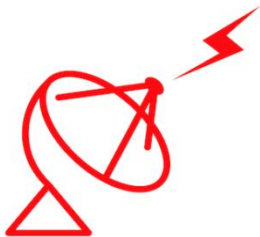




# Tracing the HI content in obscured AGN with ASKAP

**Marcin Glowacki**

**Supervisors: Elaine Sadler & James Allison**



**ASKAP-FLASH**

First Large Absorption Survey in HI

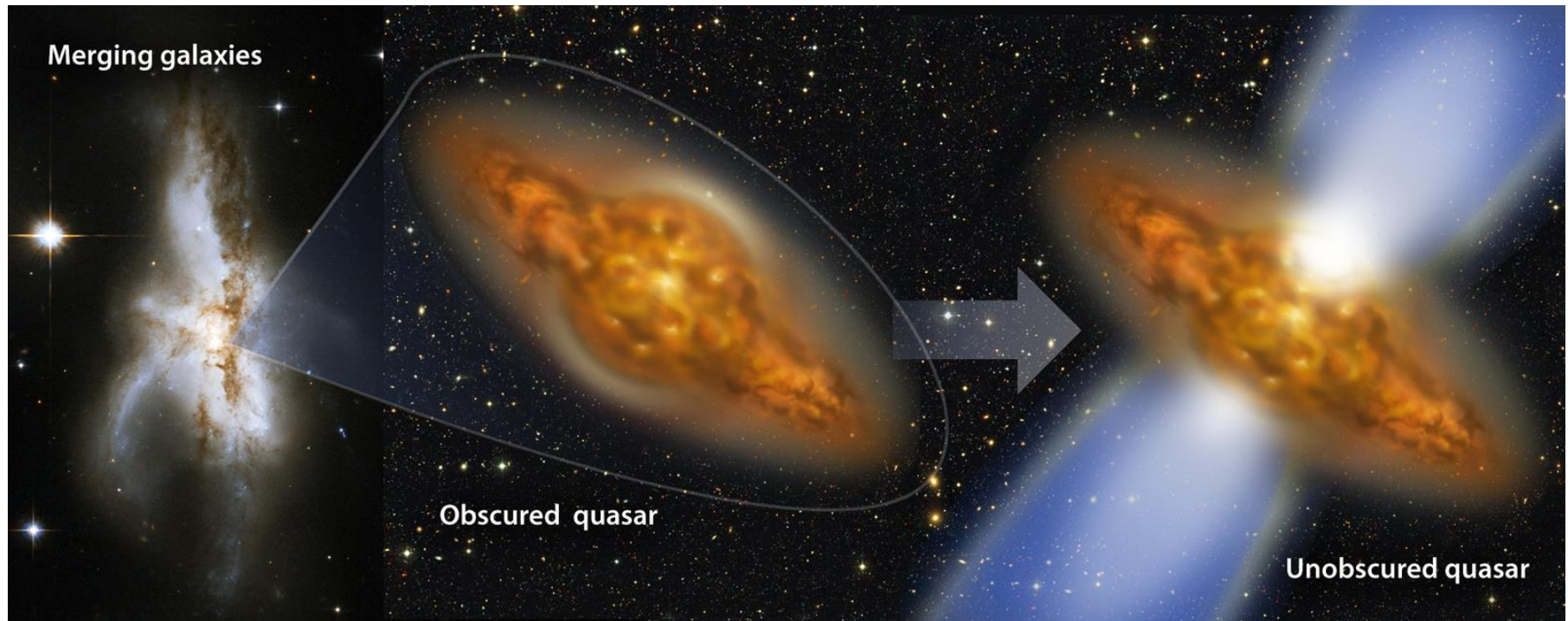




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# Obscured Quasars

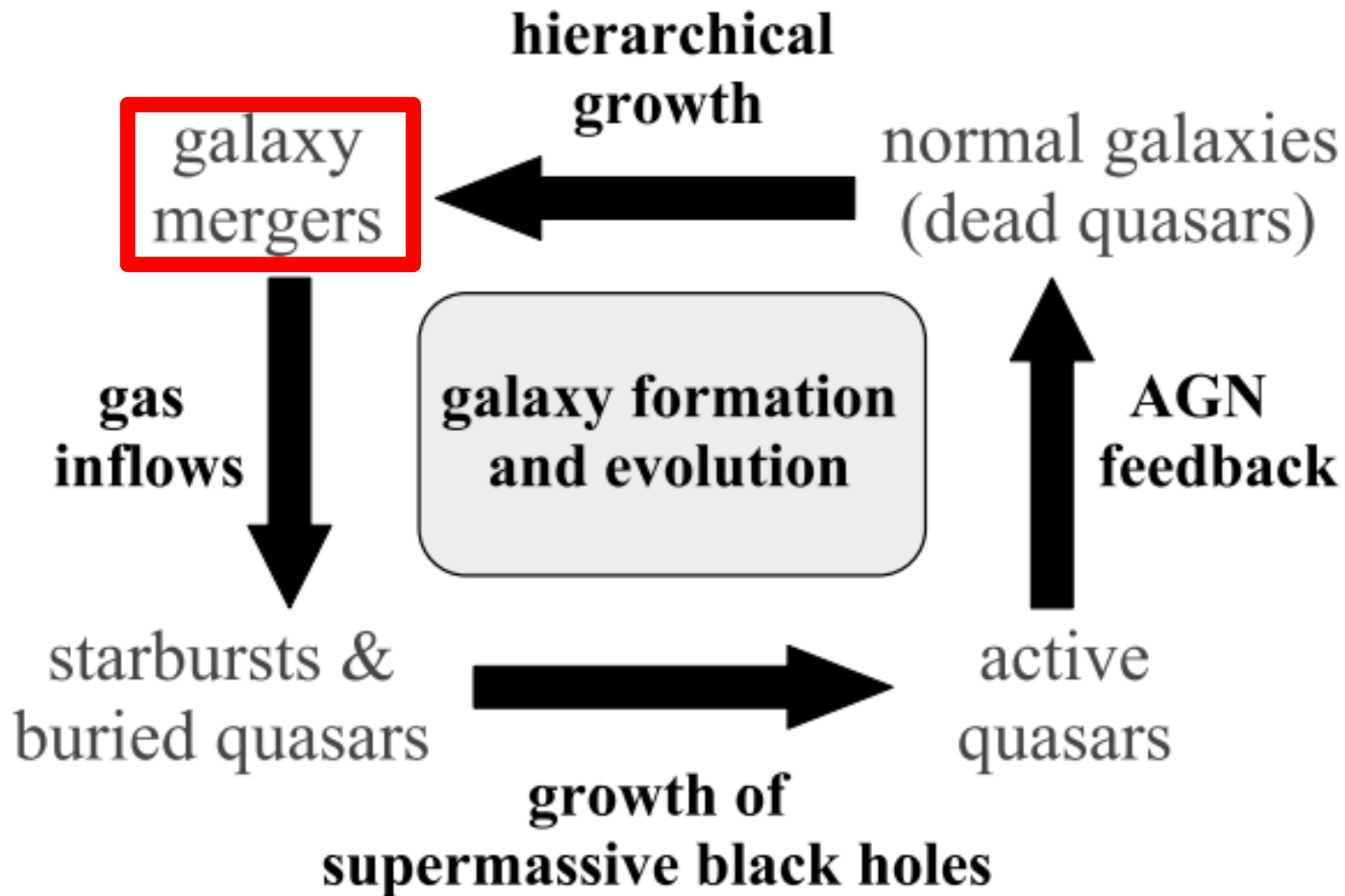
- Optically obscured quasars represent a transitional period between:
  - Merger event  $\rightarrow$  gas + dust fuels the AGN
  - AGN switches on, feedback disperses material



*Artist's impression, by Karen Teramura*



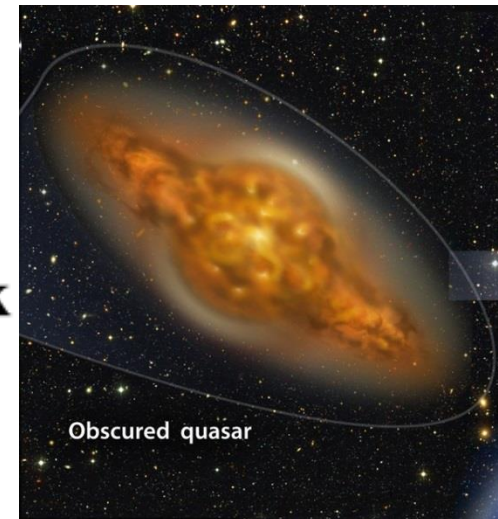
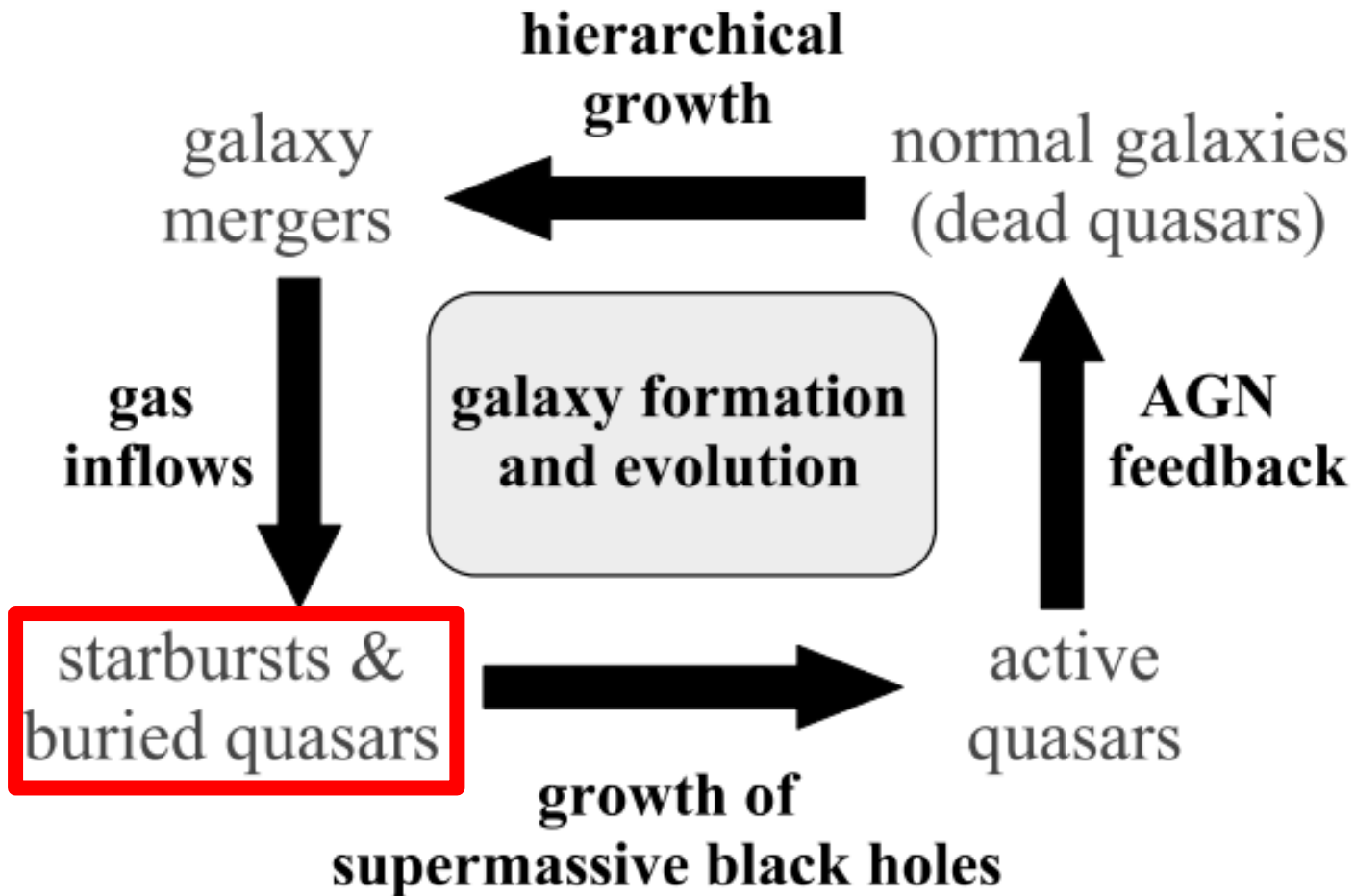
# Obscured Quasars





