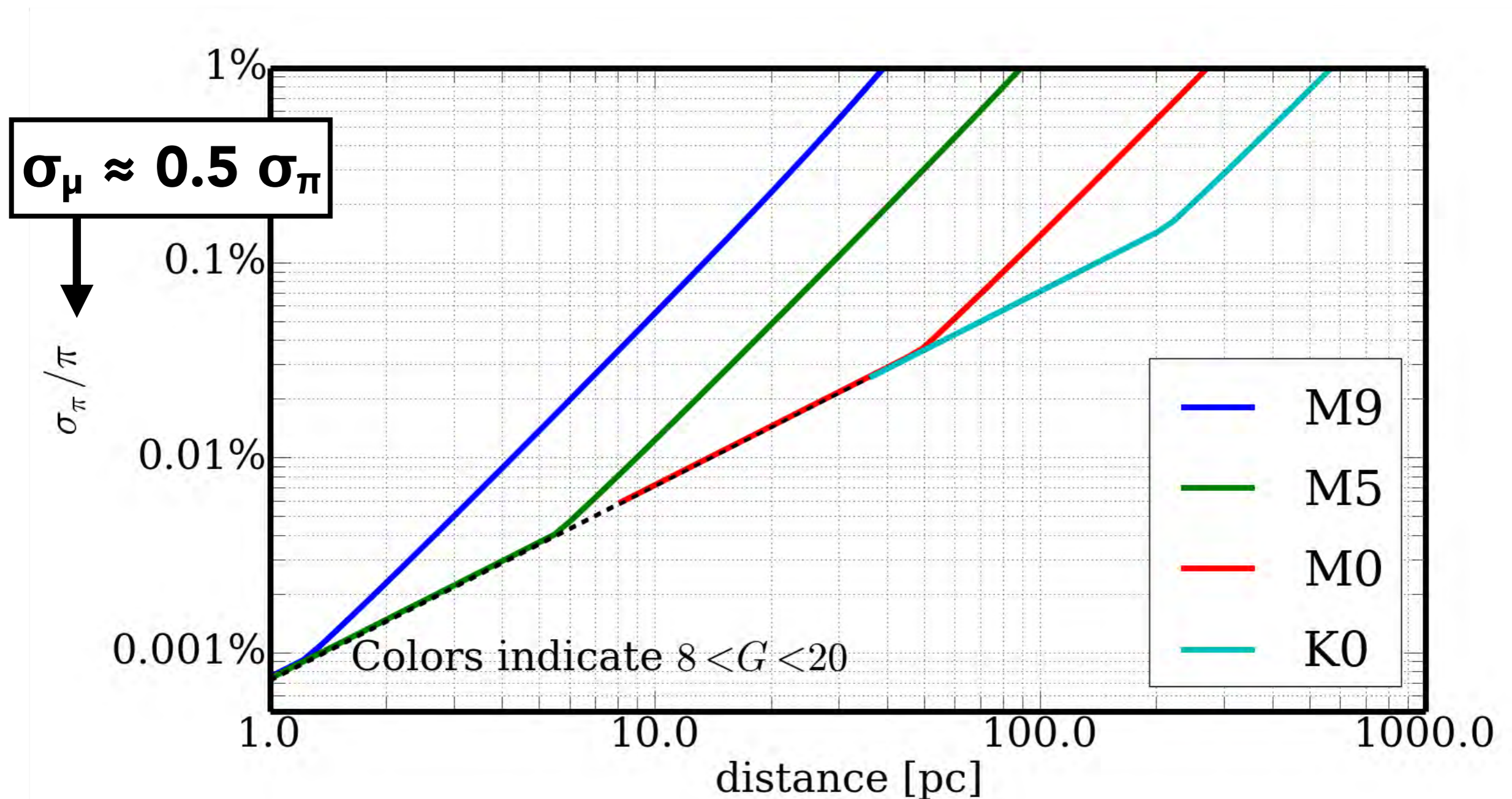
ASTROMETRY

- Much room for improvement below $\delta < -30^\circ$
- Original plan: 36 epochs/5 years \rightarrow PMs to 2-4 mas/yr



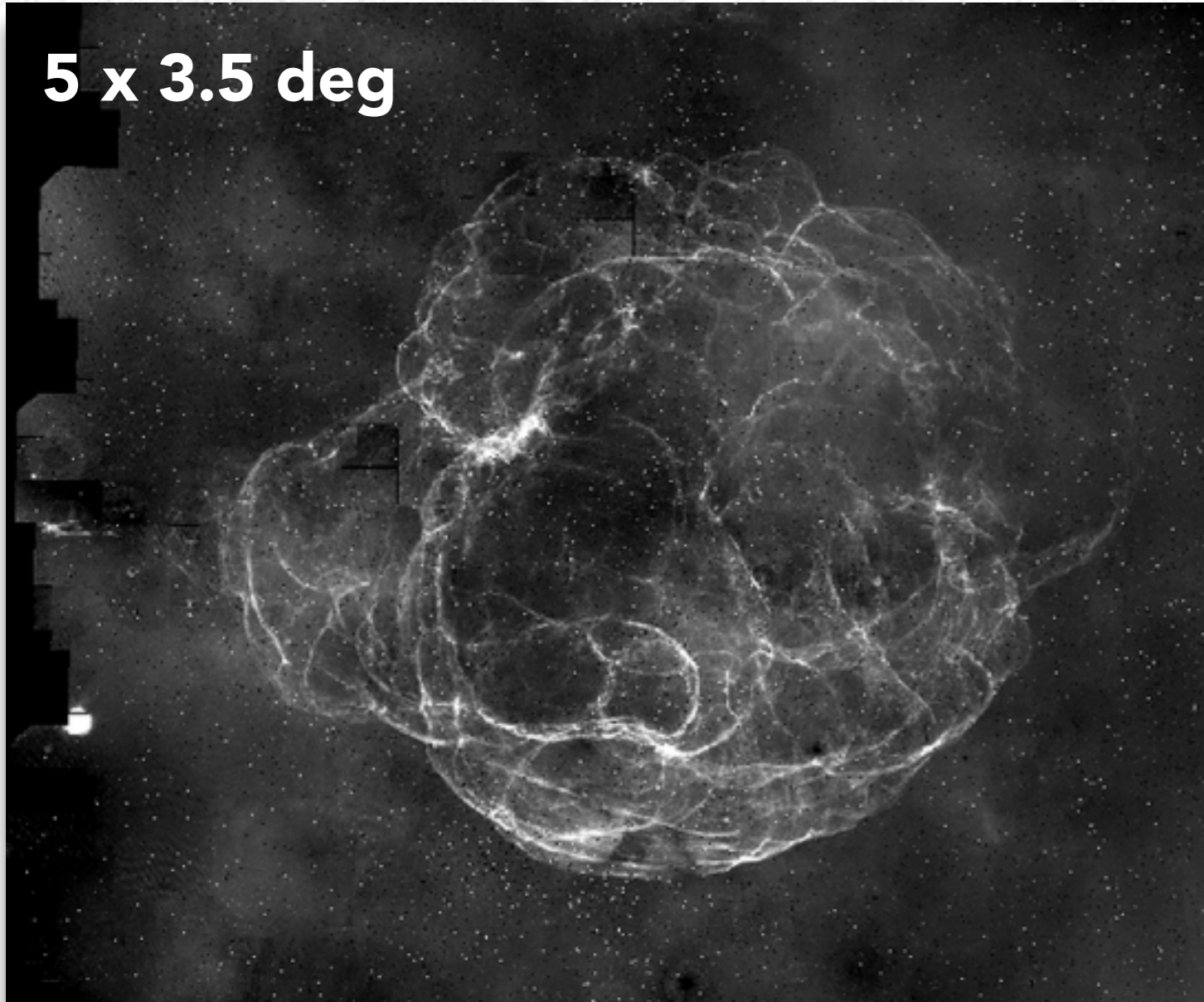
ASTROMETRY

- ...but why compete with *Gaia*?



