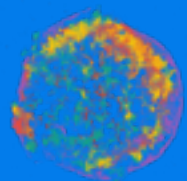


THE DARK ENERGY SURVEY: 3 YEARS OF SUPERNOVA

IN <15

MAT SMITH & THE DES SN TEAM



University of Southampton
Supernova Group



THE DARK ENERGY SURVEY

....SUMMARY: WE ROCK!

THE DARK ENERGY SURVEY: THE OVERVIEW SLIDE

AIM: “WHAT IS DARK ENERGY?”

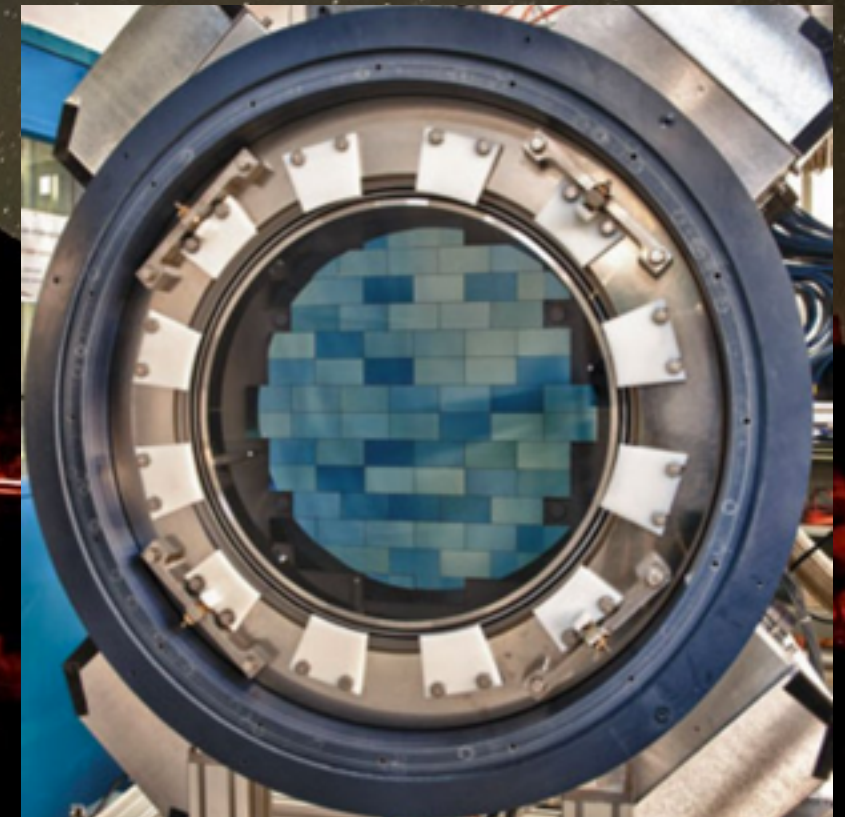
One survey, comprised of multiple probes, to determine the nature of cosmic acceleration

DES-Wide

- 5000 deg² in *grizY*: $r_{AB} \sim 24.3$, $i_{AB} \sim 23.5$ (10σ)
- Large Scale Structure, Weak Lensing & Galaxy Clustering

DES-SN

- ~ 6-day cadenced *griz* survey over 27 deg²
- Type Ia supernovae & other transients