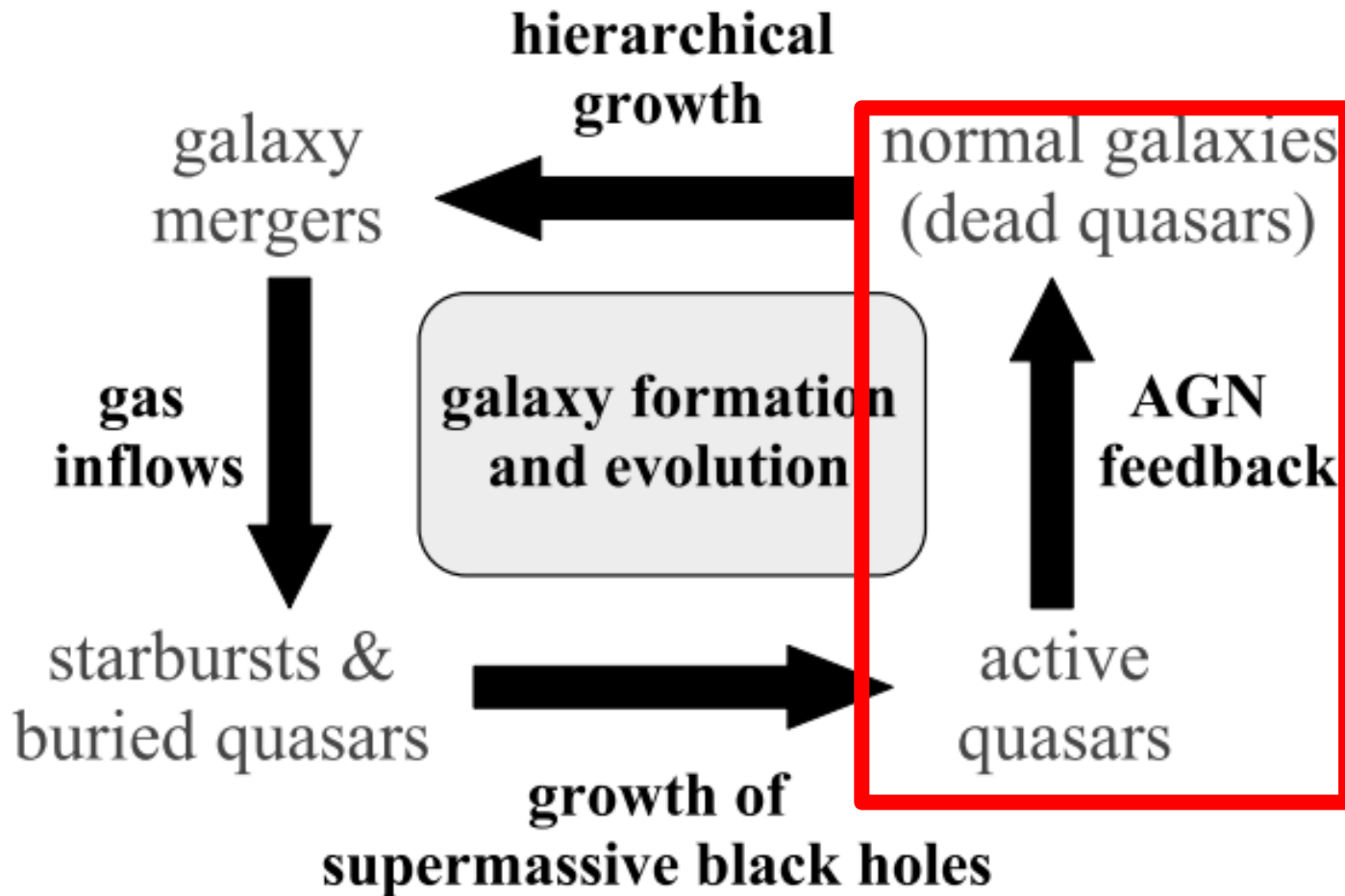


# Obscured Quasars





# Obscured Quasars

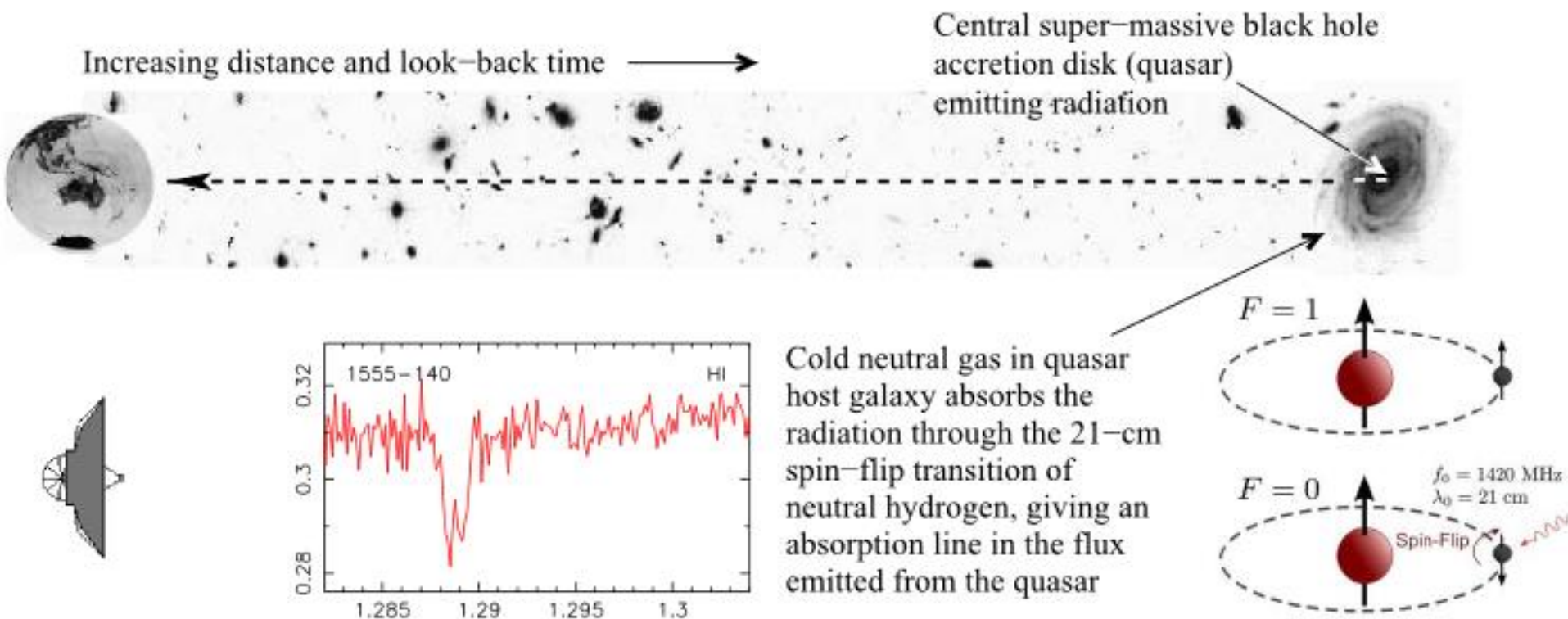




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# HI 21 cm in Absorption

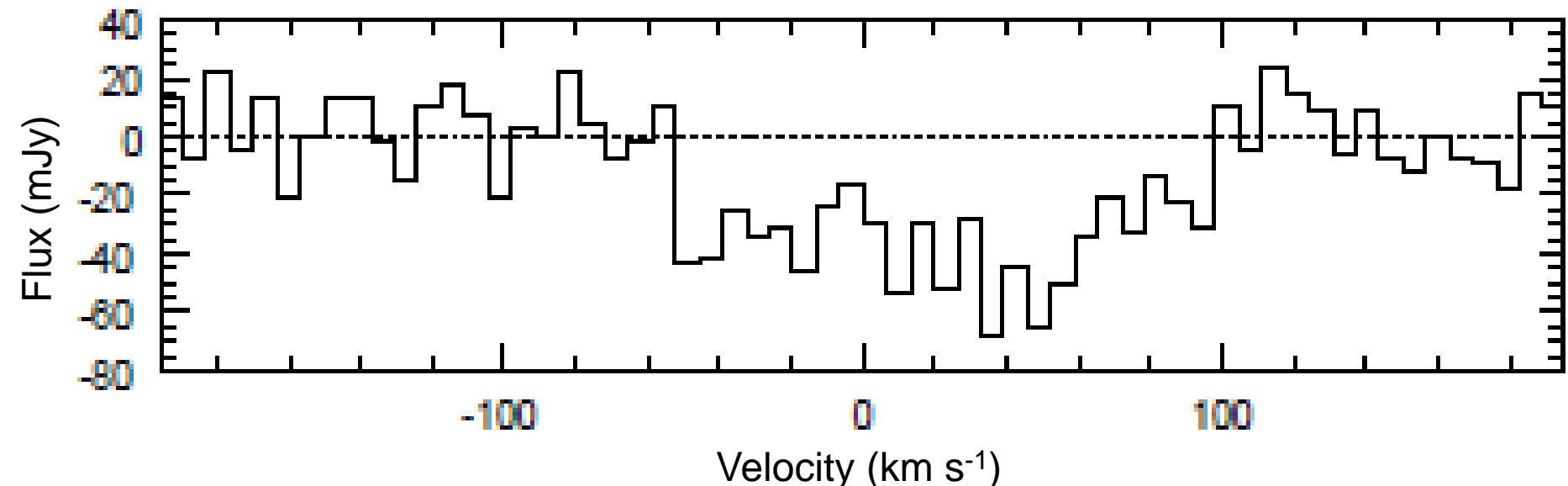
- Can trace gas within host galaxy or along line-of-sight of a radio-bright source (e.g. AGN)
- No redshift limit for transition in absorption





# How much HI content?

- Carilli et al. (1998) found an **80% HI absorption detection rate** (associated and intervening)
- **Only 5 objects** searched to low optical depths
- Need to study a larger sample to learn about gas kinematics near quasars in obscured phase



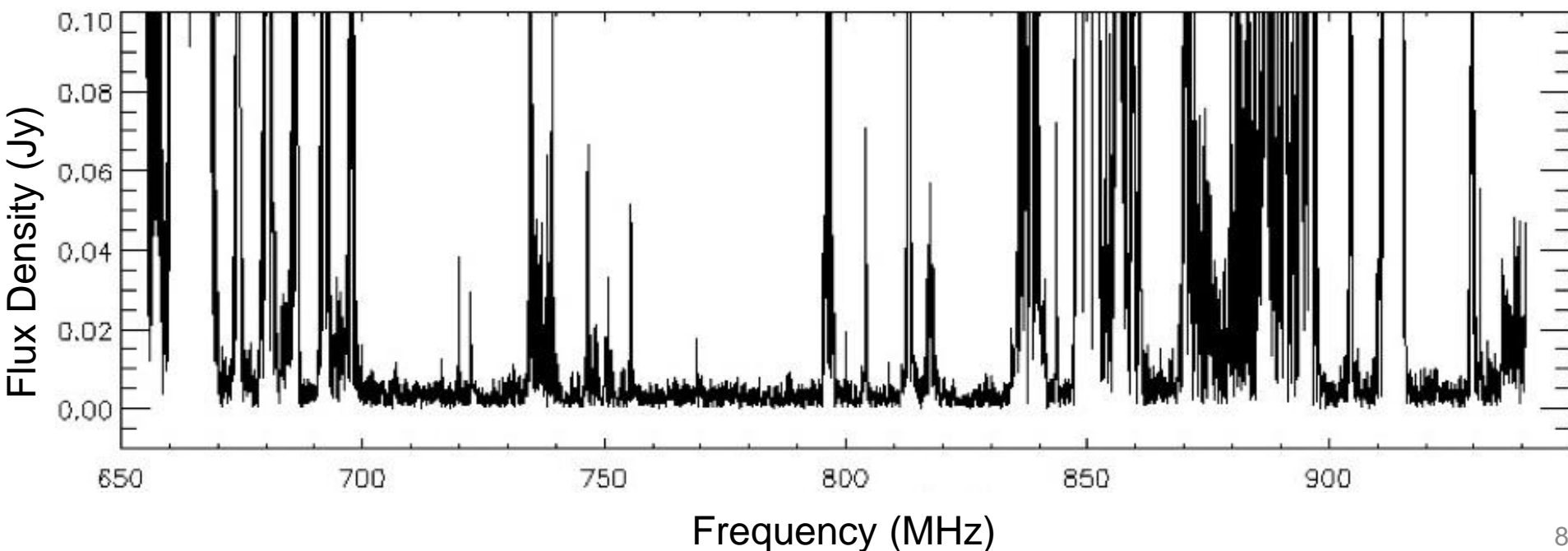


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# RFI

- Radio Frequency Interference (RFI) is problematic for radio surveys, particularly spectral line surveys
- Need HI absorption feature to fall in an uncorrupted location of the spectrum to be able to see it