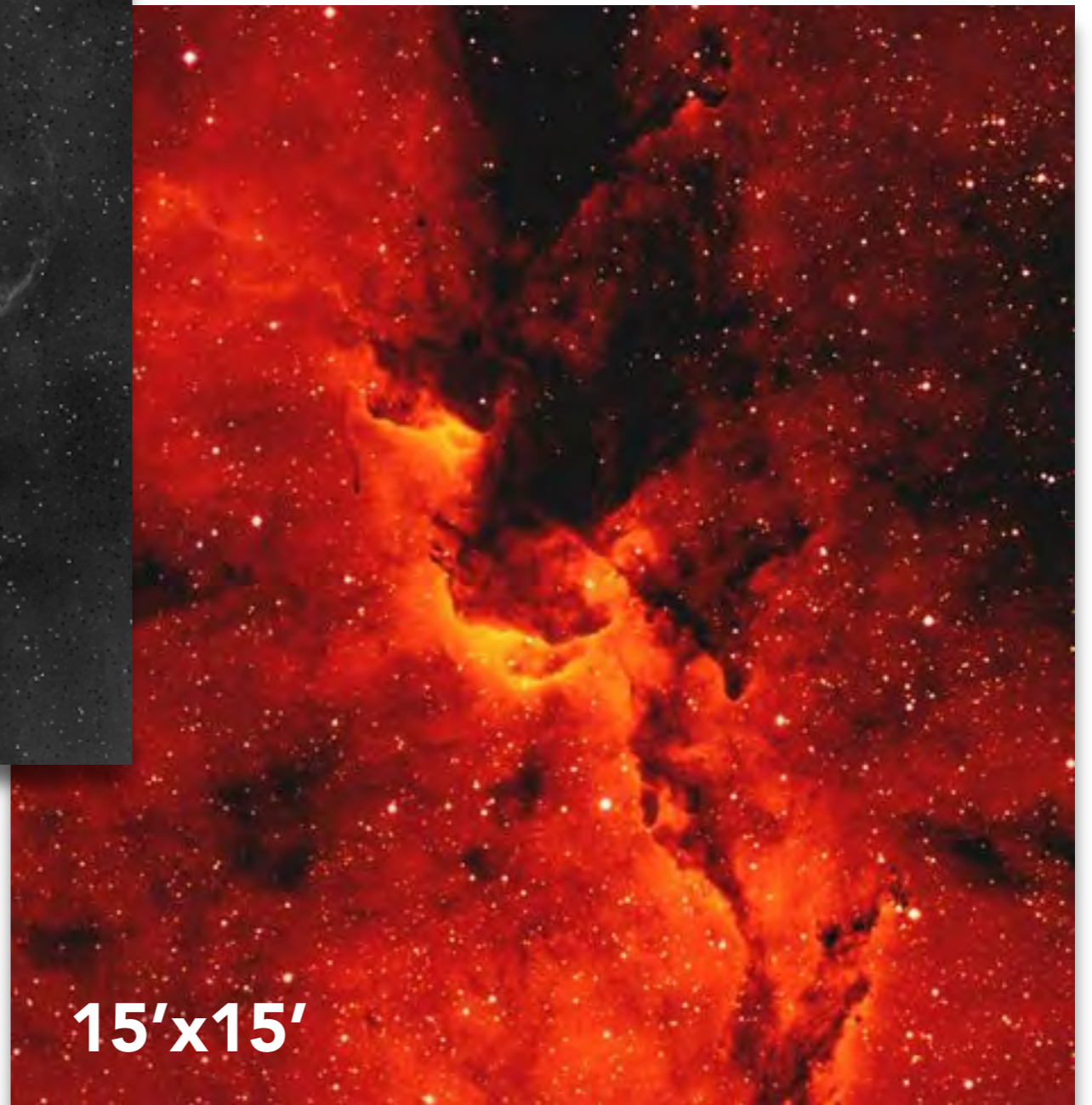
H-ALPHA PHOTOMETRY

5 x 3.5 deg



SNe remnant S147 (A. Zijlstra & J. Irwin)

IC 1396B (Nick Wright)



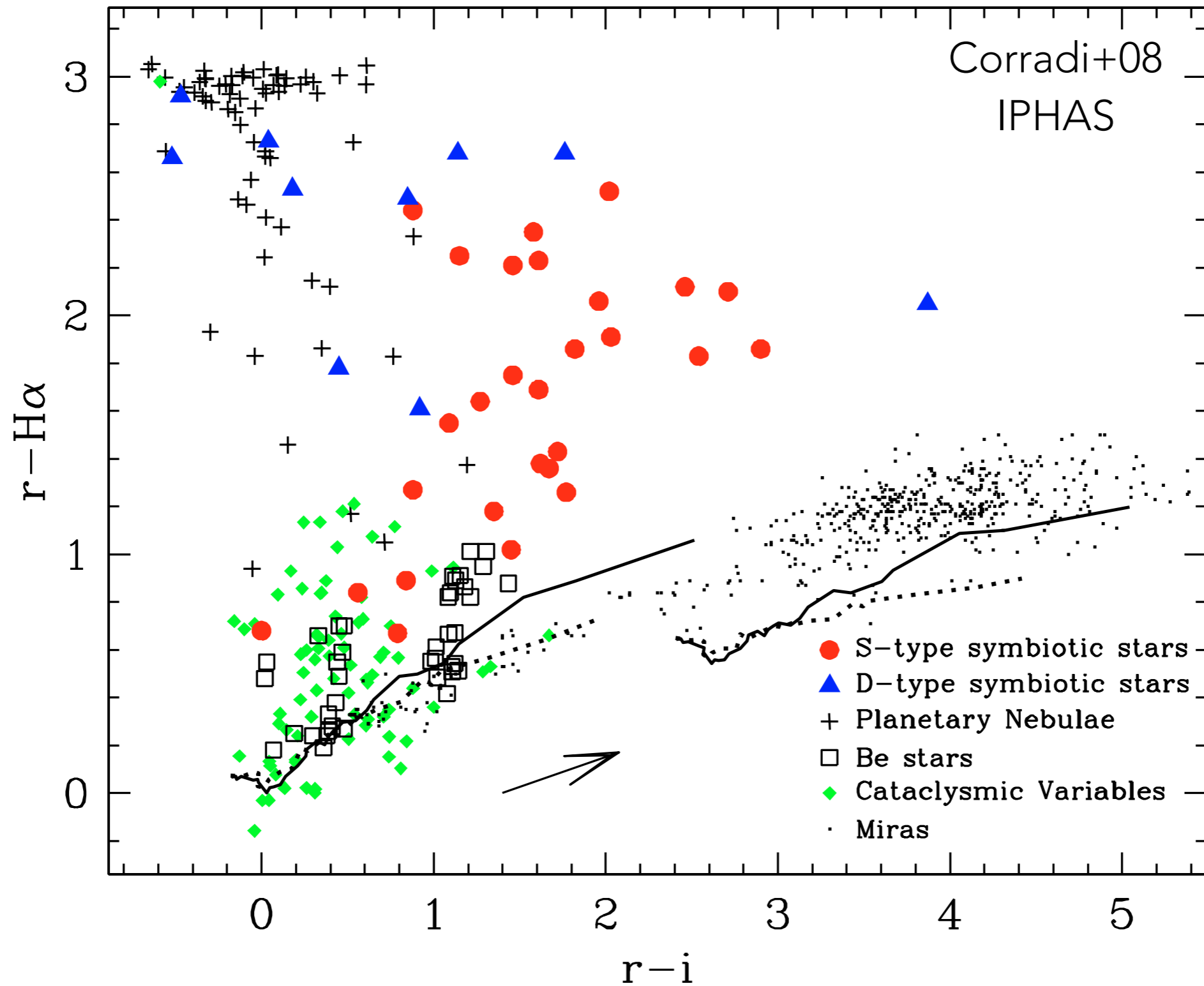
15'x15'