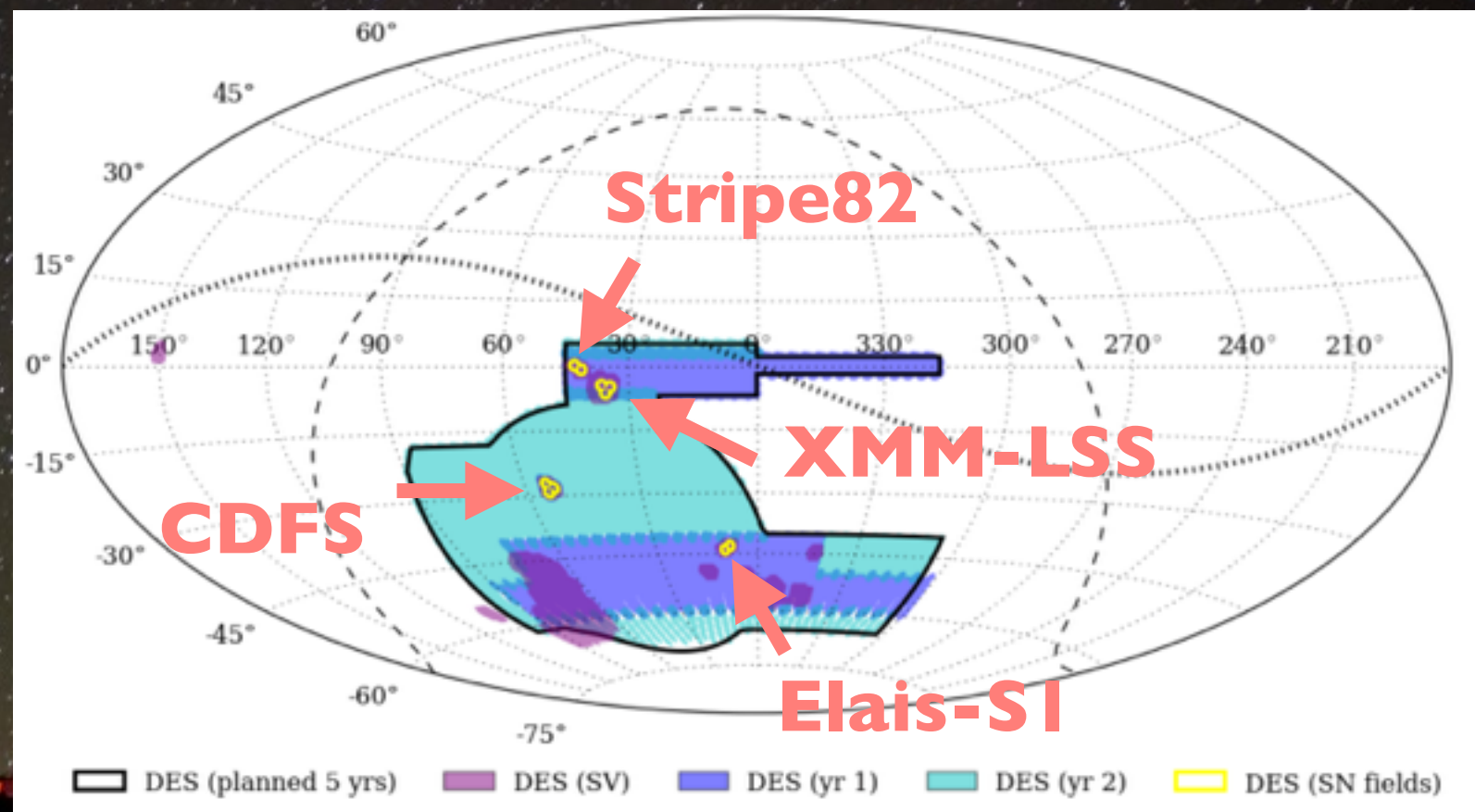
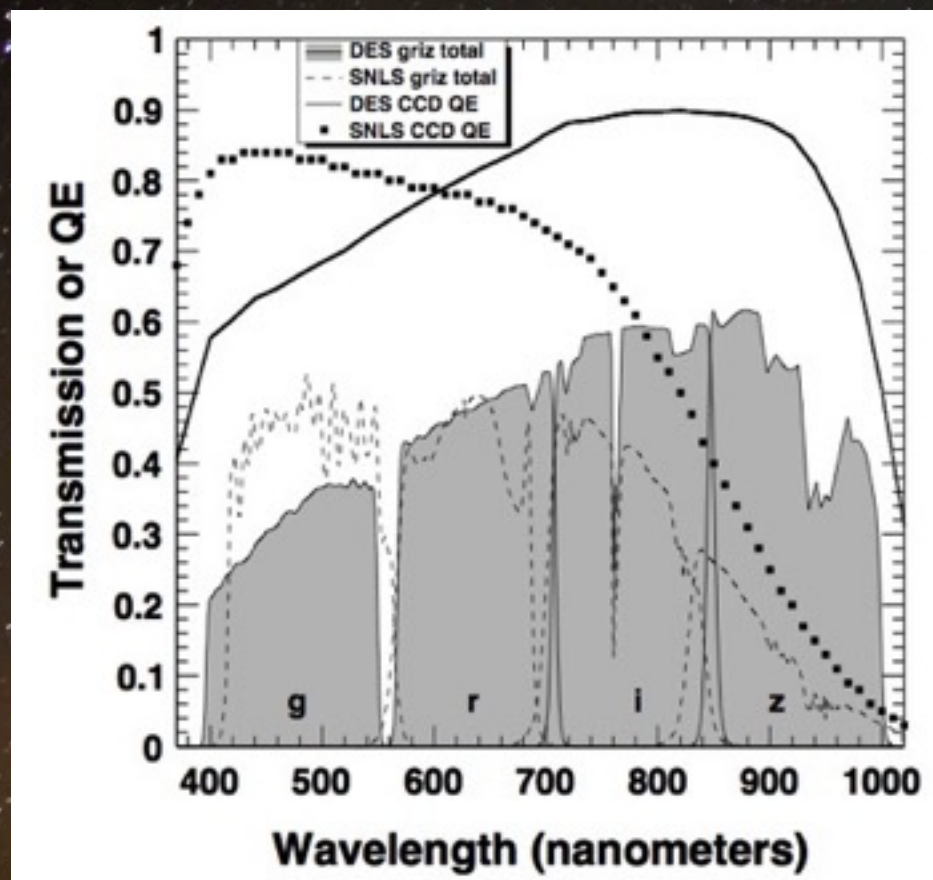
Overall

- Probes of both Distance v Redshift & Growth of Structure
- Multiple probes to break degeneracies & minimise systematics
- Lots of ancillary science (understatement alert)