Green Bank Telescope spectrum, 650 – 950 MHz

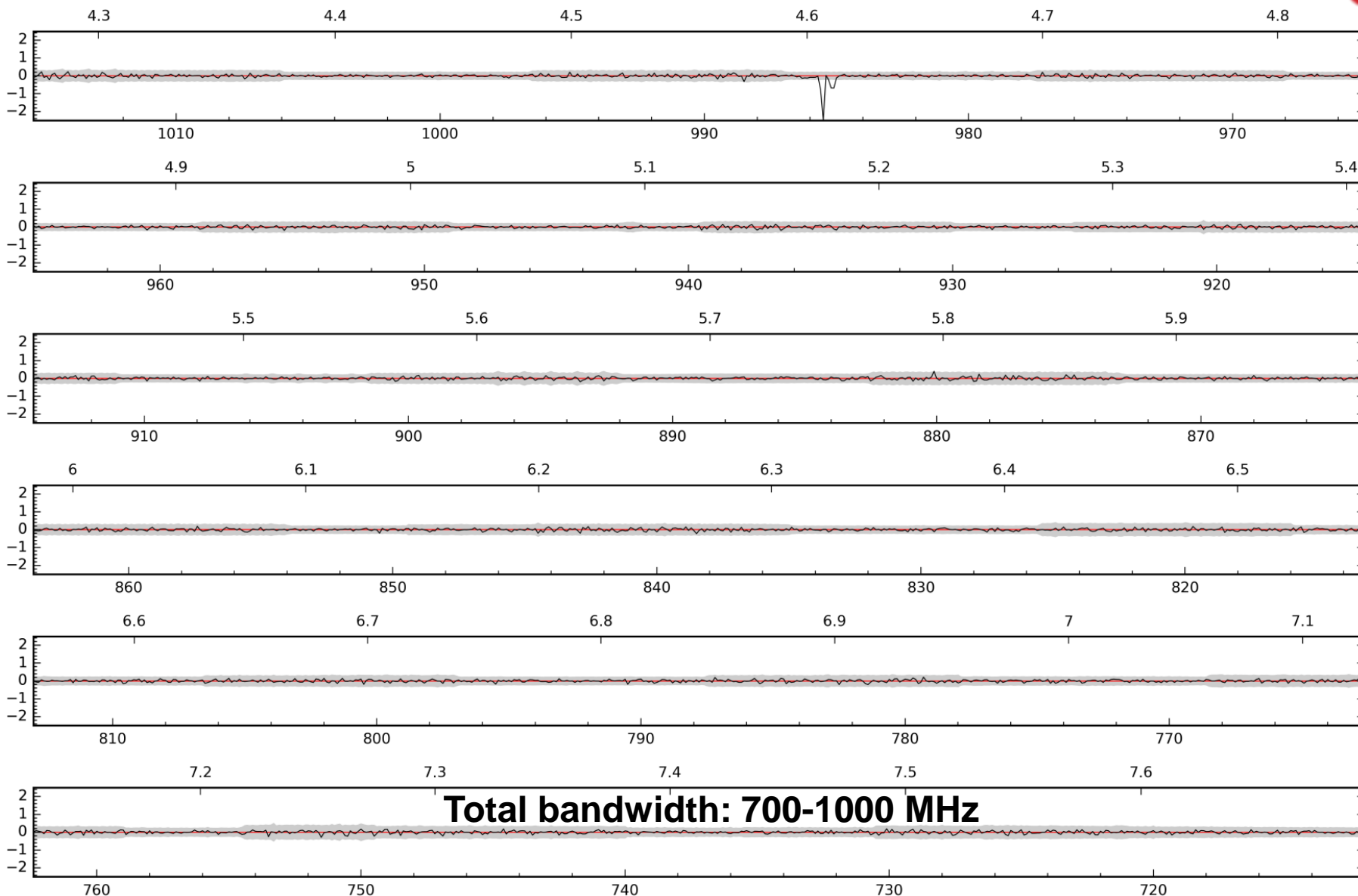






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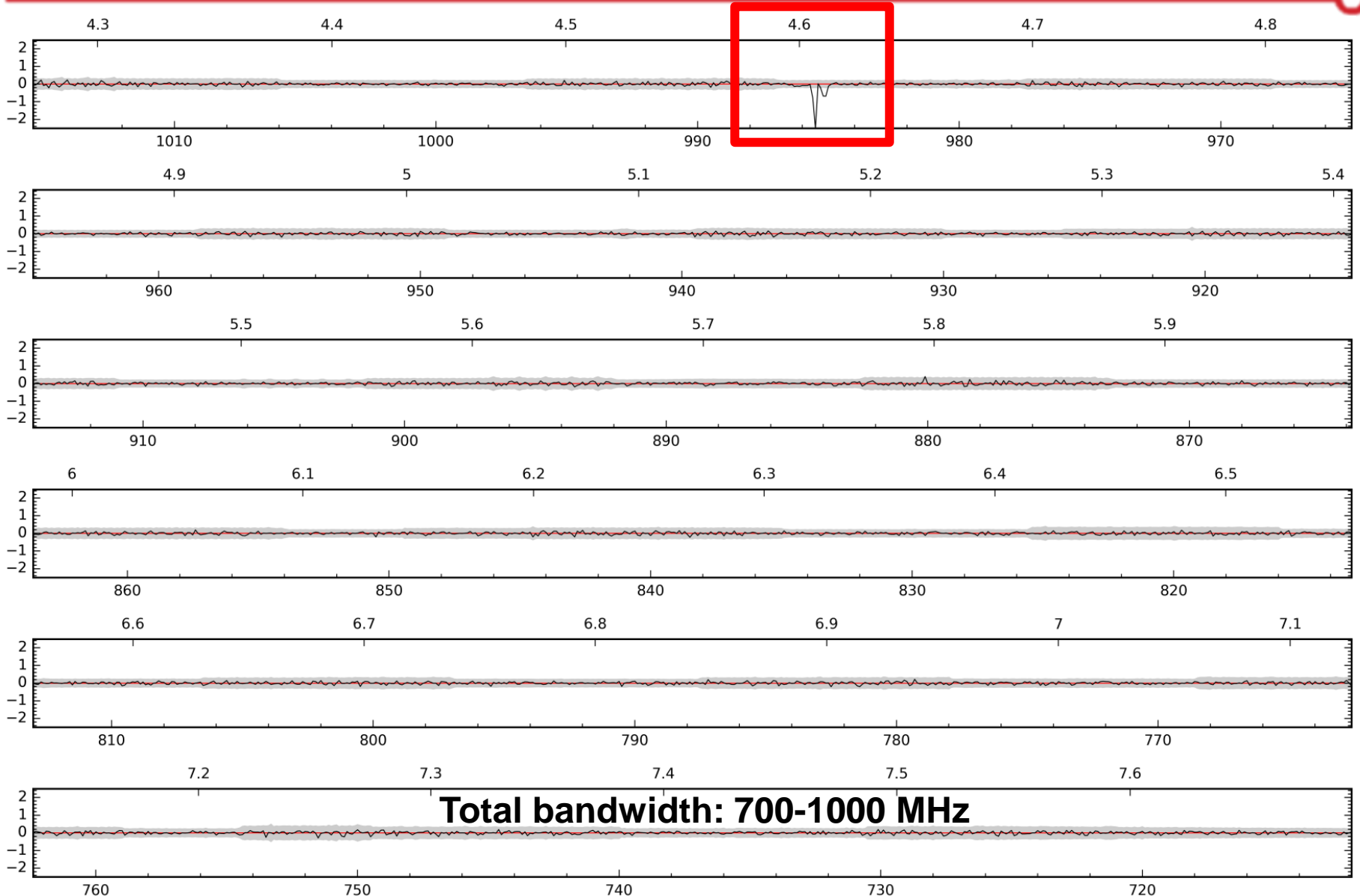
# “Bam, and the RFI is gone”





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# “Bam, and the RFI is gone”



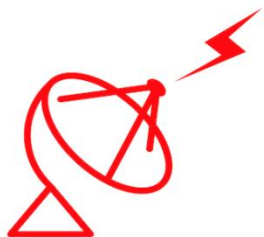


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# ASKAP and FLASH



- Array of 36 antennas (currently has ~12)
- Uses PAFs (Phased Array Feeds)
- At **radio-quiet MRO site**



**ASKAP-FLASH**

First Large Absorption Survey in HI

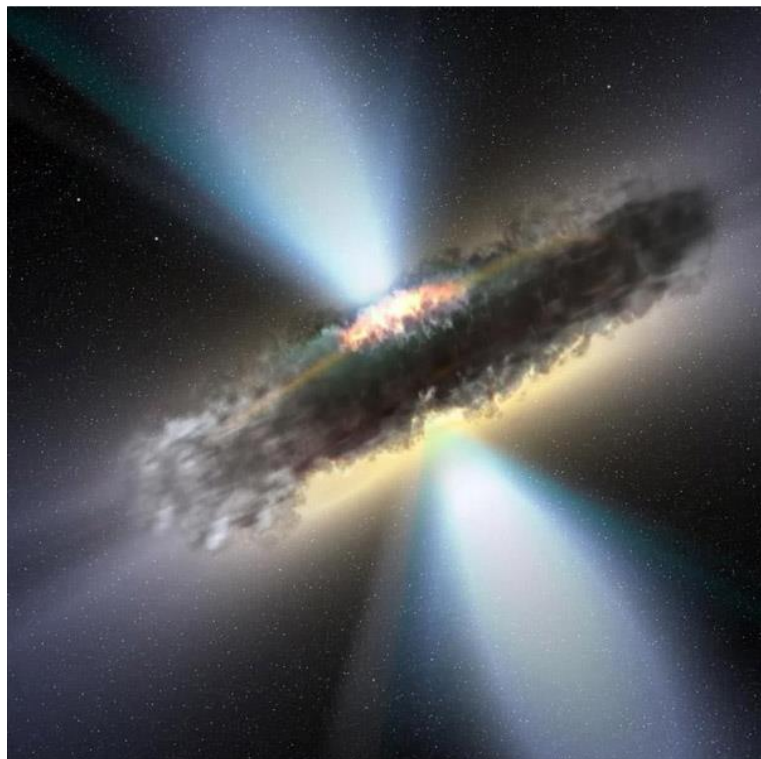
- **FLASH: All-sky HI search at redshifts  $0.4 < z < 1$**
- **Estimated ~1000 new detections against 150,000 sightlines**



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# HI survey of obscured AGN

- Aim – to further explore Carilli et al. (1998) work with current telescopes in radio-quiet site
- Used ASKAP-BETA (6 antennas) to search for HI absorption toward obscured quasars



*Credit: ESA / V. Beckmann (NASA)*

- Considerations:
  - ASKAP-BETA a prototype telescope; not as sensitive as other radio telescopes
    - Radio-loud sources selected
  - Replicate sample selection
  - Explore mid-infrared data now available

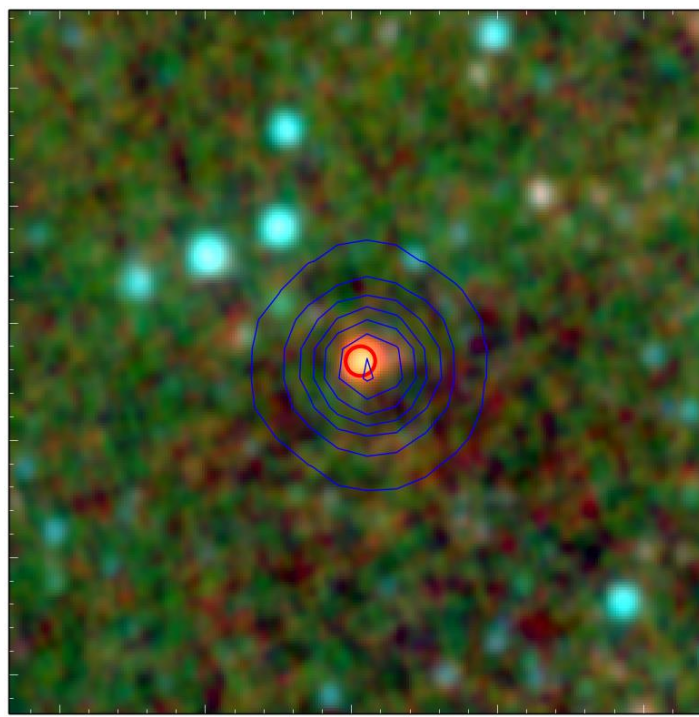


# Sample Selection

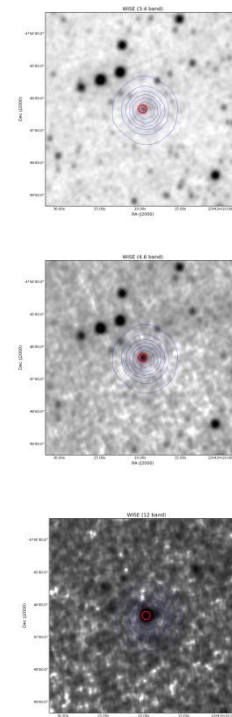
- Starting point: **Kuhr et al. (1981)** radio catalogue
- Brightness, declination and redshift cuts
- **Two subgroups: Optically faint** (R-band magnitude  $> 20$  mag) and **mid-infrared colour selected** (WISE)



*Wide-Field Infrared Survey Explorer*



*WISE 3-colour image of 3C275*



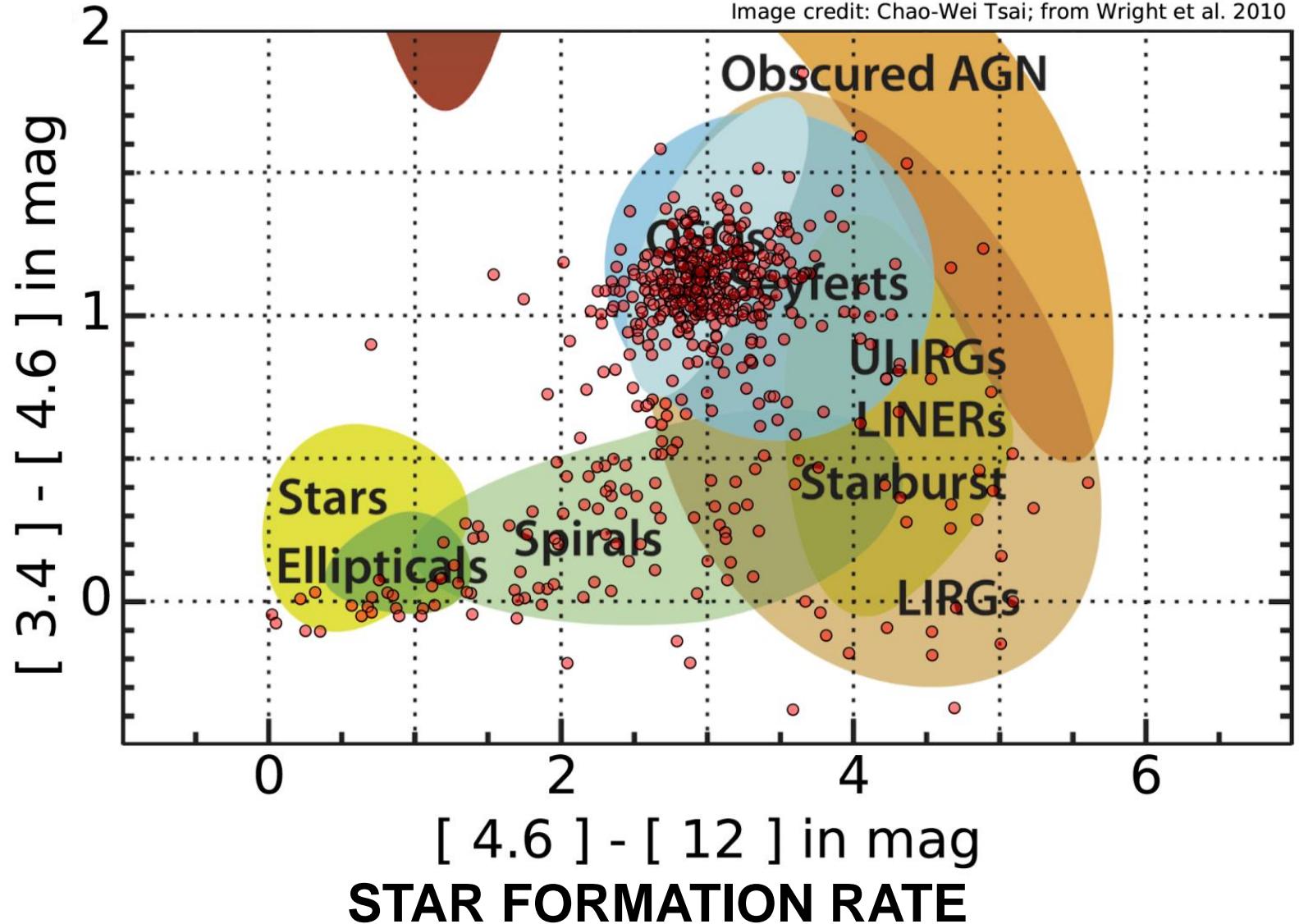


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# WISE Colour selection

Image credit: Chao-Wei Tsai; from Wright et al. 2010

AGN GAS ACCRETION RATE



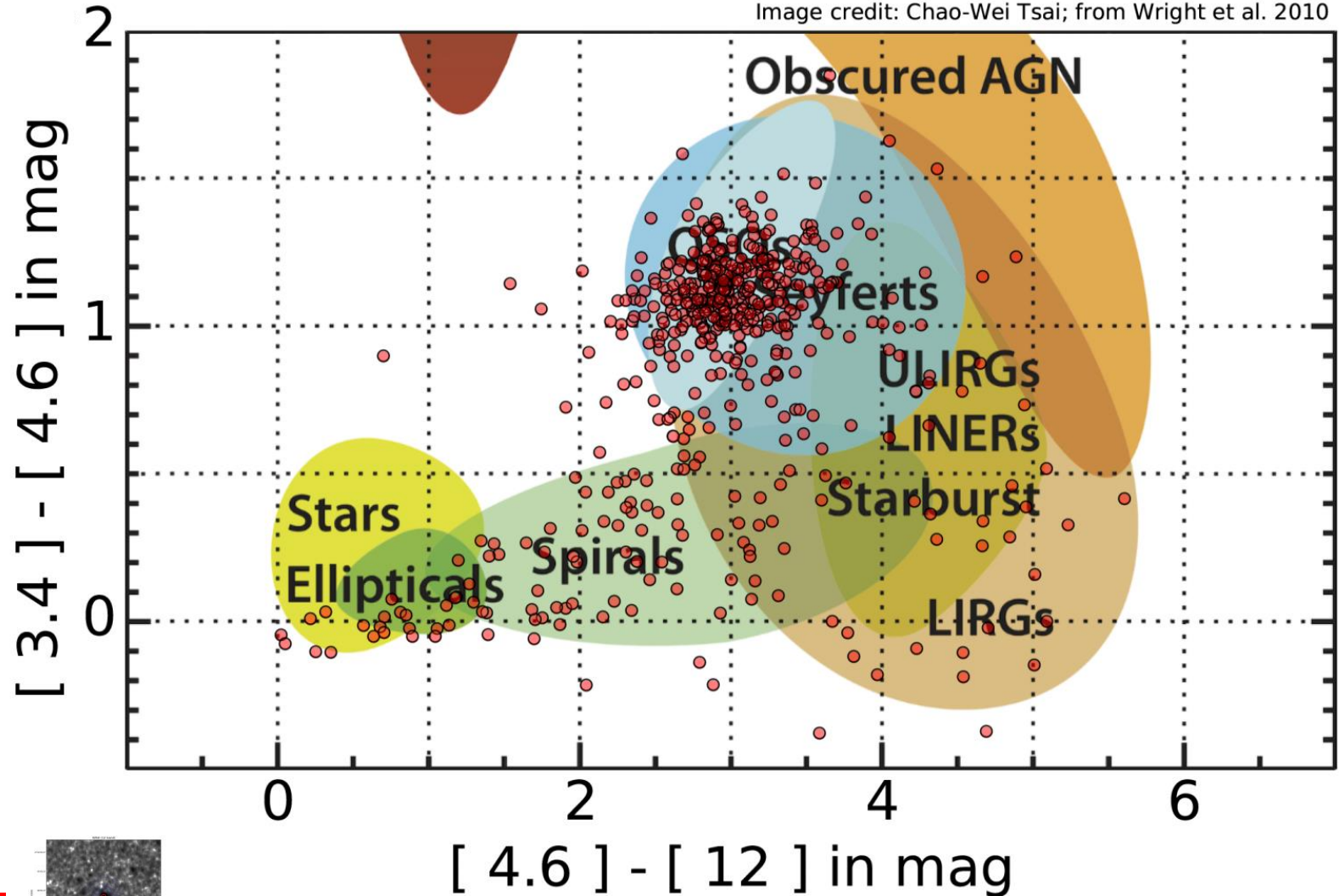


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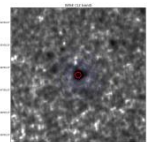
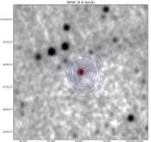
# WISE Colour selection