H-ALPHA PHOTOMETRY

- H α emission common in pre- and post-MS evolution:
 - **Accreting young stars (particularly M dwarfs)**
 - Unresolved PNe; post-AGB stars
 - Be stars (incl. **young Herbig Be**)
 - Interacting binaries (symbiotic, accreting systems)
 - H-rich white dwarfs
 - Near-MS A stars and M giants
- **Extragalactic studies (SFR etc...)**

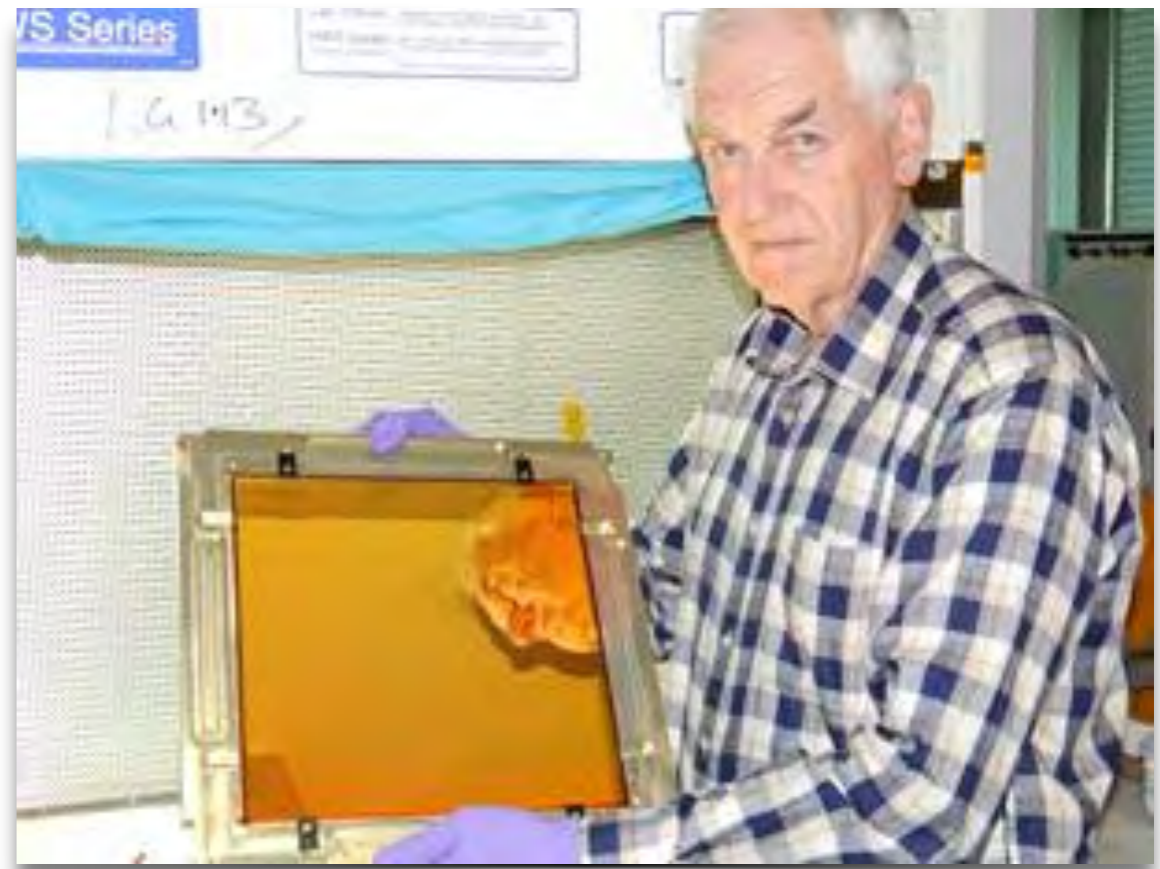
H-ALPHA PHOTOMETRY



H-ALPHA PHOTOMETRY

- 658 nm; 12 nm FWHM H α filter from Materion (Barr)
- 93% transmission
- Worlds largest (309x309 mm) and most uniform
- Will be swapped with u filter in non-survey time