- *See plenty of other talks.... (DES is great!)*

THE DARK ENERGY SURVEY: THE “HOW DOES IT WORK” SLIDE

ONE LARGE, RED, SURVEY IN THE SOUTH



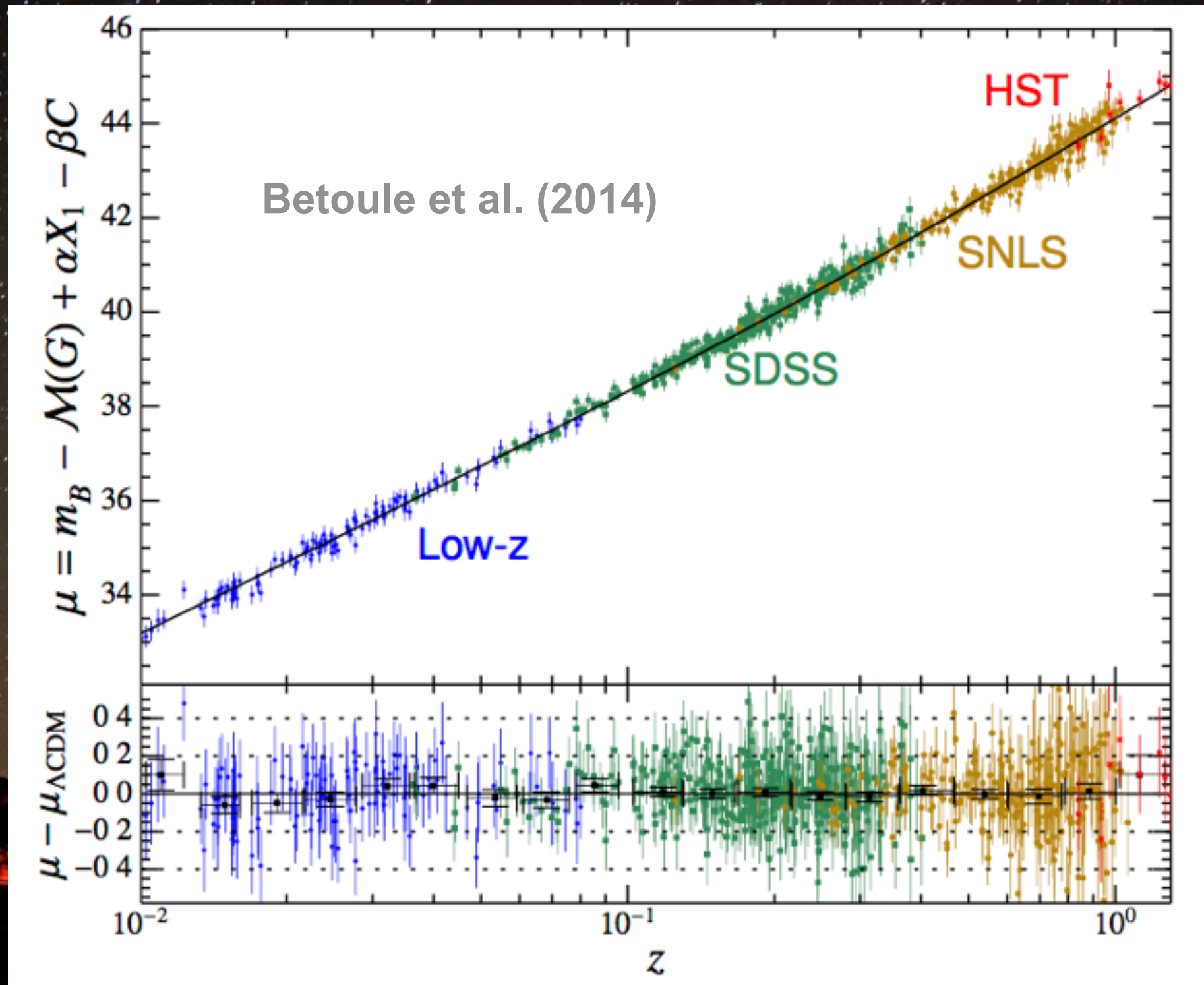
Dark Energy Camera (DECam) on 4m Blanco Telescope at CTIO

- 520 Mpx camera; 3 deg² FoV; deep-depleted LBNL CCDs
- Allocated 525 nights over 5 observing seasons (Aug - Feb)
 - *DESY4 starts 13 August 2016*



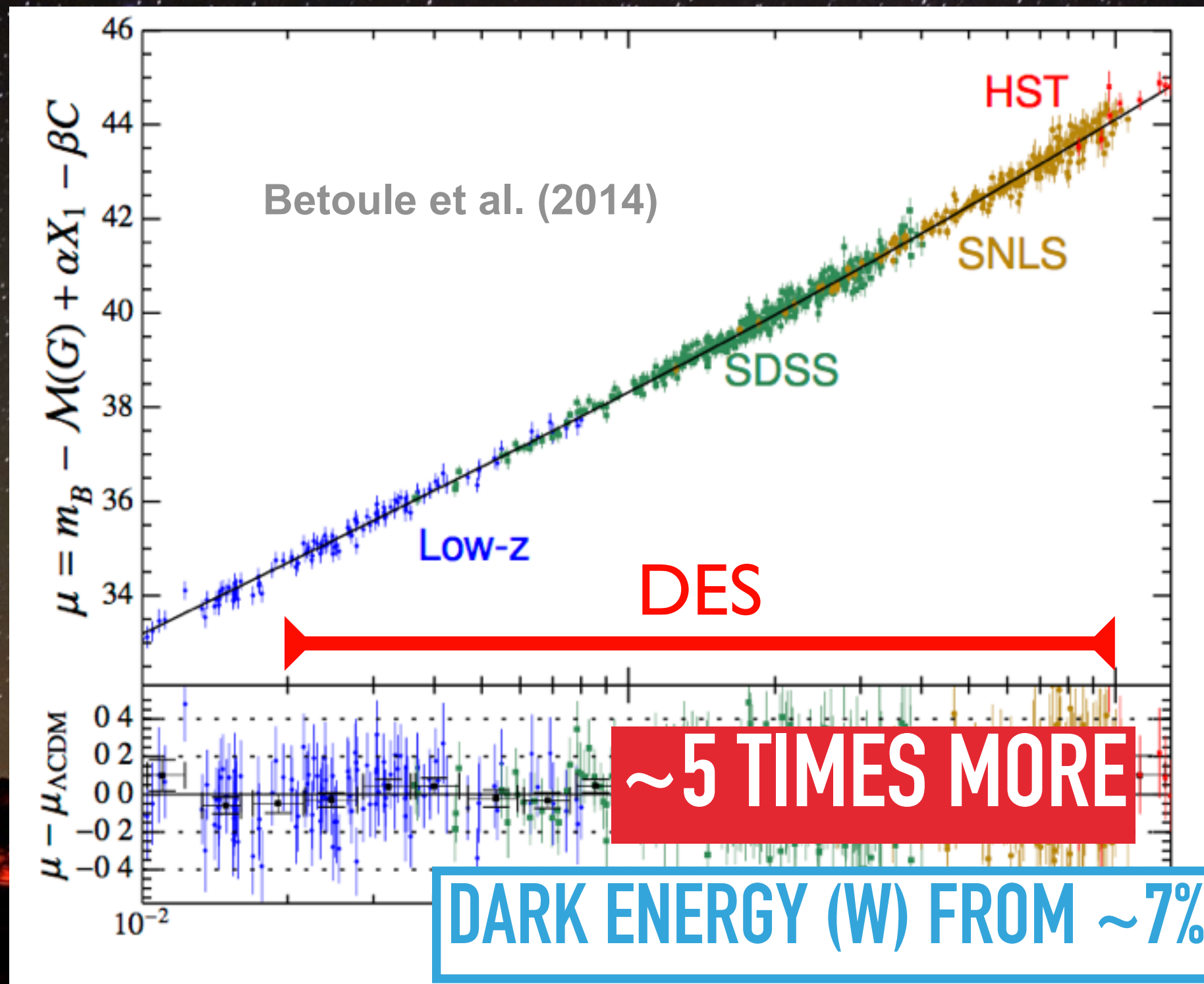
But this is a supernova talk....