Image credit: Chao-Wei Tsai; from Wright et al. 2010

AGN GAS ACCRETION RATE



STAR FORMATION RATE



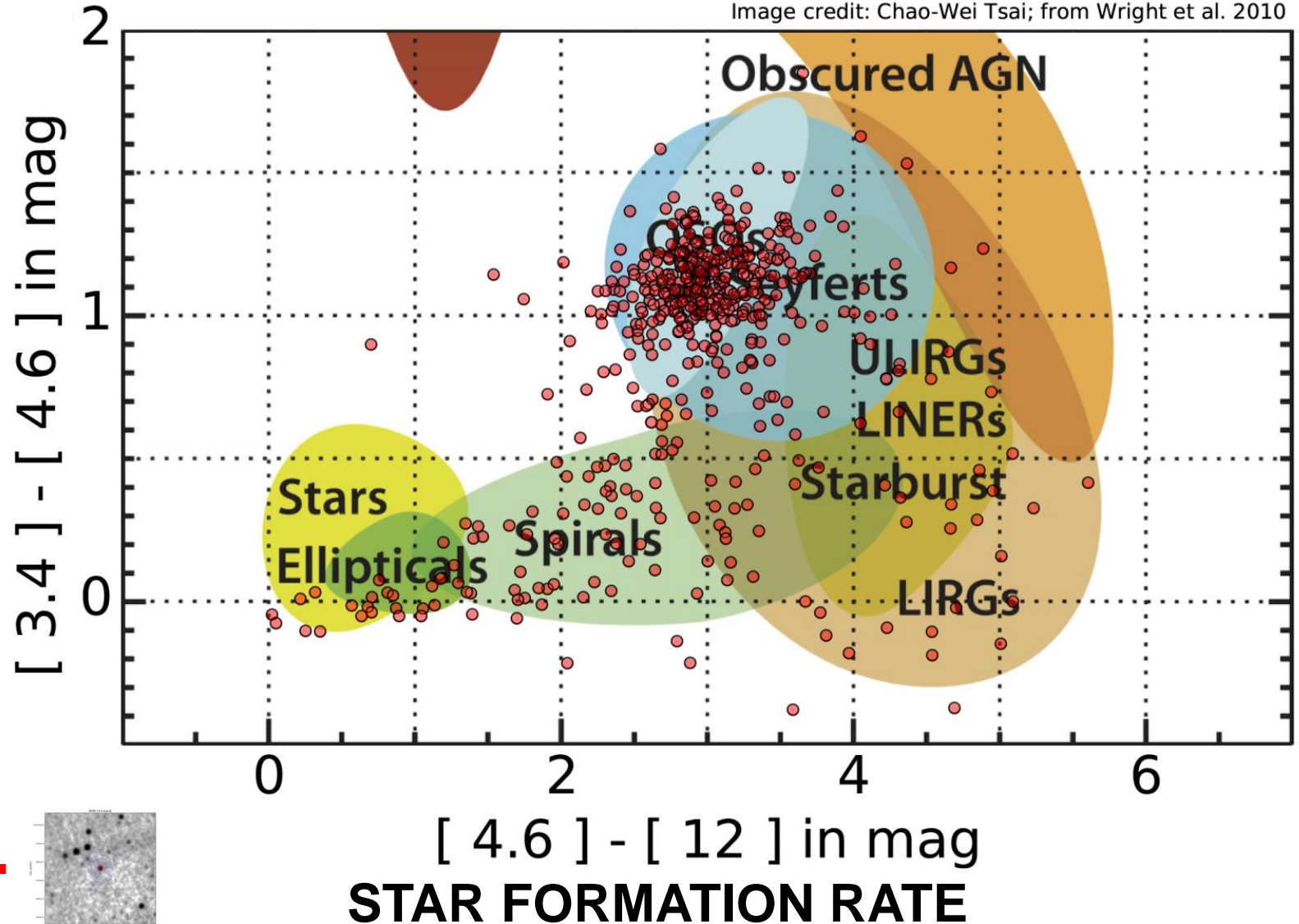




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# WISE Colour selection

Image credit: Chao-Wei Tsai; from Wright et al. 2010



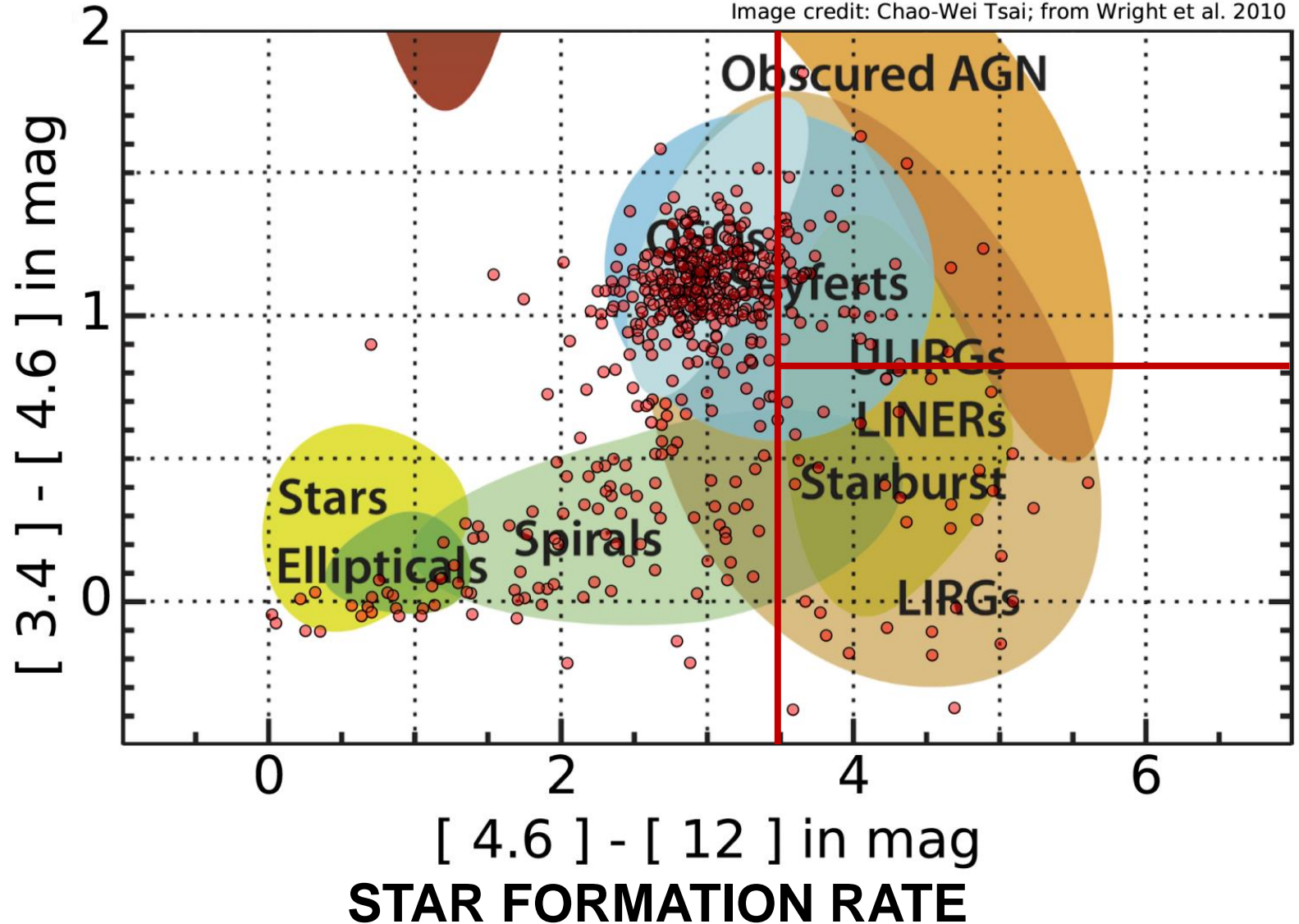




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AGN GAS ACCRETION RATE

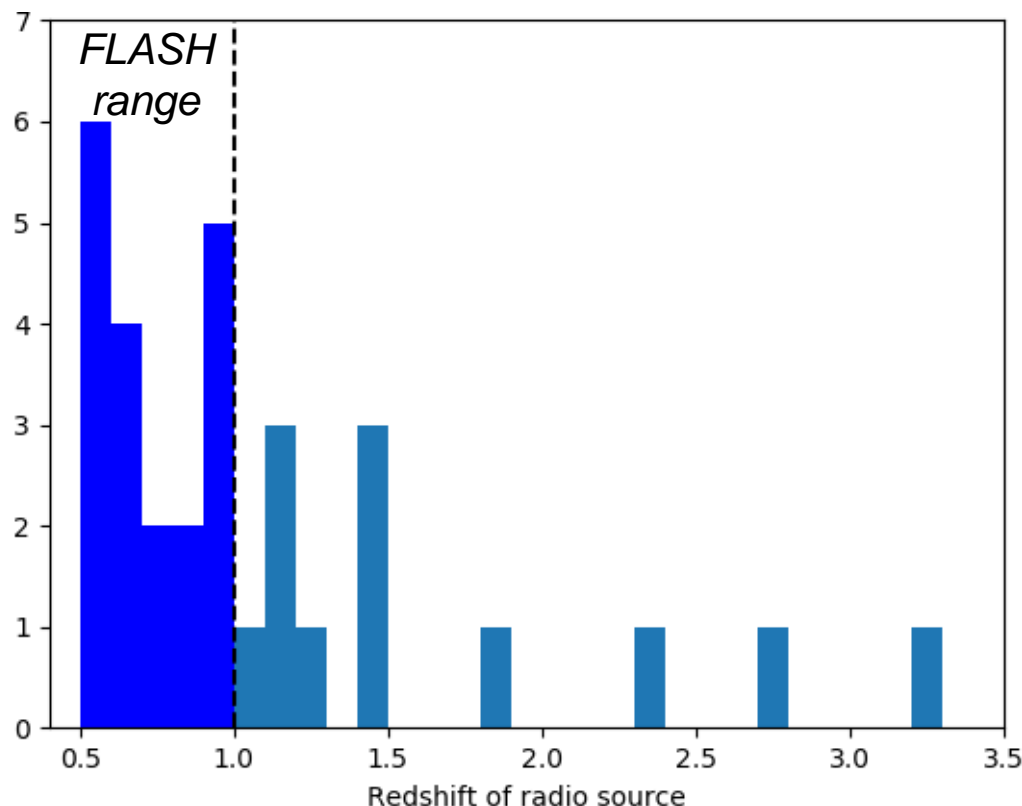




# Redshift distribution

- **Associated and intervening search for HI towards 39 radio-bright sources**
- Within  $0.4 < z < 1.0$  range: 19 sources
- Some redshifts unreliable – e.g. photometric estimates
- 7 radio sources with no known redshift

	Optically Faint	Mid-Infrared
Number	31	15
$0.4 < z < 1$	15	11
Detections	3	0



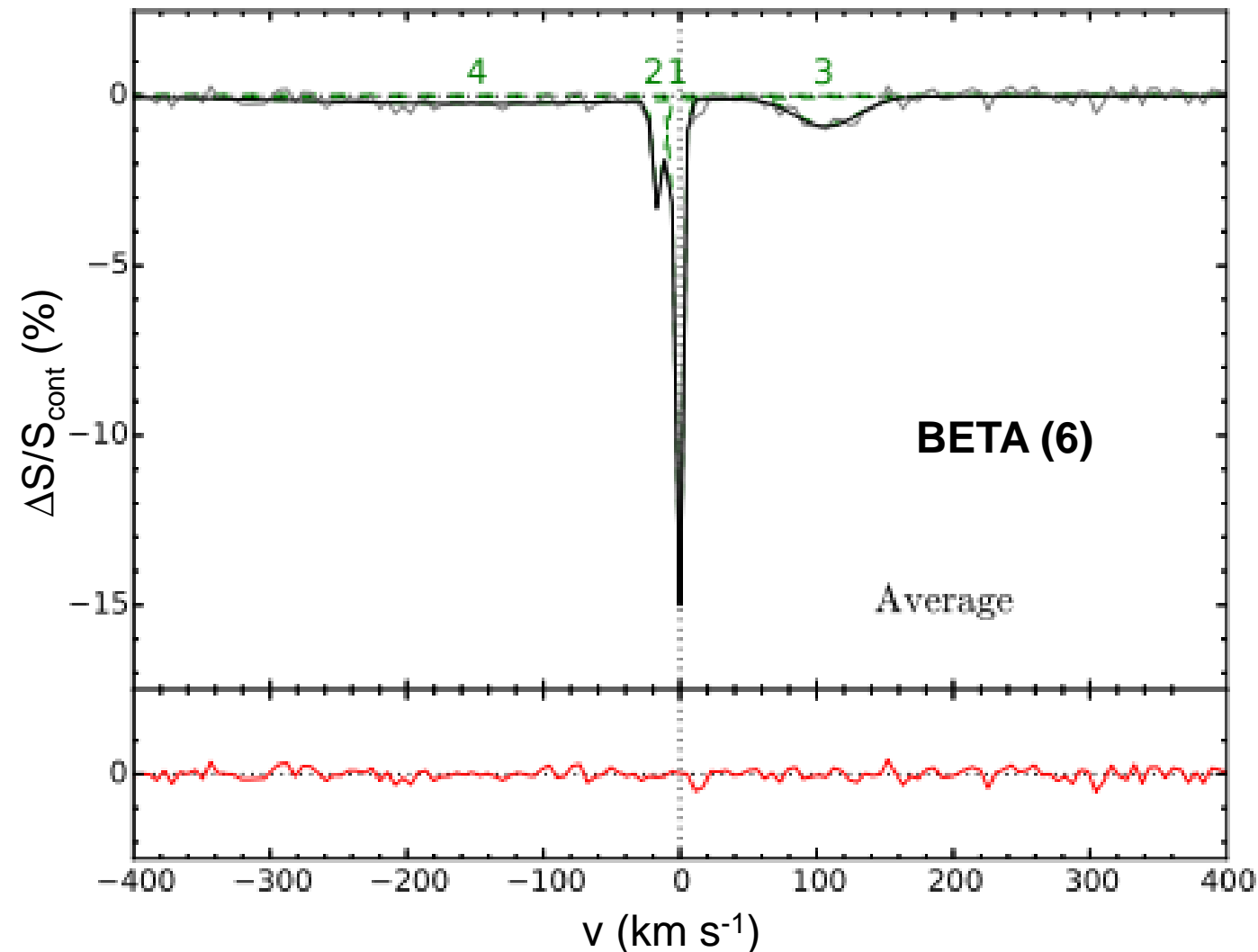


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# New Detections

**PKS 1740-517:  $z = 0.44$**

*Allison et al. 2015, MNRAS*



FLASH HI samples with ASKAP-BETA targeted for example:

- Giga-Hertz Peaked Spectrum (GPS) sources
- Soft X-ray absorbed sources (**Vanessa's talk!**)
- Optically faint sources

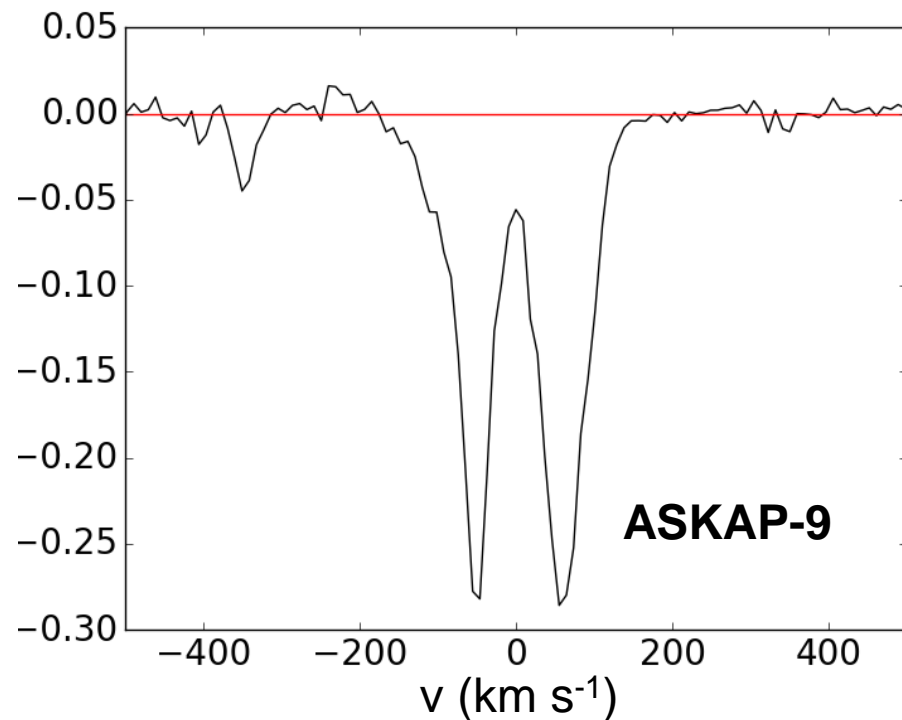
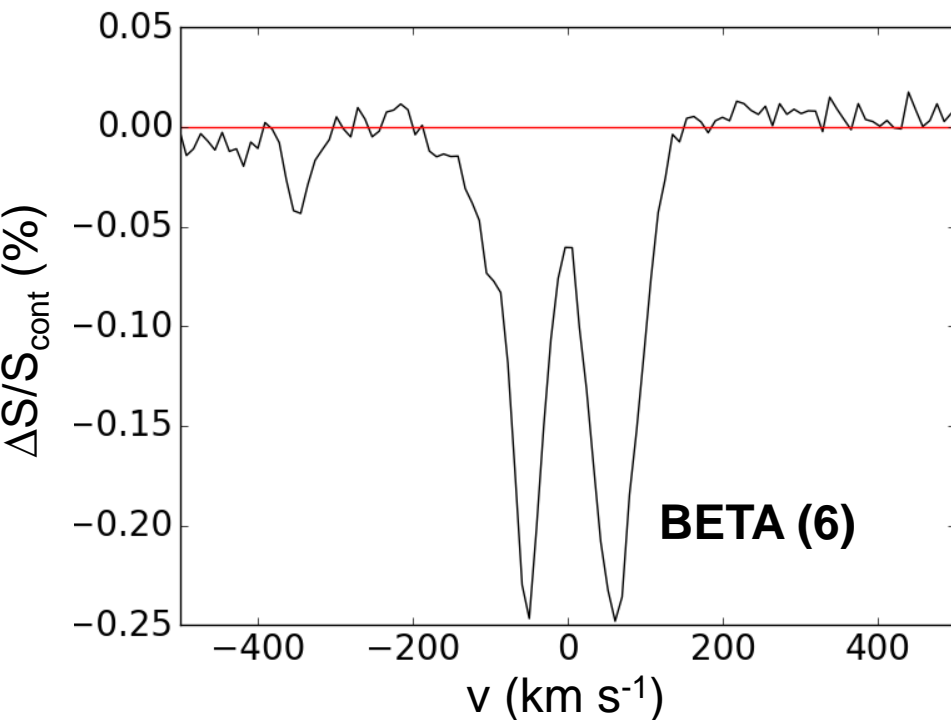
PKS 1740-517 falls into these three groups!



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# New Detections

**PKS 1829-718:  $z = 0.54$**



Searched to  $\sim 23$  mag (Costa 2001) with only 'possible identification'  
Intervening or associated? If associated, the twin dips could be gas rotating around a disk with the third feature an outflow



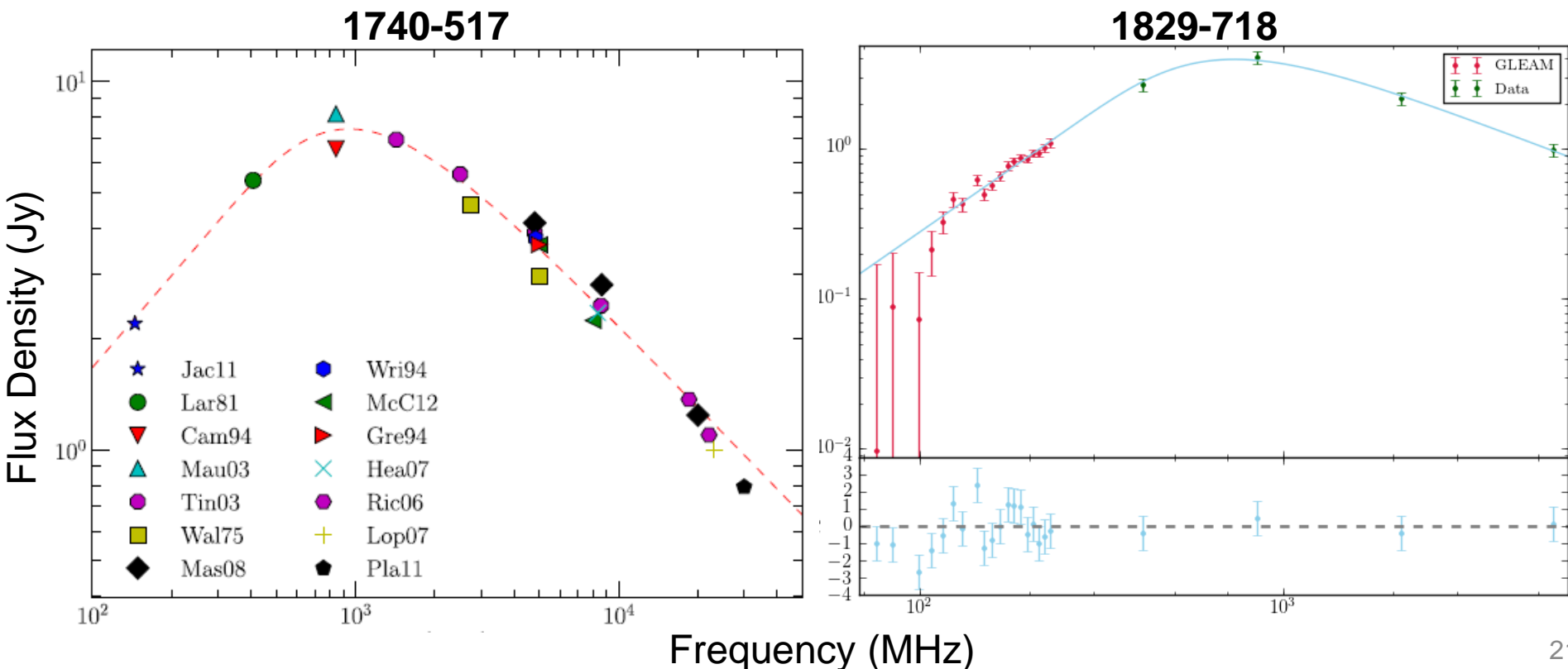


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# Common links?

- PKS 1740-517 and PKS 1829-718 are associated with Giga-Hertz Peaked Spectrum (GPS) sources
- GPS sources thought to be young/recently retriggered AGN

*Left: Allison et al. 2015. Right: MWA GLEAM data, courtesy of Joseph Callingham.*

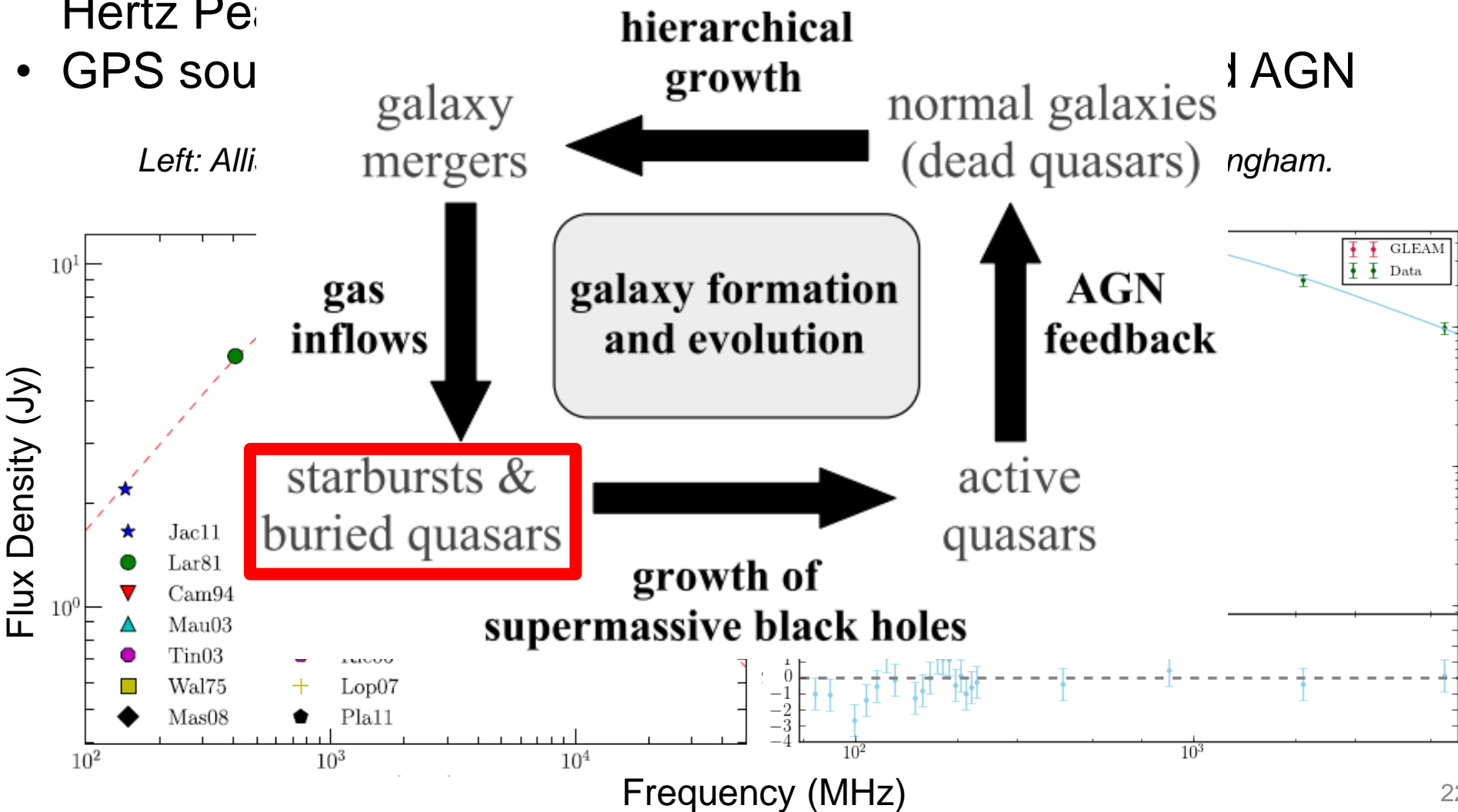




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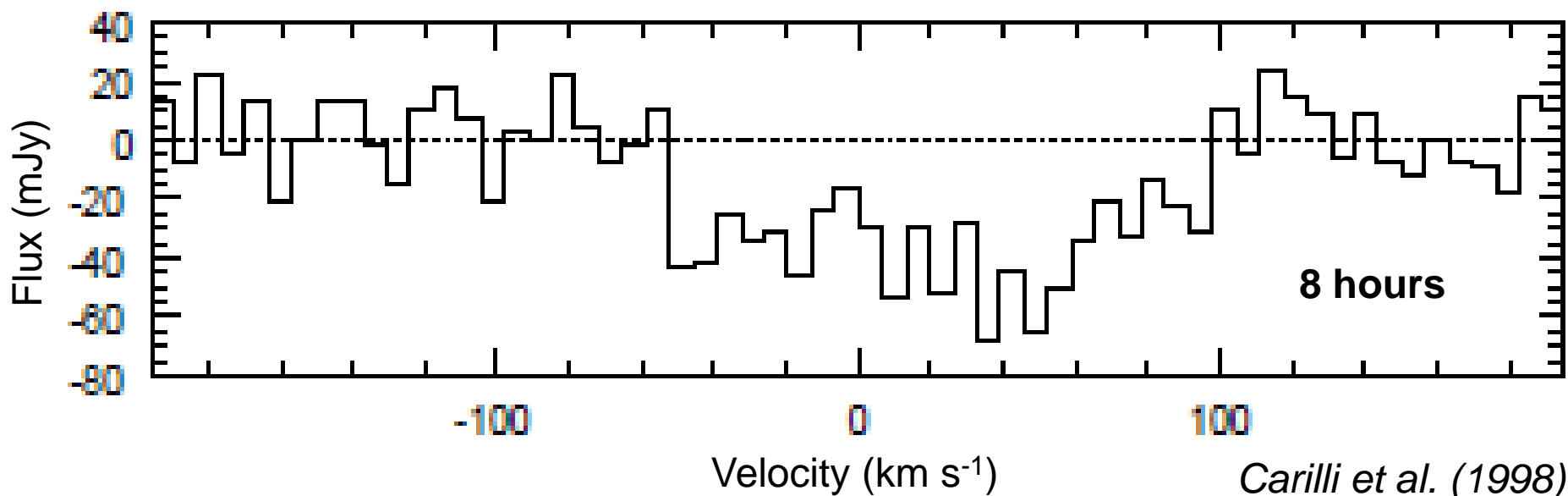
# Common links?