The happy father after a long and painful delivery



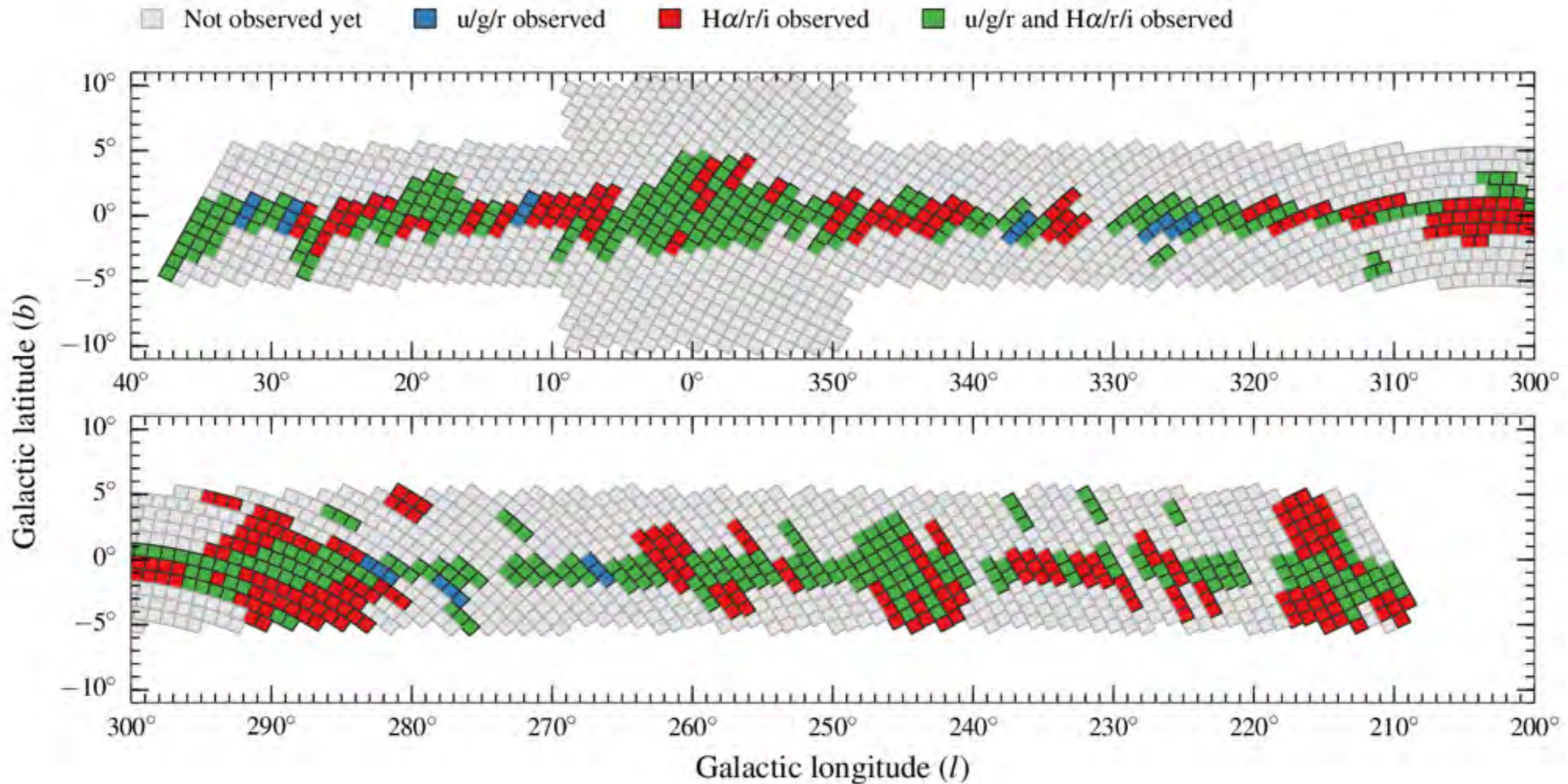
H-ALPHA PHOTOMETRY

- UKST SuperCosmos H α Survey (SHS, Parker+05)
 - 4000 deg² along southern Galactic plane, photographic; $R=20.5$
- VST Photometric H α Survey (VPHAS+, Drew+14)
 - 2000 deg² along northern Galactic plane ($b \pm 5^\circ$) + bulge
 - VST/OmegaCam; $ugri+H\alpha$ to 20th mag