TYPE IA SUPERNOVA TODAY



~750 SNIa, ALL SPECTROSCOPICALLY CONFIRMED, FROM DIFFERENT SURVEYS

TYPE IA SUPERNOVA AFTER DES

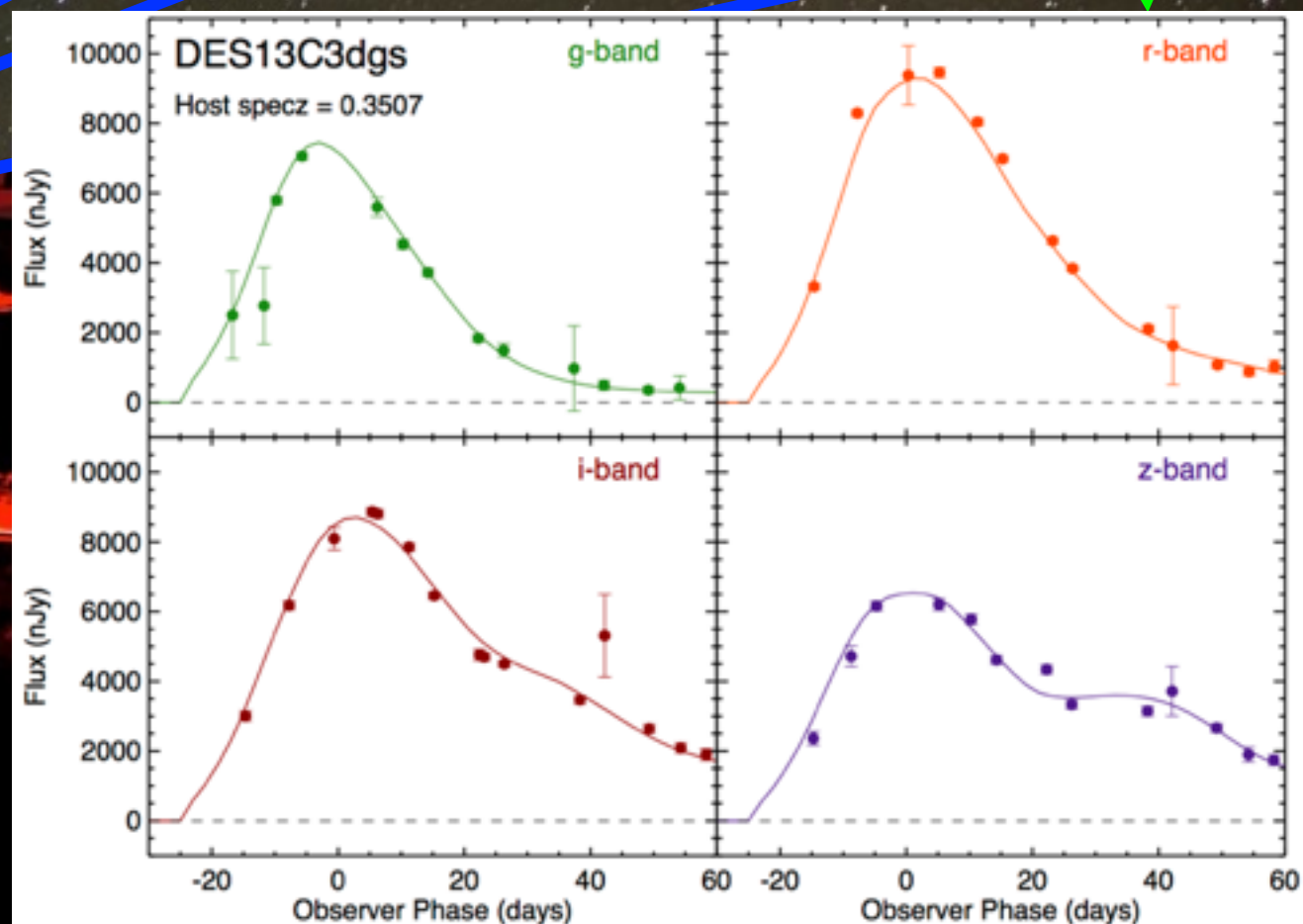
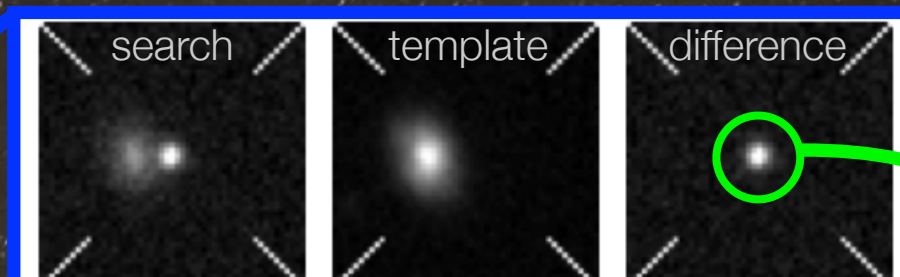
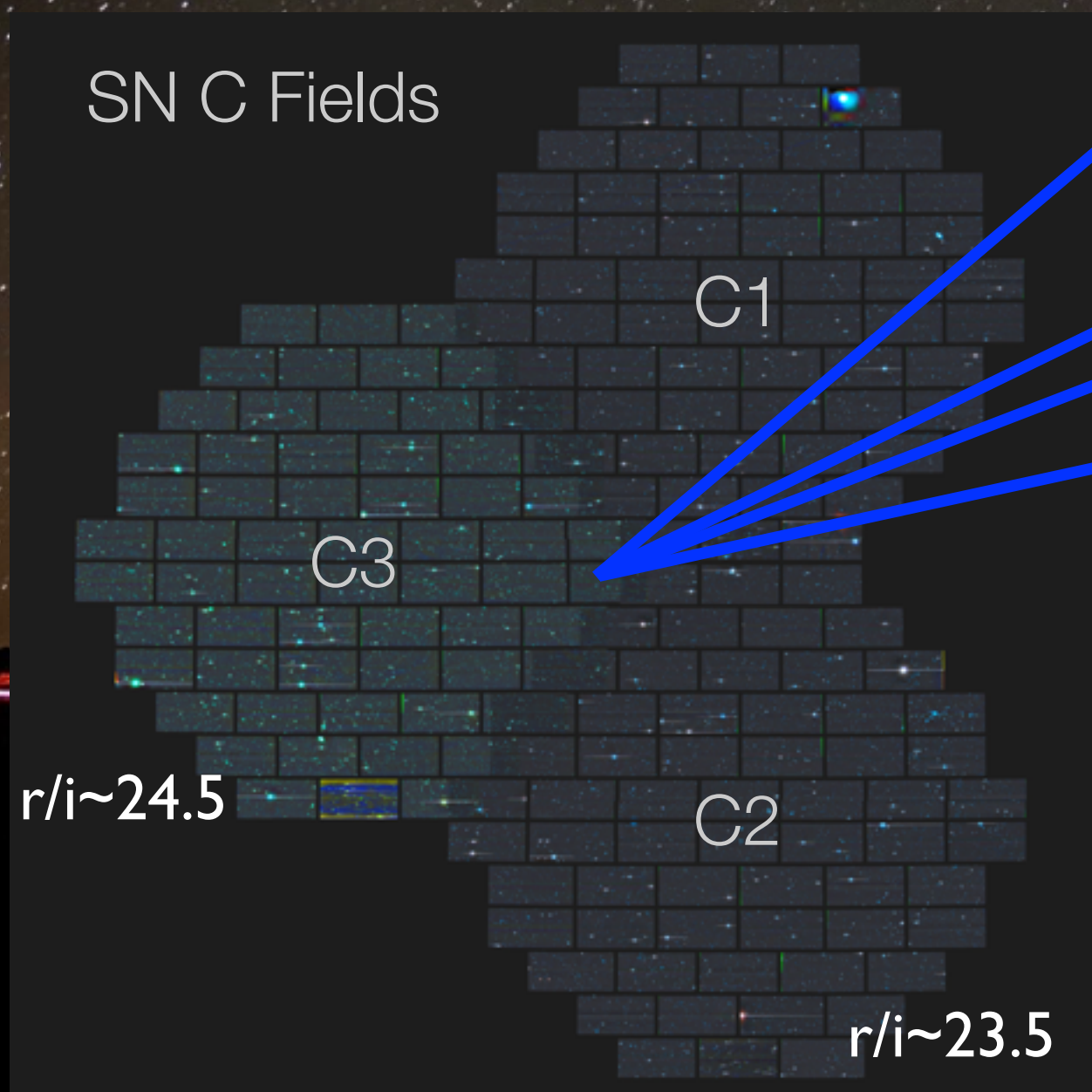