- PKS 1740-517 and PKS 1829-718 are associated with Giga-Hertz Peaked Spectrum (GPS)
- GPS sources are associated with AGN



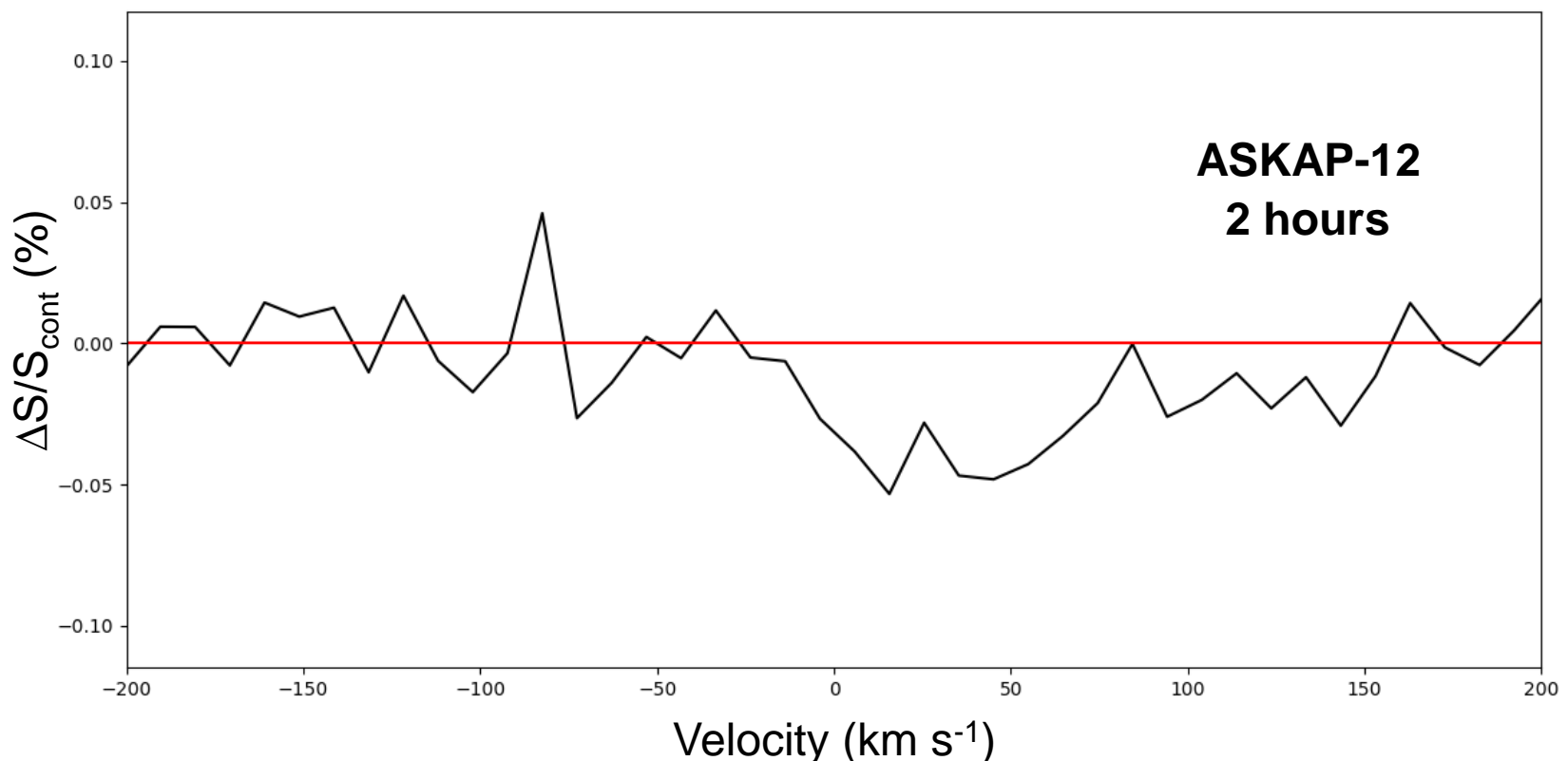


- **PKS 0500+019**
- Redshift of absorption feature is  $z = 0.585$
- Optical depth  $\sim 4\%$
- Possibly within an intervening galaxy – background quasar?





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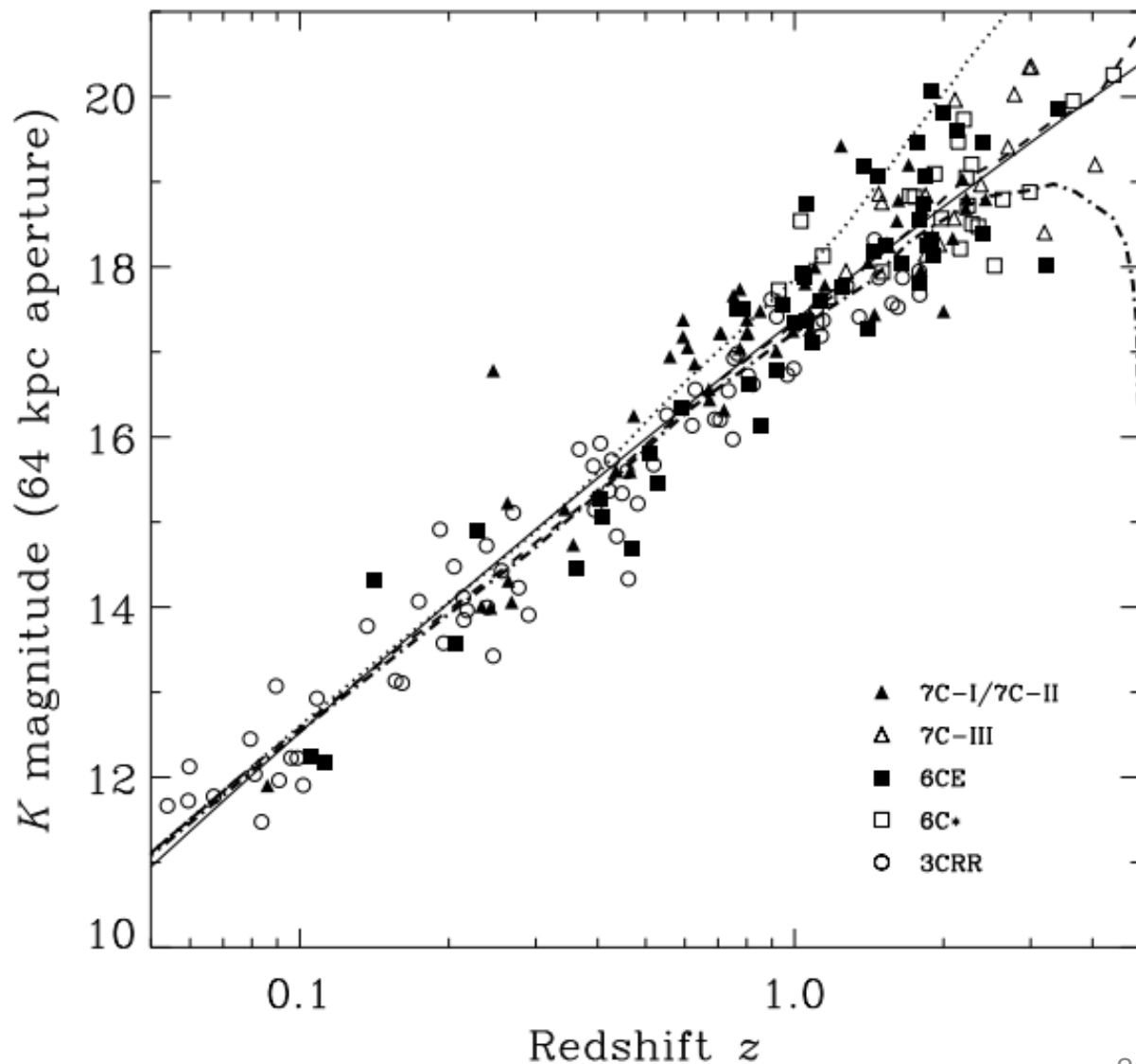






# Redshift Estimate?

- Can we estimate the redshift of radio galaxies without optical spectroscopy?
- K-z relation exists for radio galaxies (Willott et al. 2003)
- Can explore similar wavelengths for relationship with redshift
- WISE: an all-sky survey that offers this!

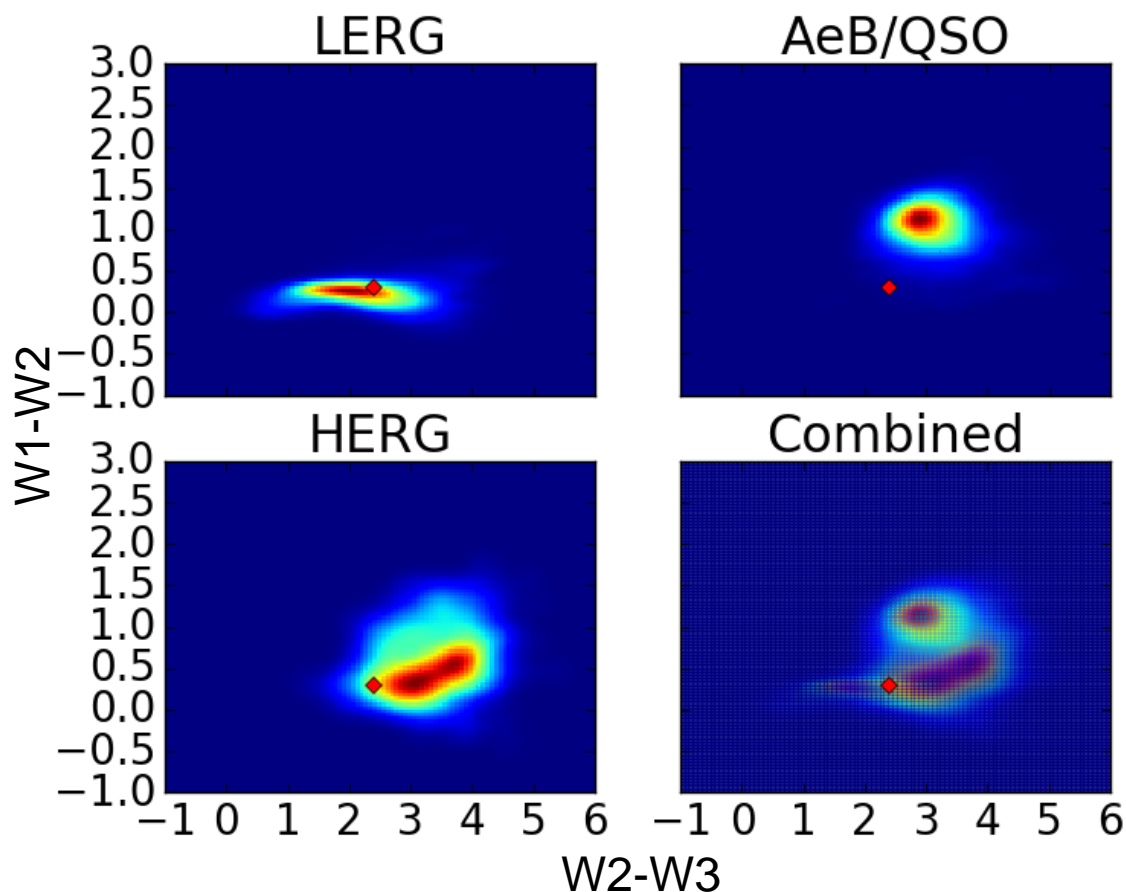


*Willott et al. (2003)*



# WISE class estimate

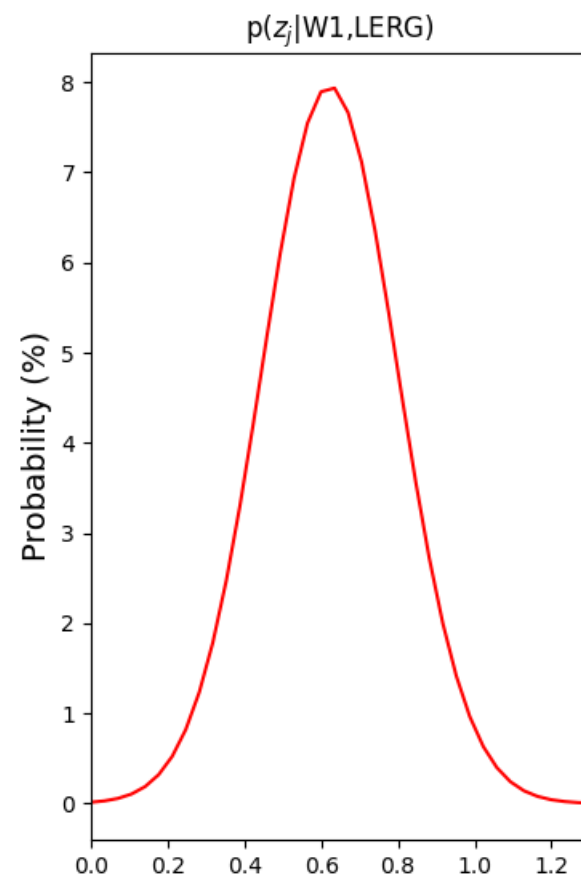
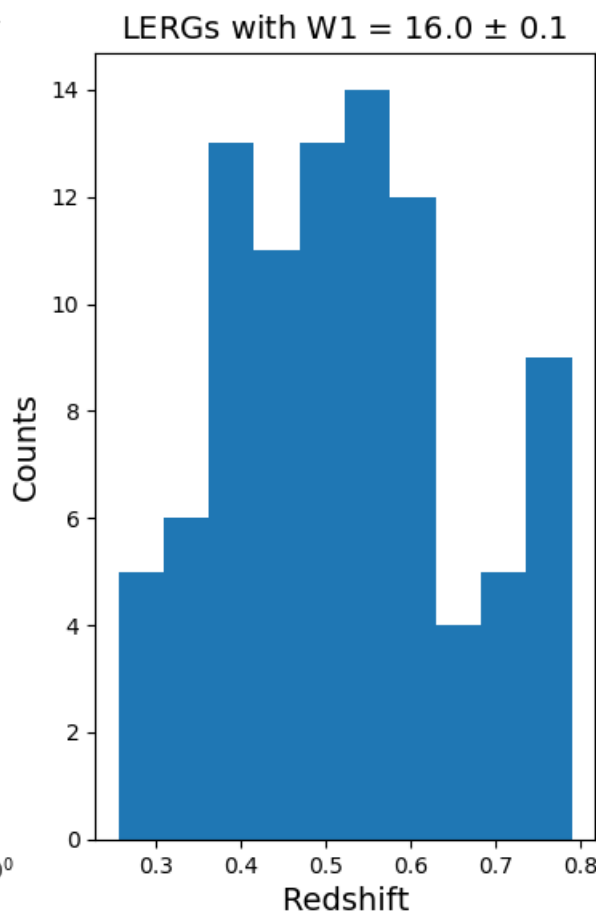
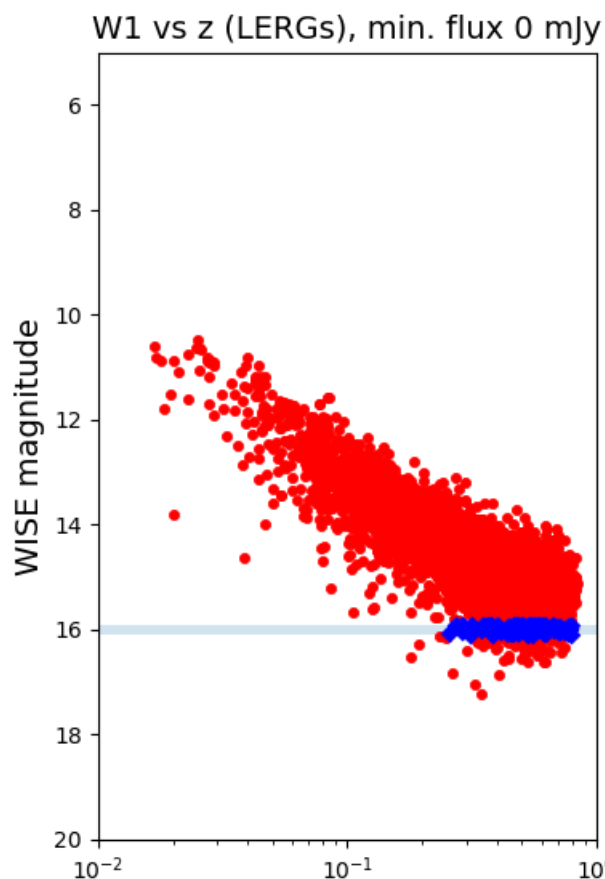
- Use LARGESS sample (Ching et al. 2017) as training set
- Spectroscopic redshifts and class identification:
  - Low-excitation radio galaxy (LERG)
  - High-excitation radio galaxy (HERG)
  - QSO
- Can estimate radio galaxy class based on WISE colour information