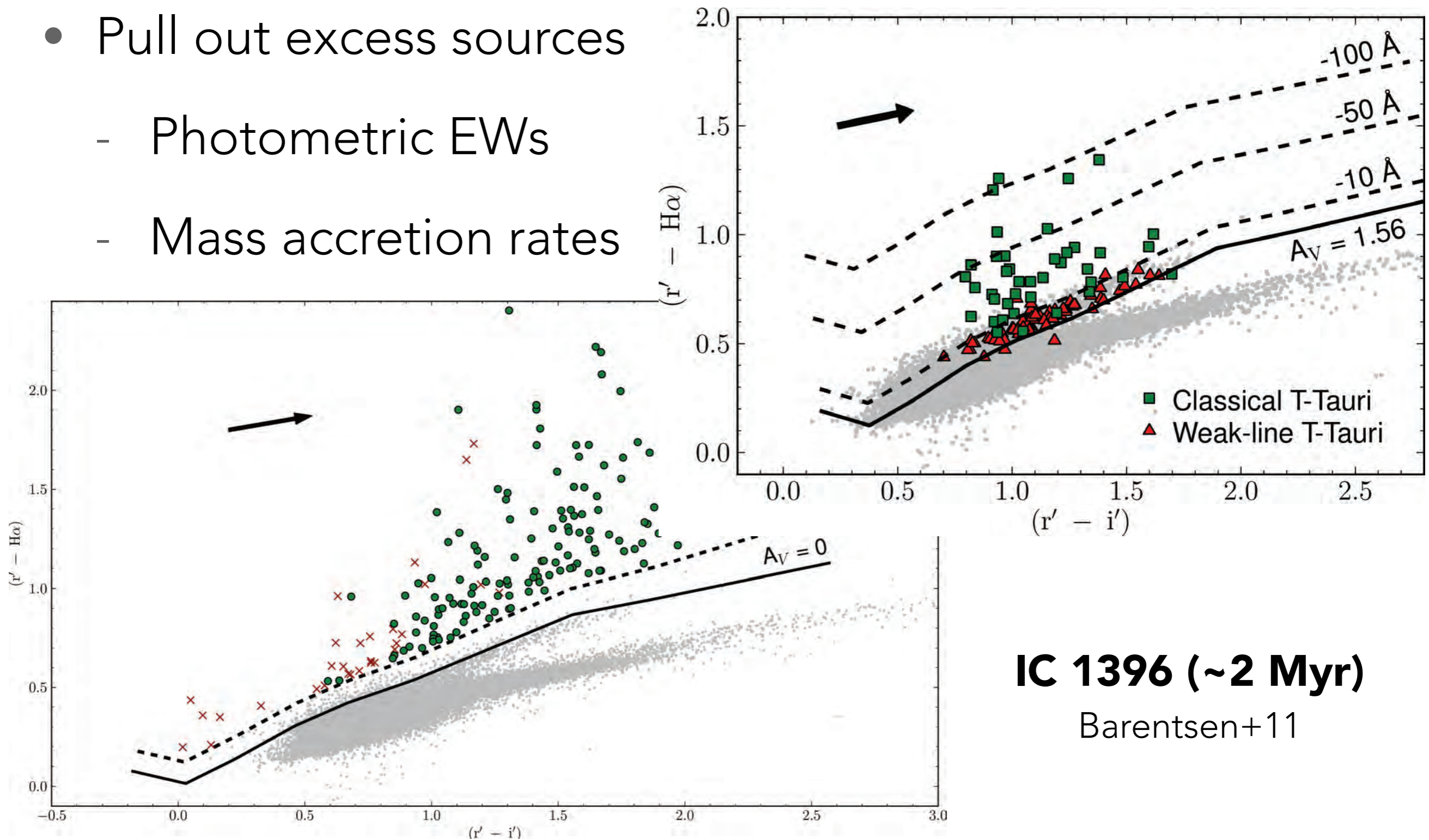
H-ALPHA PHOTOMETRY

VPHAS+ at January 2014



H-ALPHA PHOTOMETRY

- Pull out excess sources
 - Photometric EWs
 - Mass accretion rates

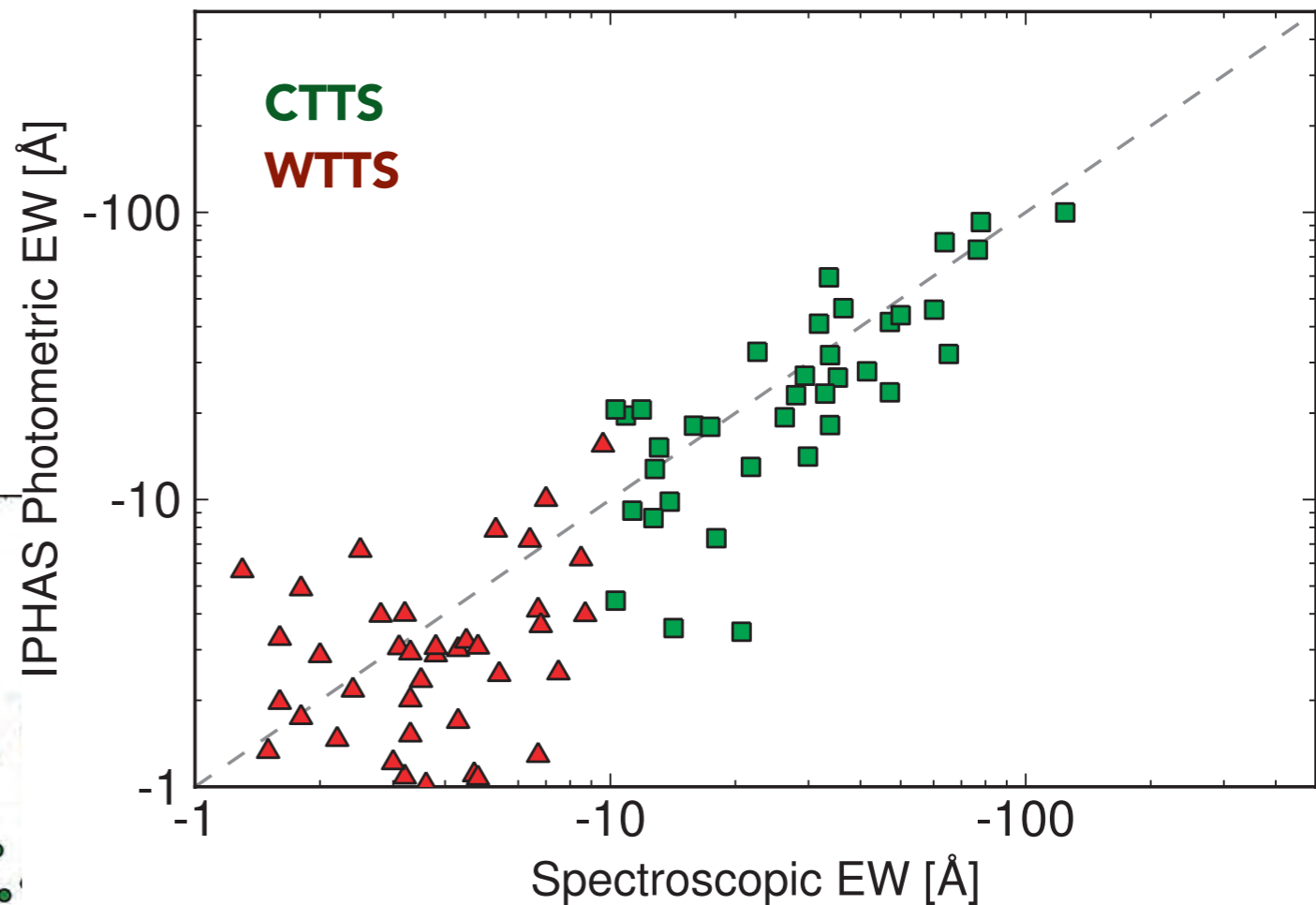
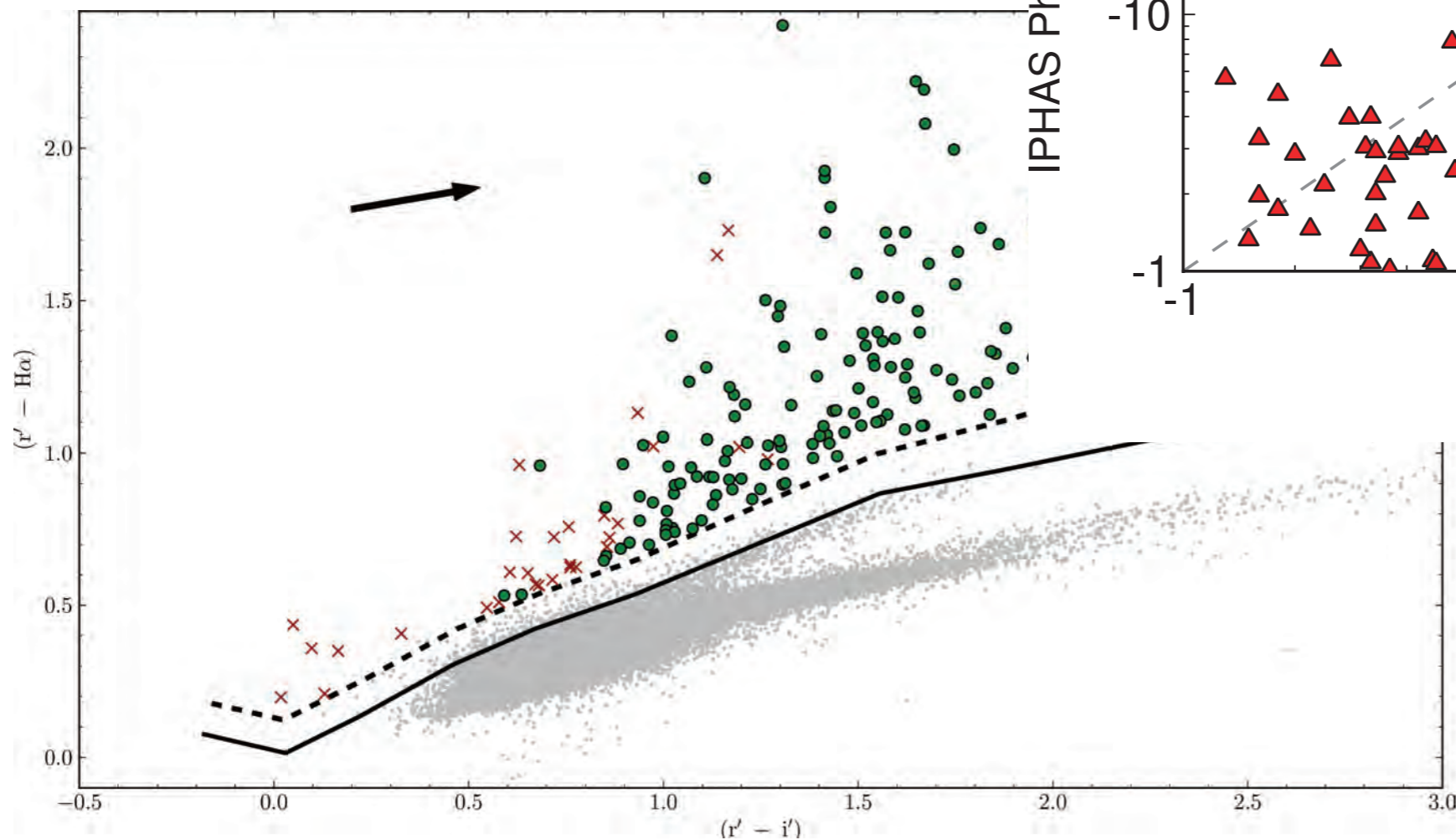


IC 1396 (~2 Myr)