ALL IN ONE SURVEY: LIMITED CROSS-CALIBRATION SYSTEMATICS

THE DARK ENERGY SURVEY: SUPERNOVA PROGRAM

TO FIND, CHARACTERISE & FOLLOW 3000+ SN TO $z > 1$

- 10 fields (8 shallow, 2 deep) in 4 filters (griz)
 - Every 7 nights, for 5+ months for 5 years
- } **~3500 SNIa**

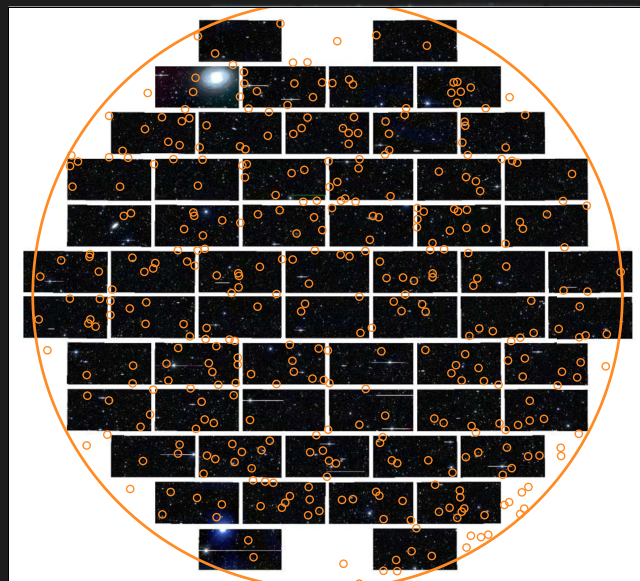


DES-SN: SPECTROSCOPIC FOLLOW-UP: THE WAY FORWARD

OZDES: REDSHIFTS FOR EVERYTHING

- 100 night survey on the AAT, overlapping with the DES fields
- Target the host galaxies of DES transients (and plenty of other things)
- Repeat observations gives an effective limiting magnitude of $r \sim 24.0$