# WISE redshift estimate

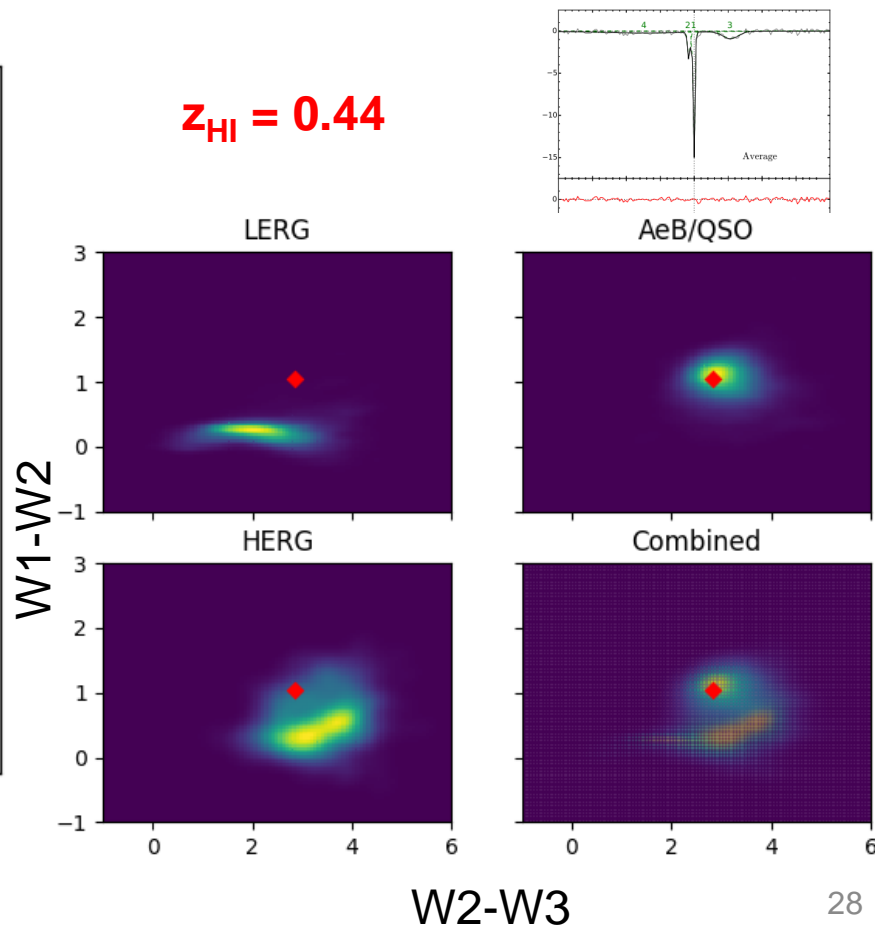
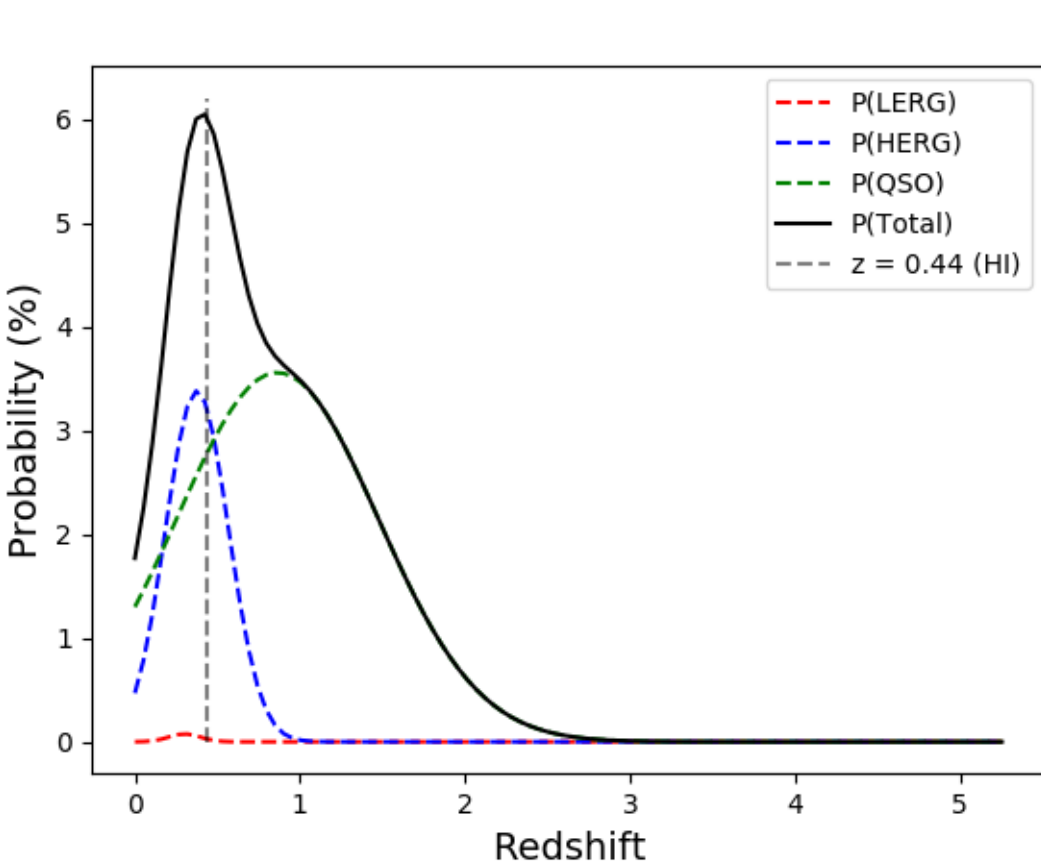
- From WISE magnitude, generate redshift distributions
- Repeat for all radio galaxy classes, then weight by likelihood of each class





# WISE redshift estimate

- **PKS 1740-517** redshift probability distribution
- Estimate from WISE magnitudes agrees well with redshift
- QSO or HERG classification from WISE colour information

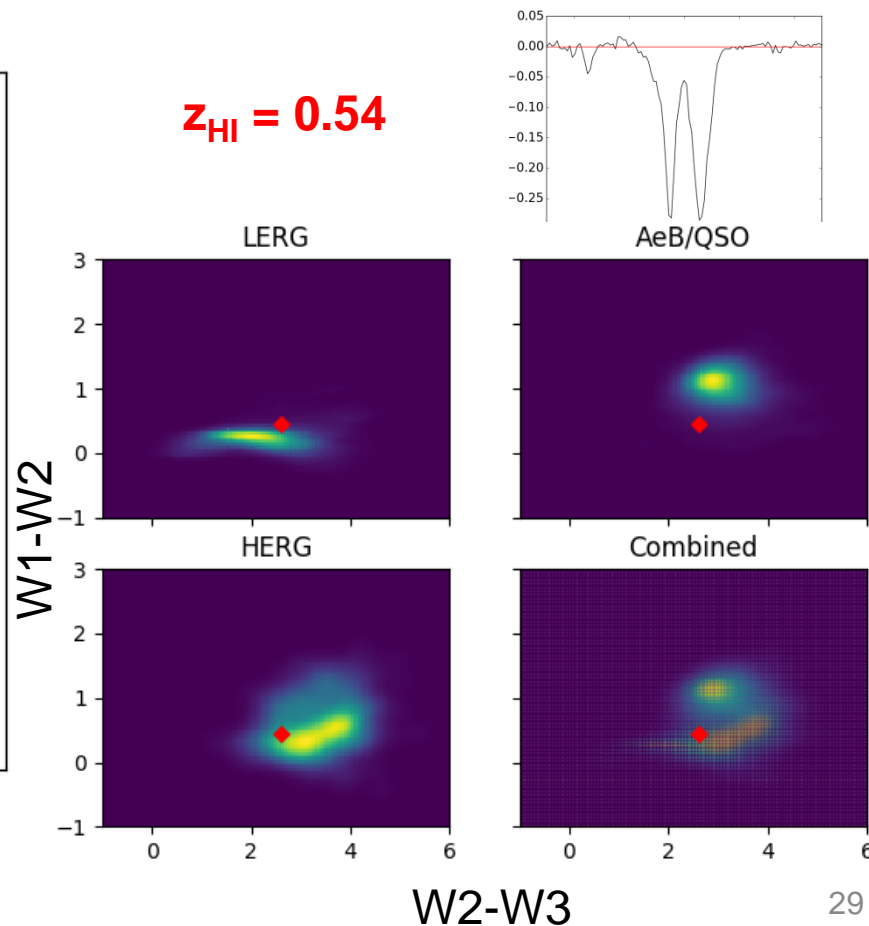
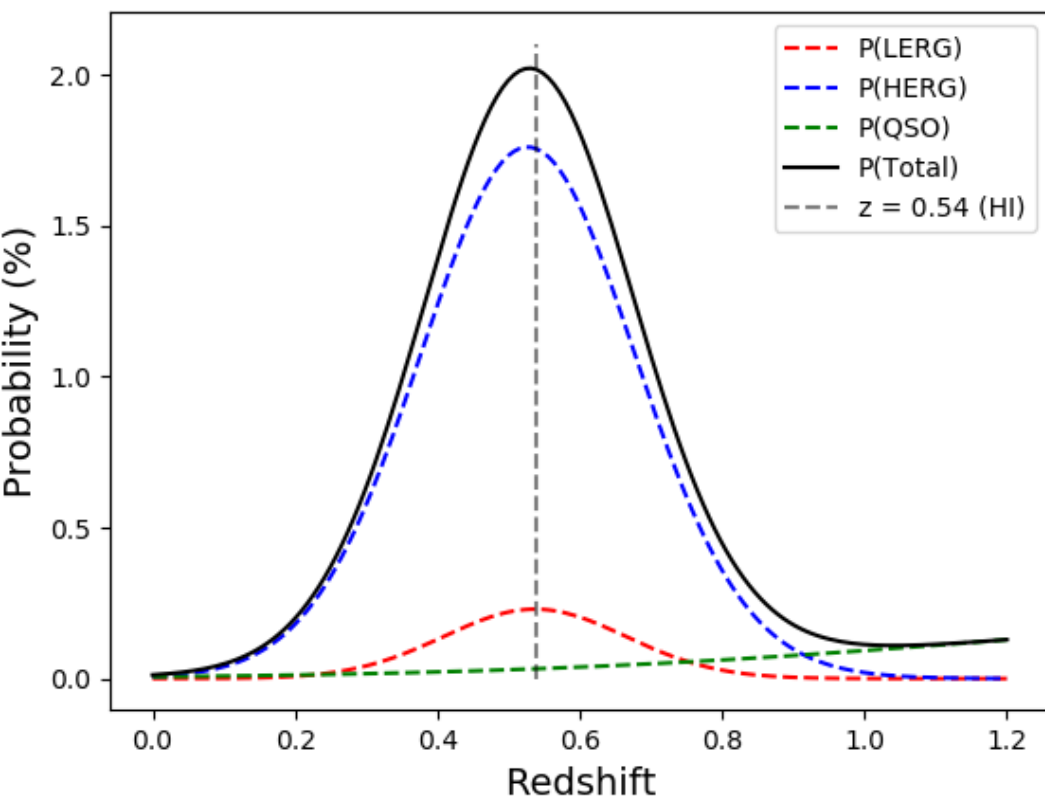




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# WISE redshift estimate

- **PKS 1829-718** redshift probability distribution
- Estimate from WISE magnitudes agrees well with  $z_{\text{HI}} = 0.54$
- Most likely to be a HERG from WISE colour information







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# Summary

- **Optically obscured quasars represent a short transitional stage in galaxy evolution**
- **Used the HI 21-cm transition in absorption to trace gas kinematics in a commissioning sample with **ASKAP-BETA**.**
- **Made three detections of HI ( $z = 0.44, 0.54, 0.58$ )**
- **Use mid-infrared WISE magnitudes to estimate redshift**
- **FLASH** will find HI in high-redshift galaxies across southern sky

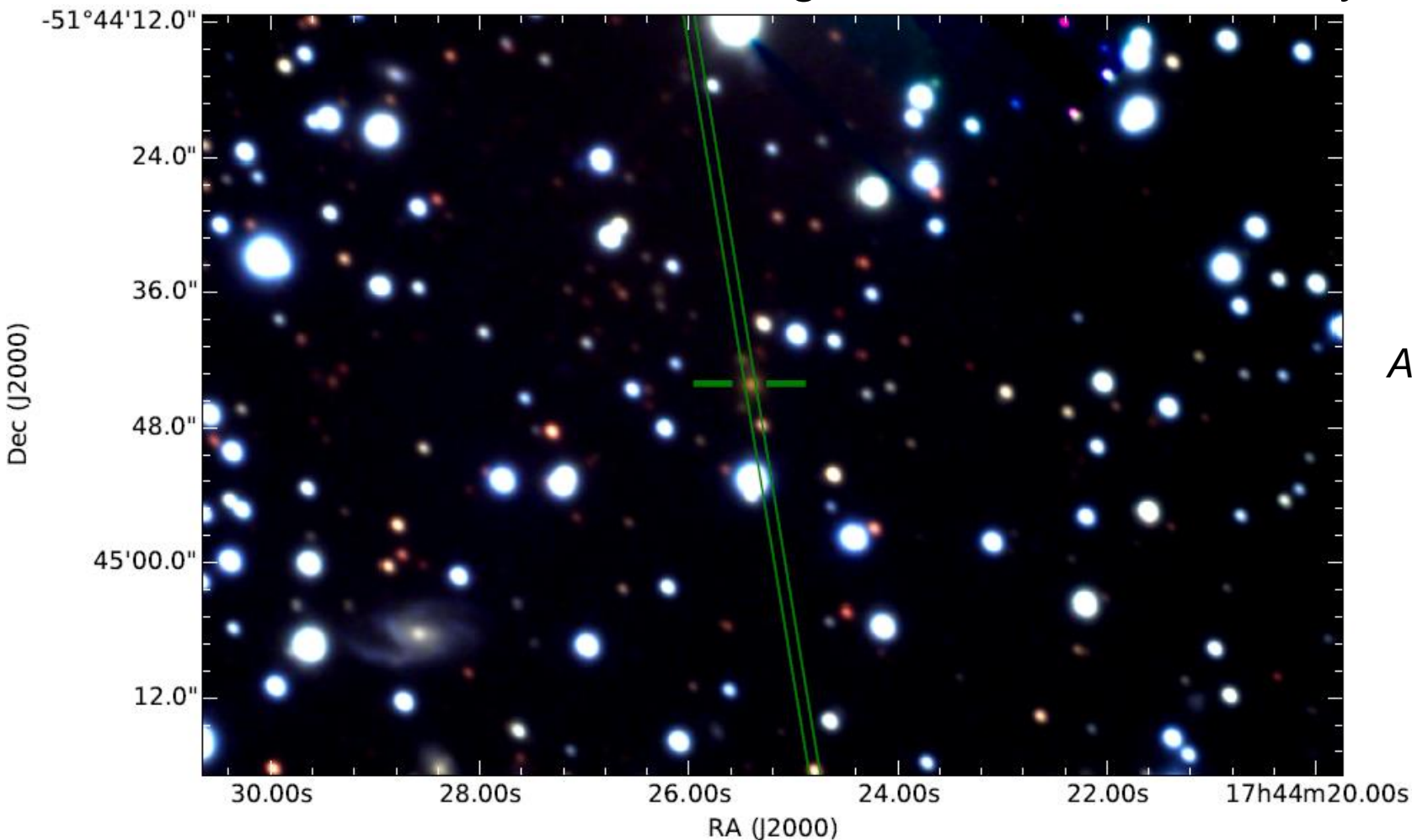




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# Multi-wavelength follow-up

- PKS 1740-517 followed up with Gemini to verify redshift
- What about all the other radio galaxies in southern sky?