Barentsen+11

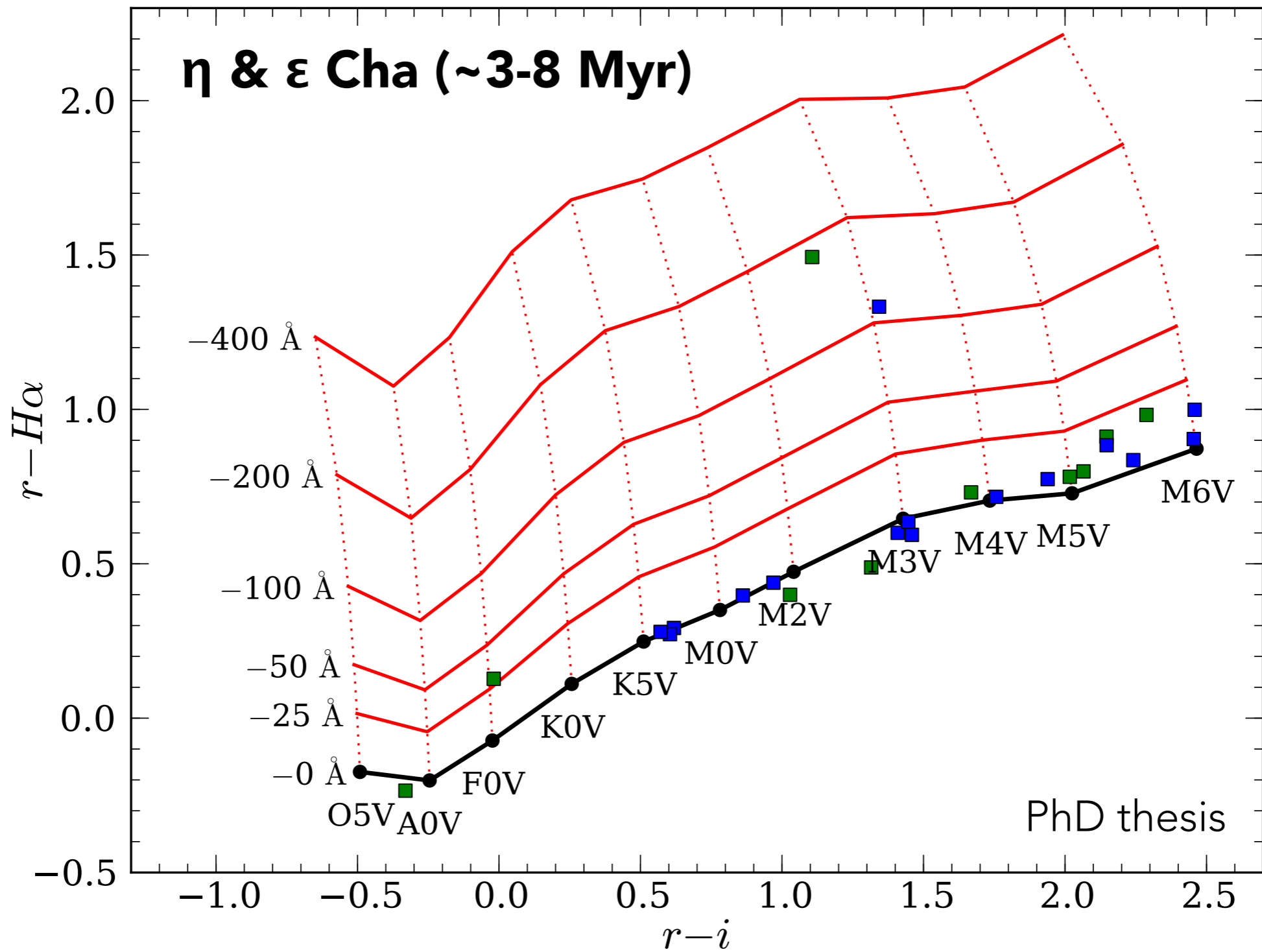
H-ALPHA PHOTOMETRY

- Pull out excess sources
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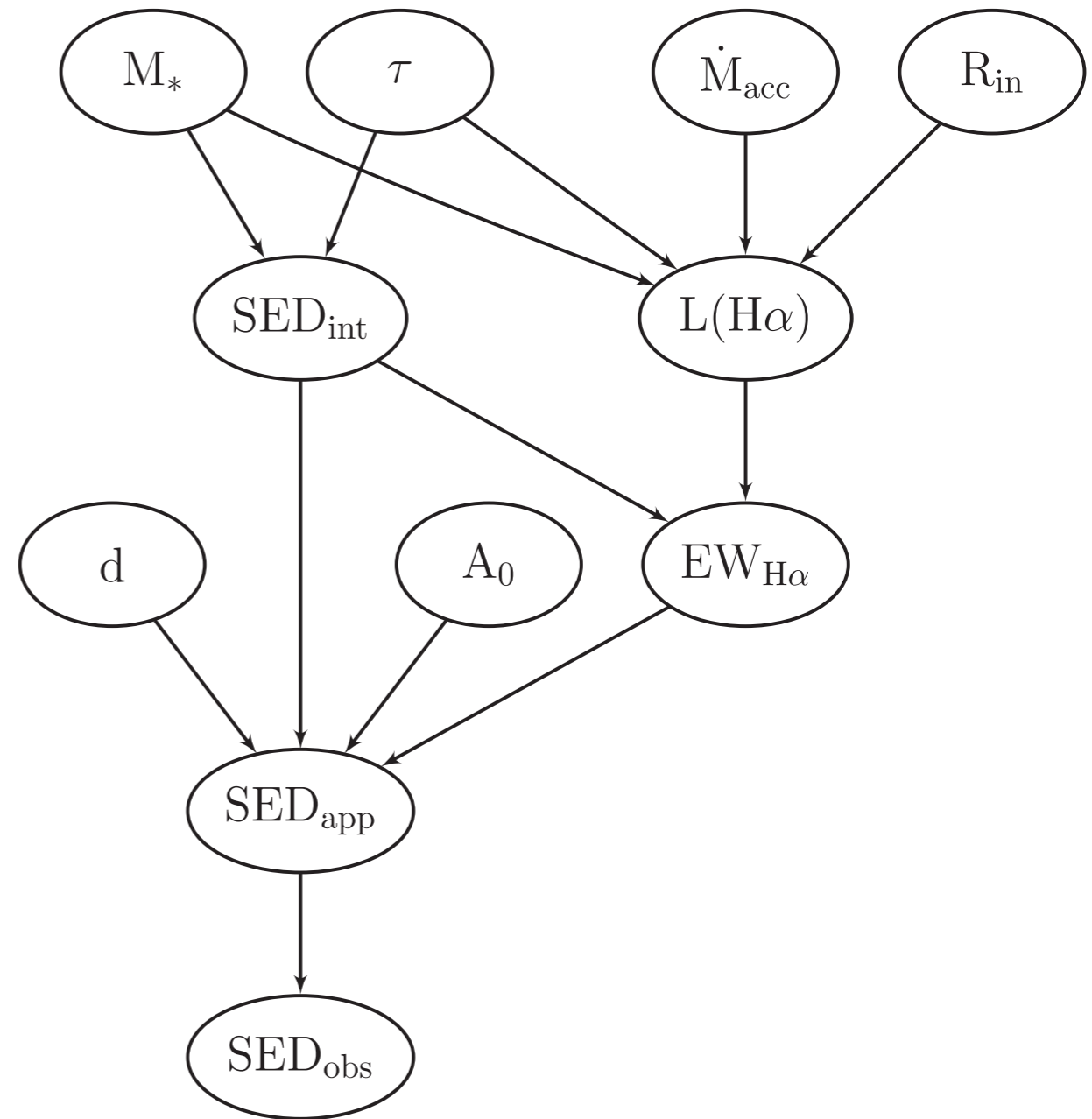
IC 1396 (~2 Myr)
Barentsen+11