SN C Fields



C1

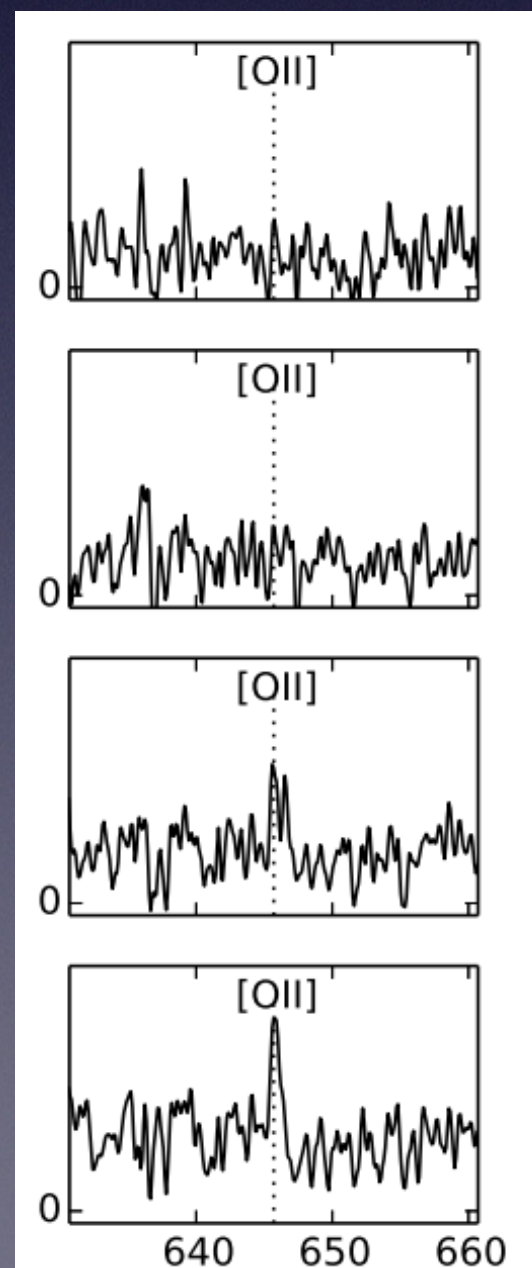
C2

Exp=7200 s

Exp=14400 s

Exp=32400 s

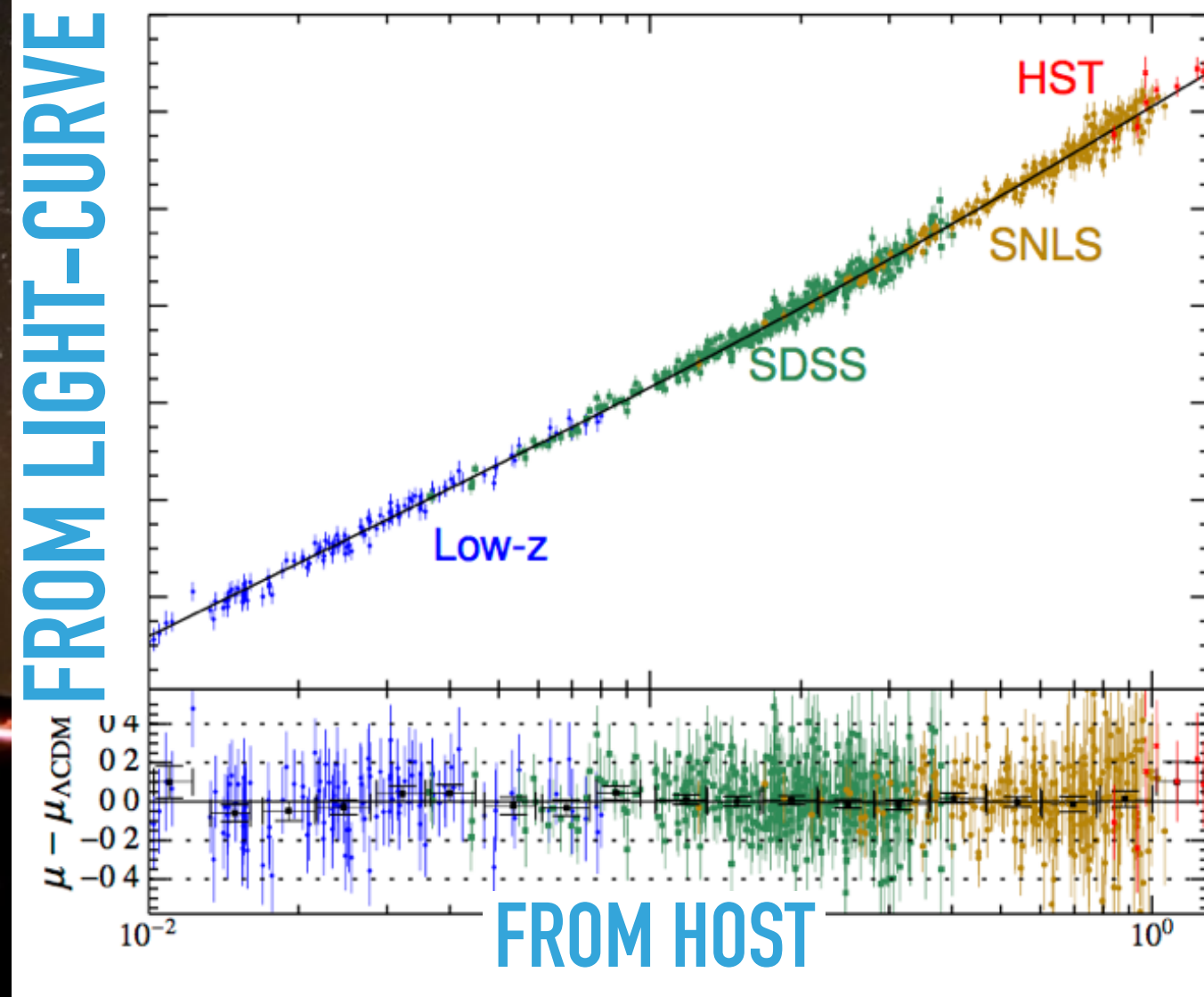
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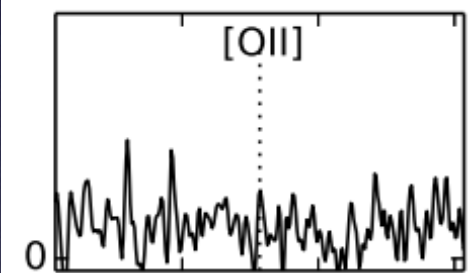
DES-SN: SPECTROSCOPIC FOLLOW-UP: THE WAY FORWARD