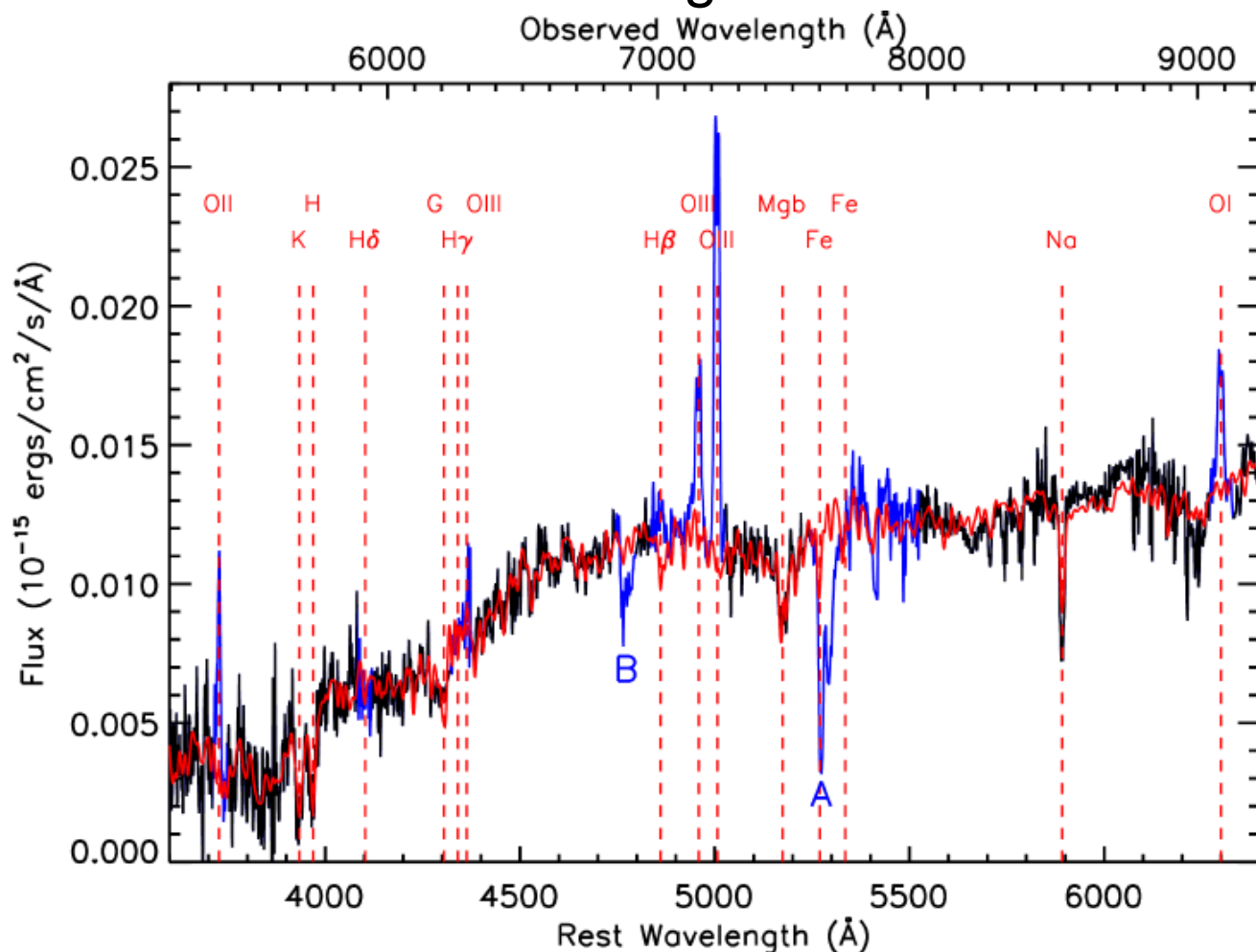
*Allison et al.  
(2015)*



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# Multi-wavelength follow-up

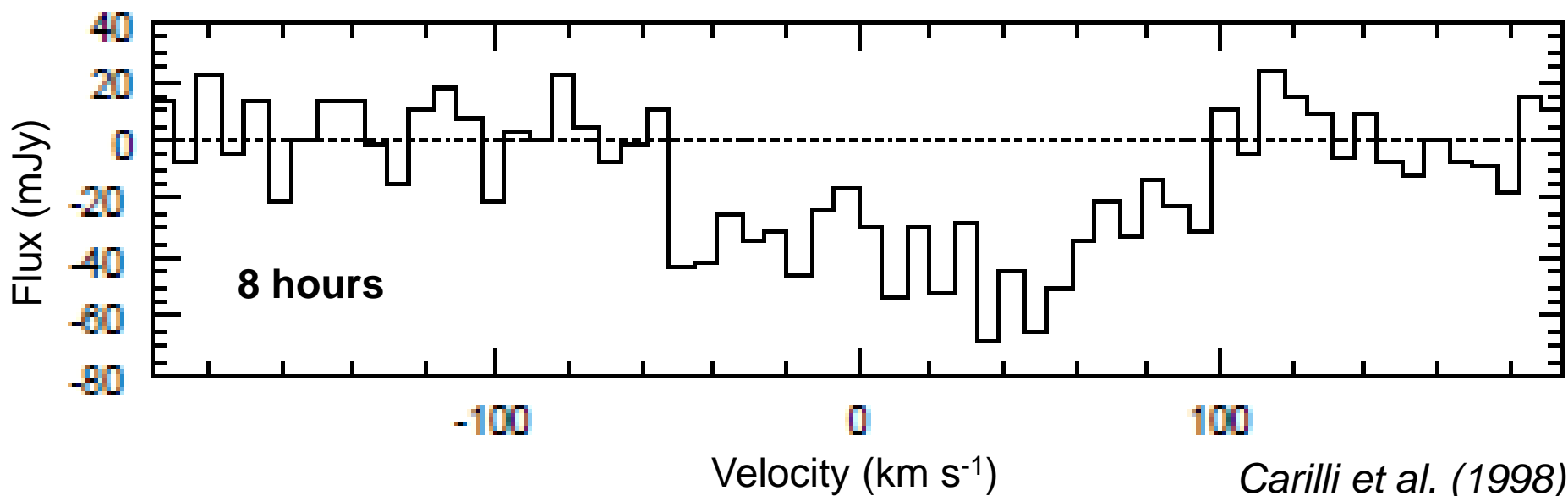
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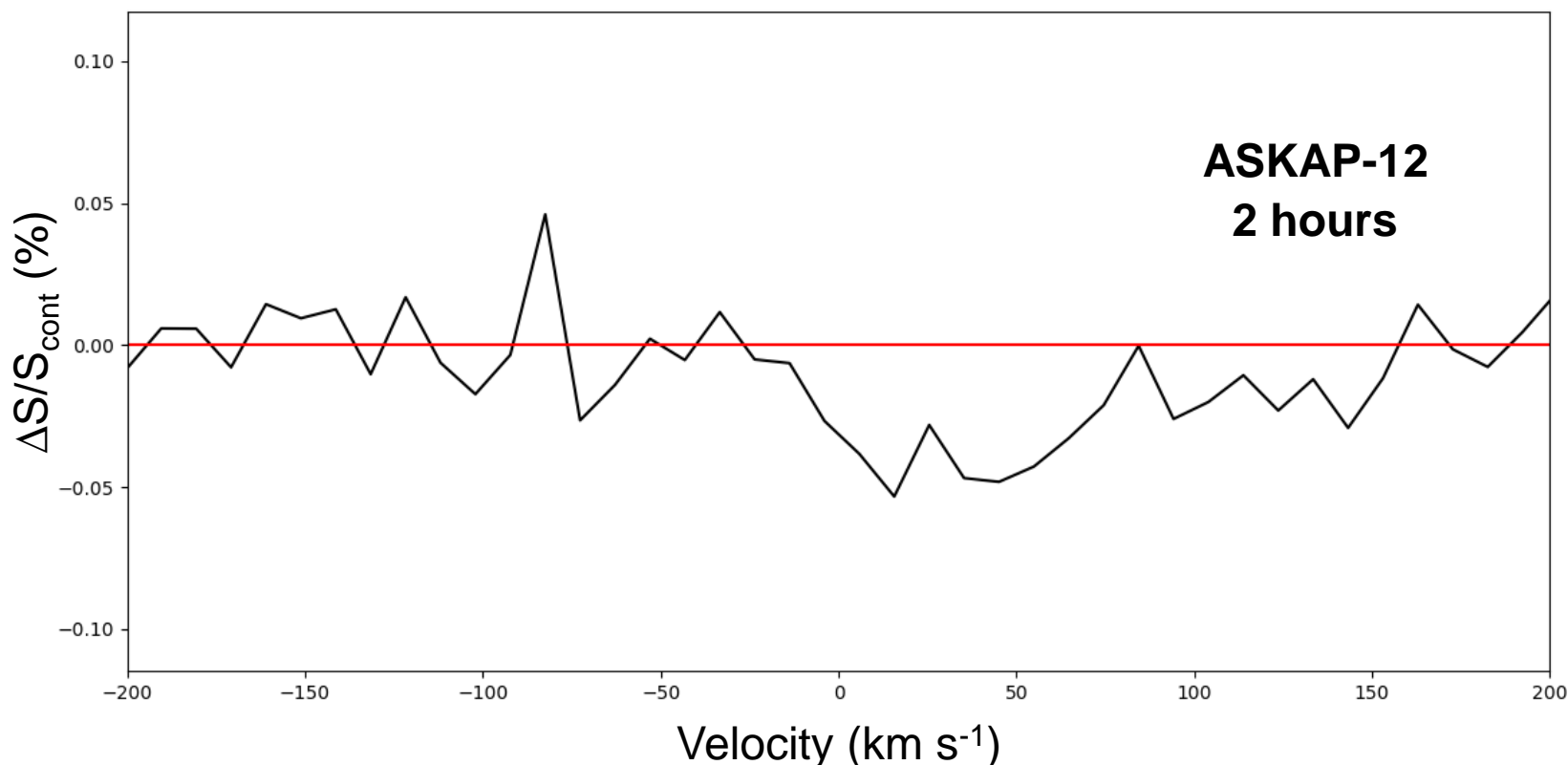


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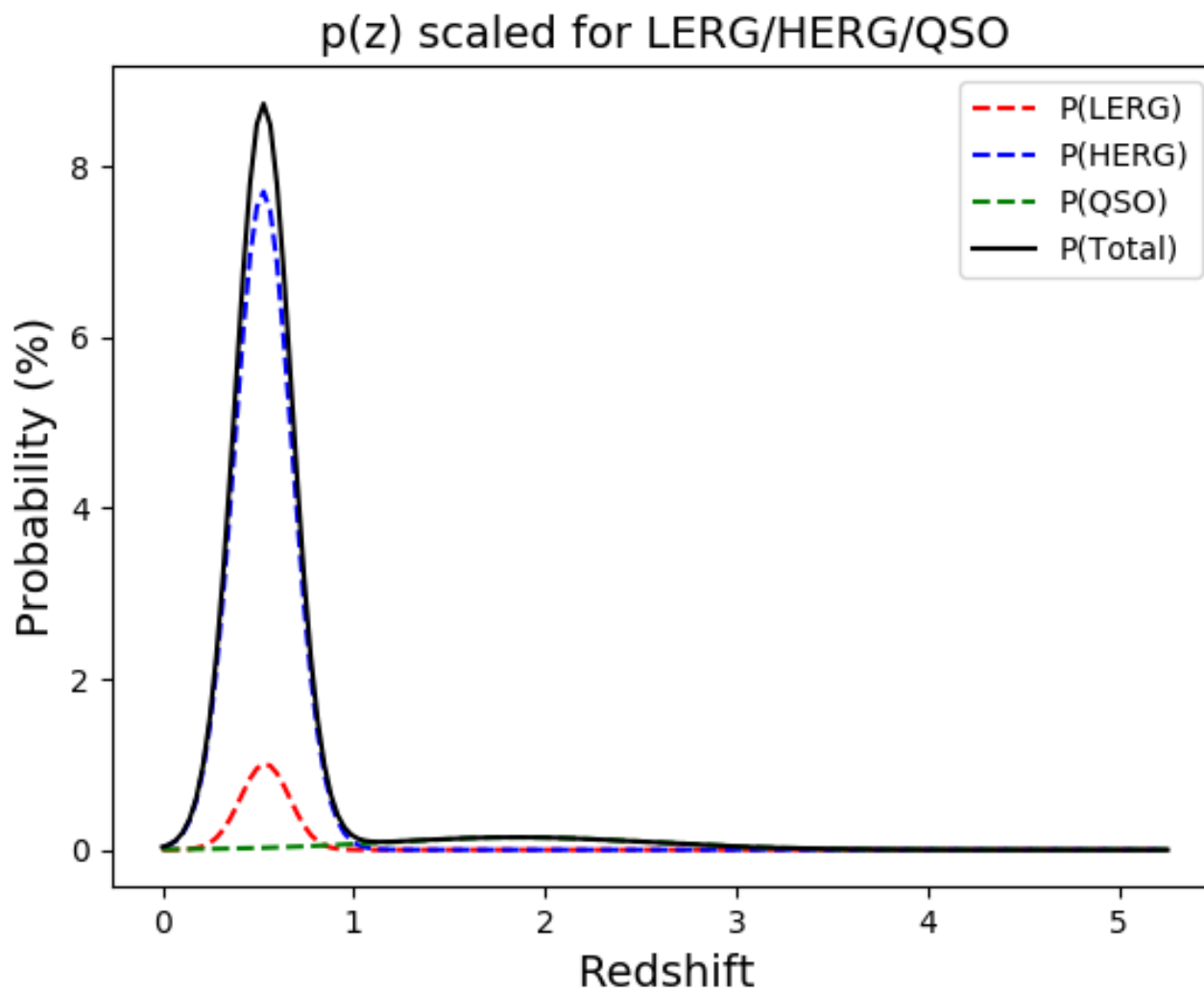




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# WISE Photometric Redshift

- PKS 0500+019 redshift probability distribution
- Lower redshift peak < HI redshift  $z = 0.585$
- Speculated to be a background galaxy (Stickel et al. 1996)