H-ALPHA PHOTOMETRY



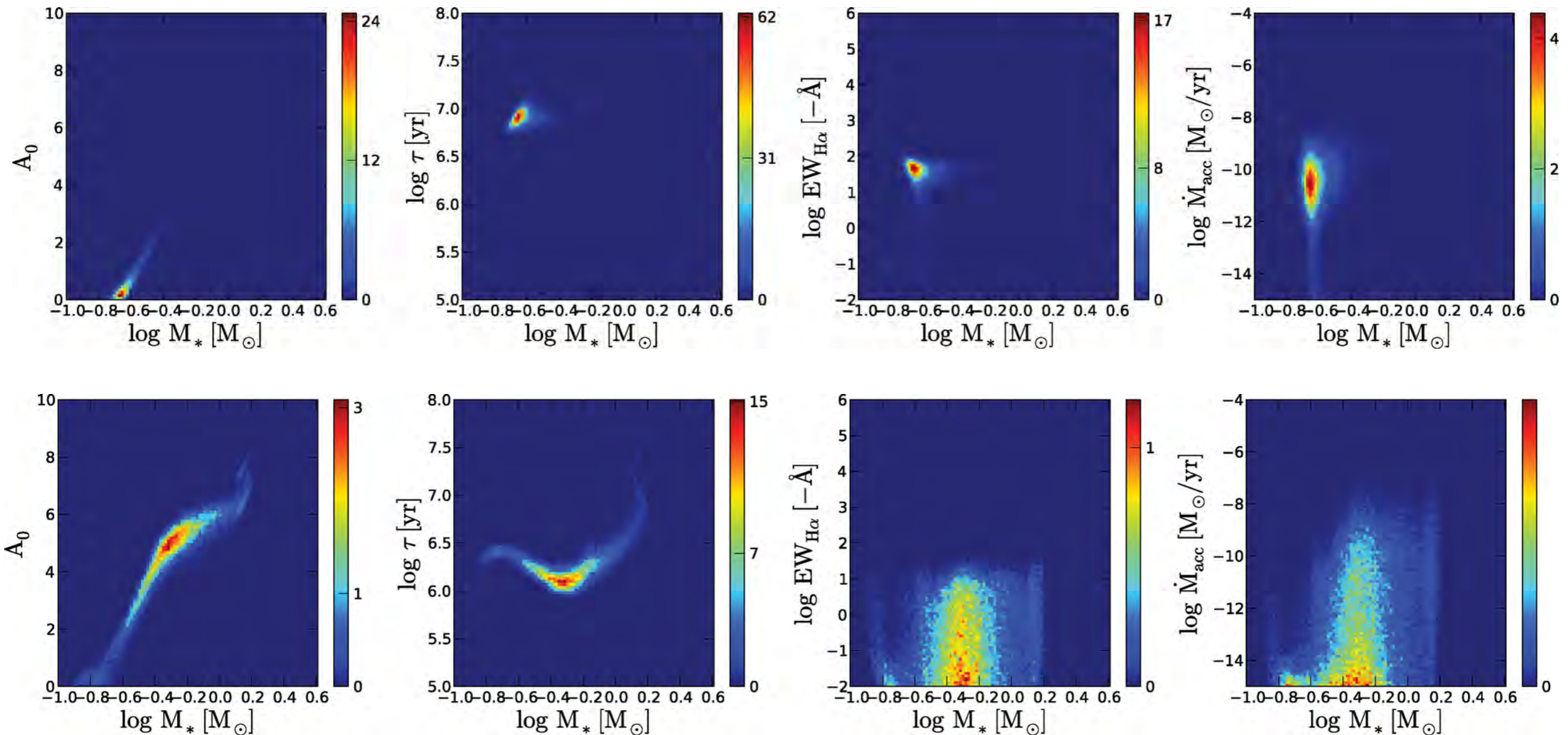
STELLAR PARAMETERS

- **Degeneracy between M , age, A_0 and H α emission**
- Bayesian network
- Priors on **$M, d, \text{age}, dM/dt$**
 R_{in} (disk truncation) & **A_0**
- Compare model ($r, i, H\alpha$) to observations
- Sample joint distribution using MCMC