OZDES: REDSHIFTS FOR EVERYTHING

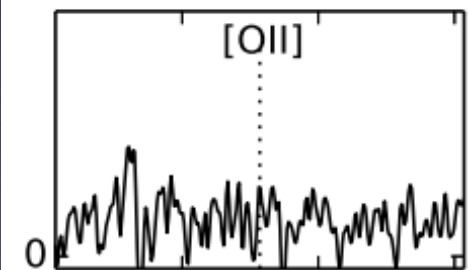
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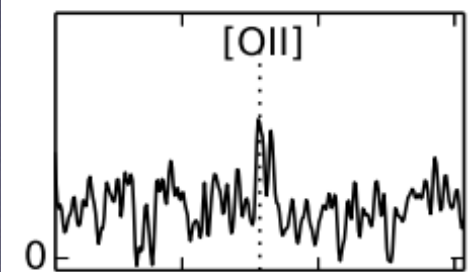
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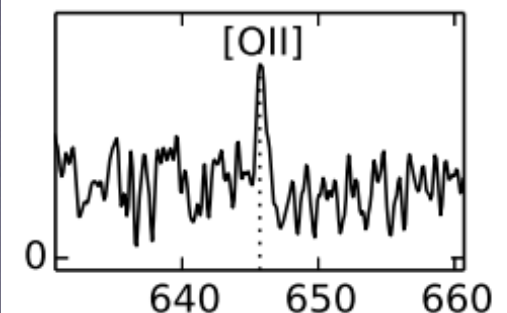
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Exp=32400 s