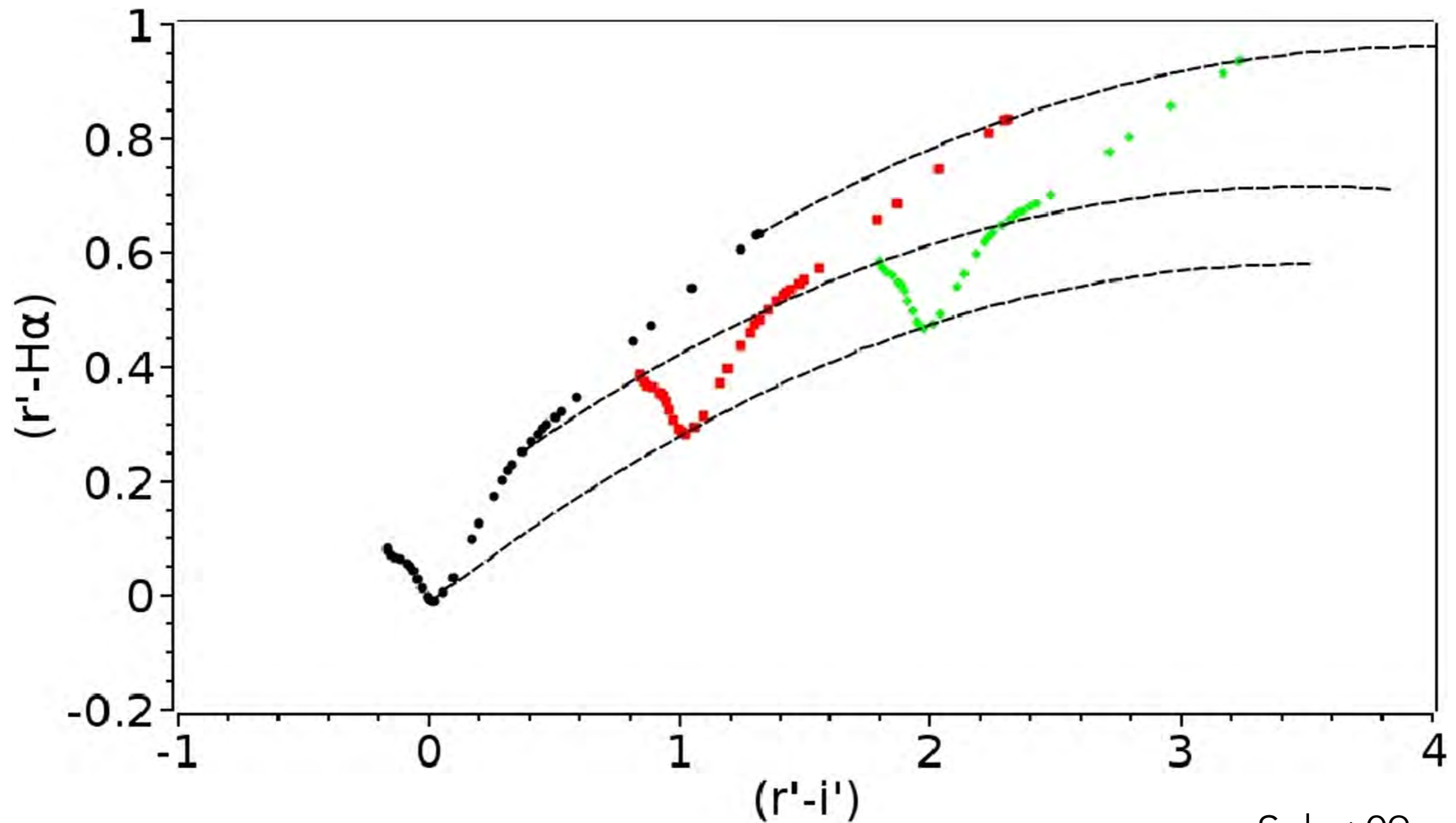
STELLAR PARAMETERS

- Able to explore degeneracies, prior dependencies



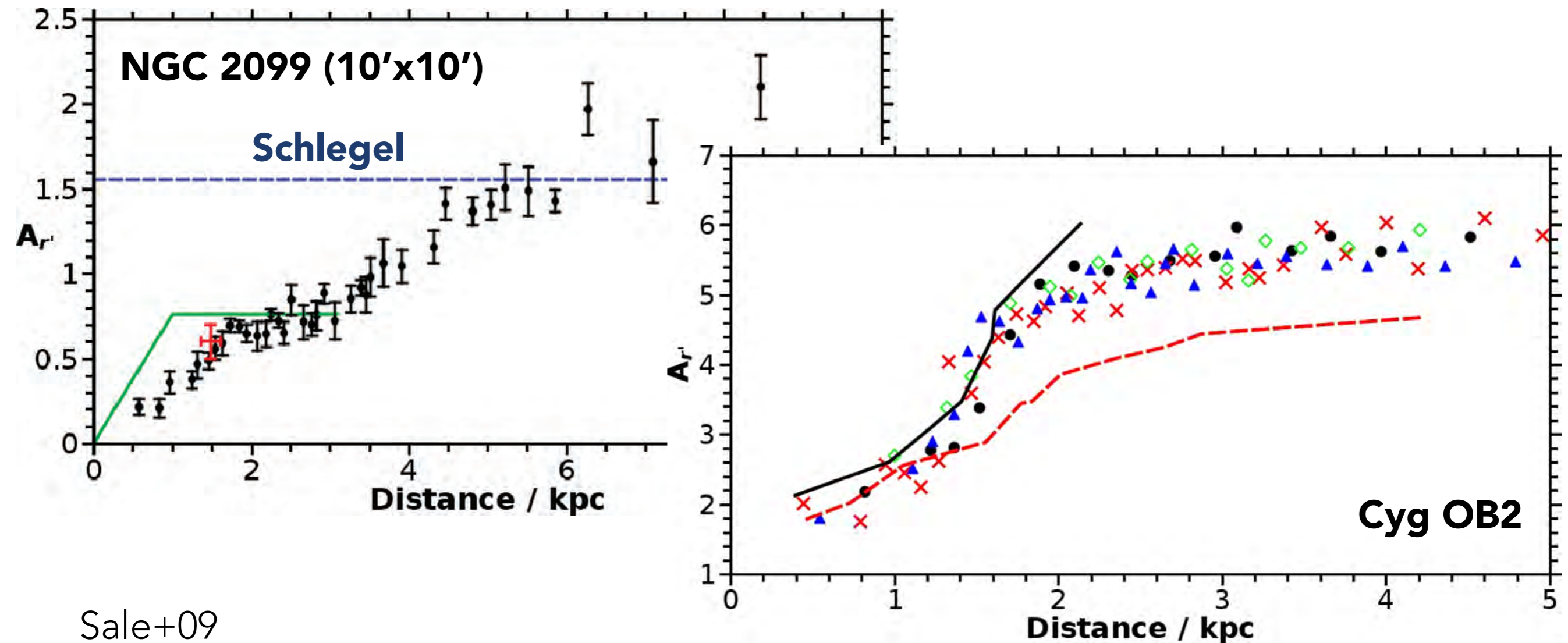
EXTINCTION MAPPING

- $(r-i, r-H\alpha)$ reddening vector at large angle to MS



EXTINCTION MAPPING

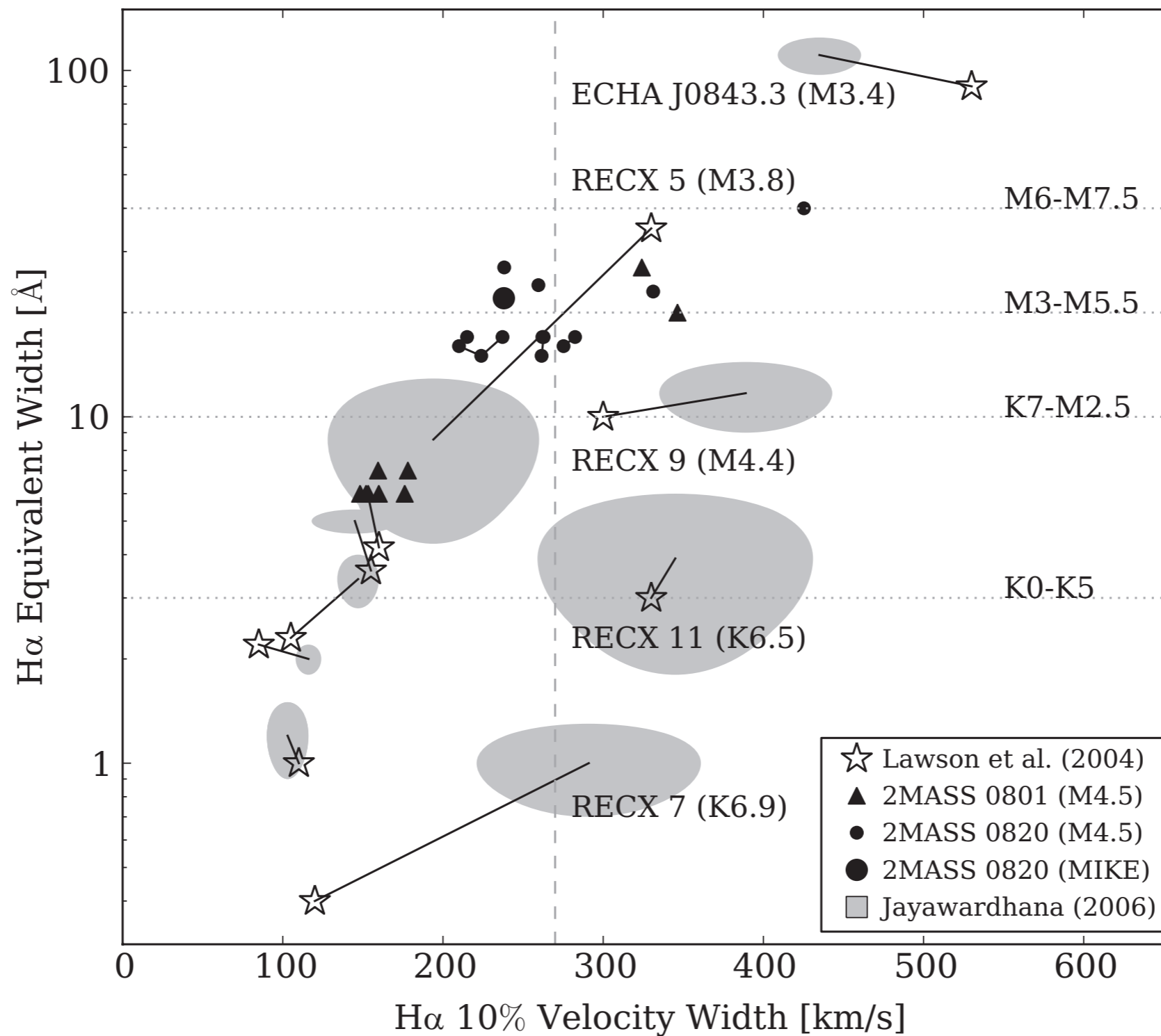
- Extract $A_{V,d}$ for early-A to K4 stars
- MEAD (Sale+09): Map $A_{r'}$ at $10'$, 0.1 kpc resolution



SURVEY STRATEGY

- Filter available in (bright) non-survey time (25%)
- **VPHAS+ bright limit is $r=12-14$**
- Cover Galactic plane to start, entire 2π deep survey?
- MC/galaxy survey
- Monitoring of southern SFR

ACCRETION VARIABILITY



Murphy+11