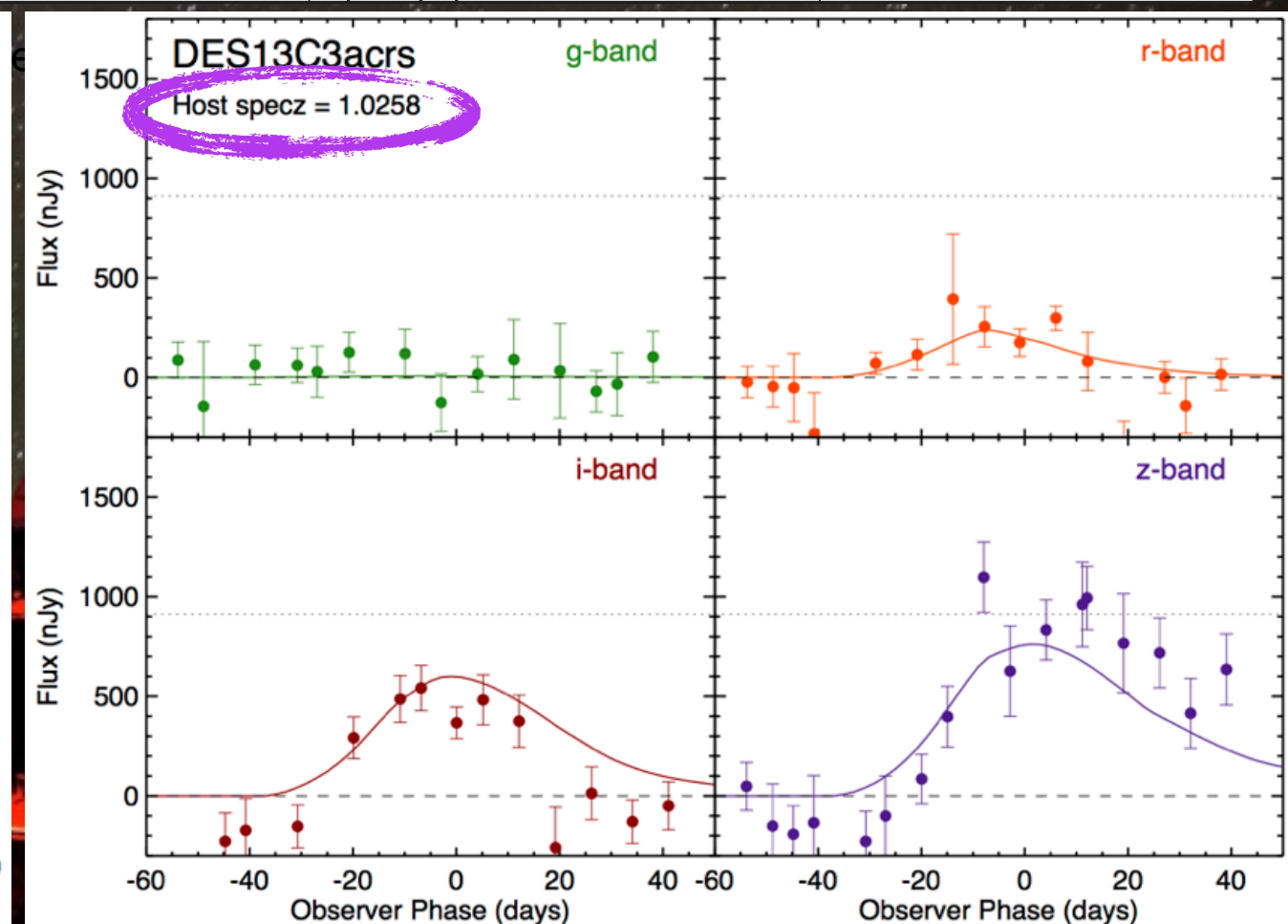
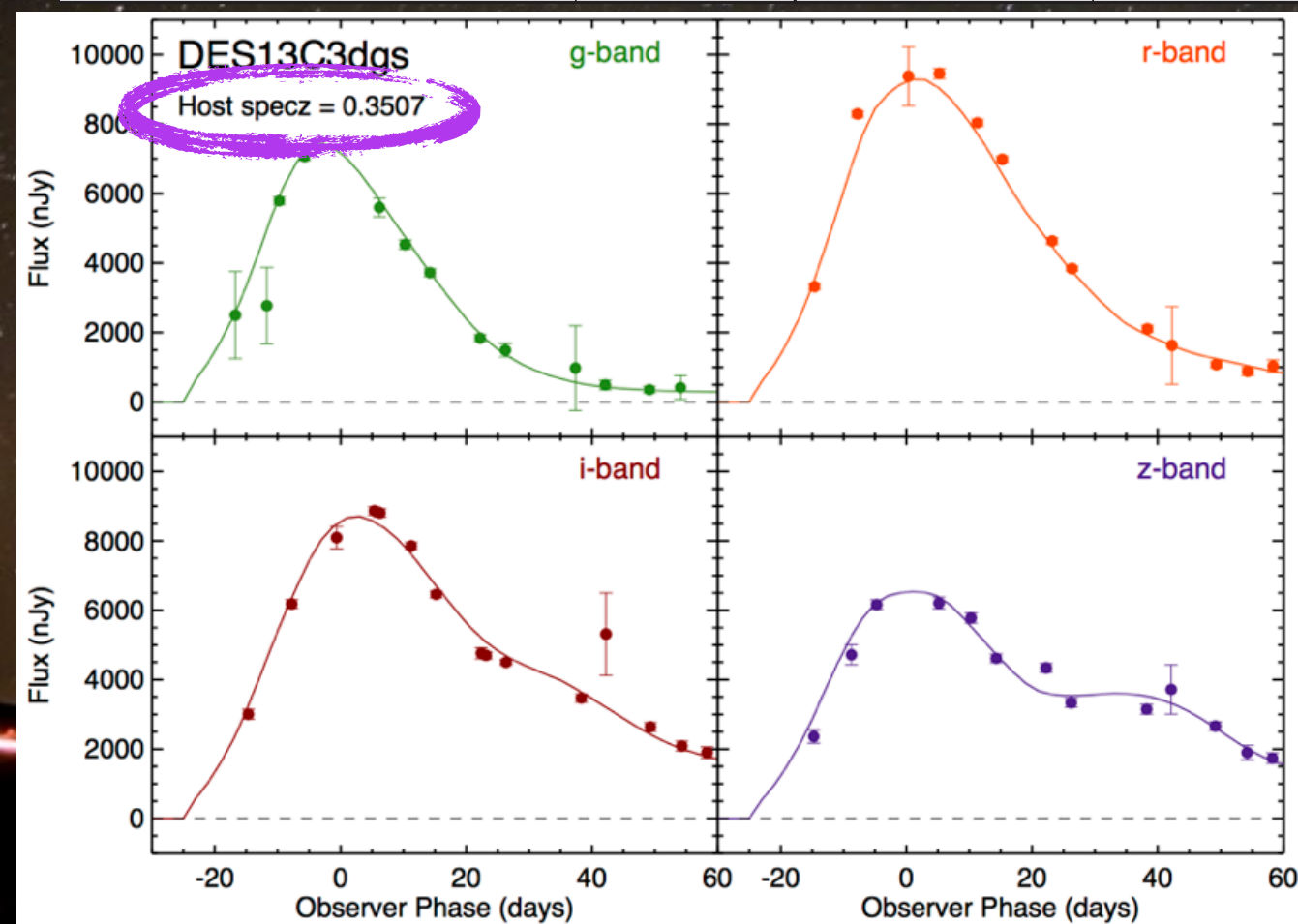
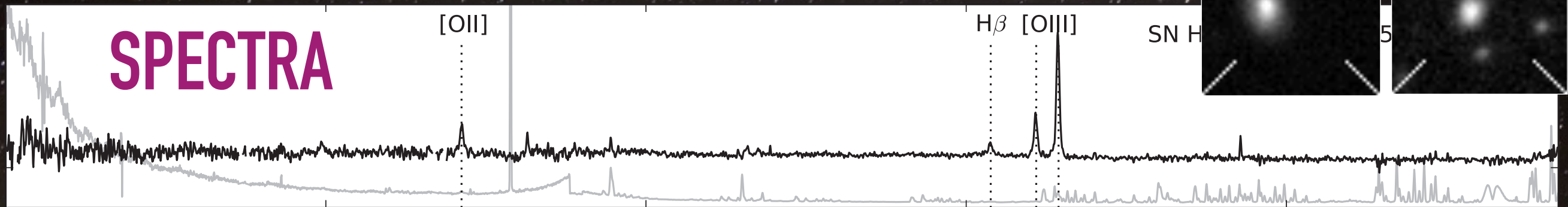


Exp=51600 s



DES-SN: SPECTROSCOPIC FOLLOW-UP: THE WAY FORWARD

DOES IT WORK?



- *Chris L will explain all in <10!, but this takes 5 years.....*

DES-SN: SPECTROSCOPIC FOLLOW-UP: IN THE MEANTIME...

LOTS OF SUPERNOVA, TOO FEW (BUT STILL LOADS OF) TELESCOPES

	Instrument	Time	S	Targets	DESY4+
AAT	AAOmega/2dF	48n	Y123	1027/544	52n over Y4+Y5
VLT	X-Shooter	14n+13h	Y23	91/89	7n + 7n (2016/7:A/B)
Magellan	LDSS-3, IMACS	14.5n	Y23	88/86	6n (2016B)
SALT	RSS	79h	Y123	52/44	17h (to Oct31)
Gemini	GMOS-N/S	53.2h	Y3	32/29	95.1h (2016B) LPP: 99h 2017B
MMT	BCS	8n	Y23	31/28	2n (2016B)
Keck	LRIS, DEIMOS	6n	Y123	23	2n (2016B)
GTC	OSIRIS	55h	Y123	19/18	12h (2016B)

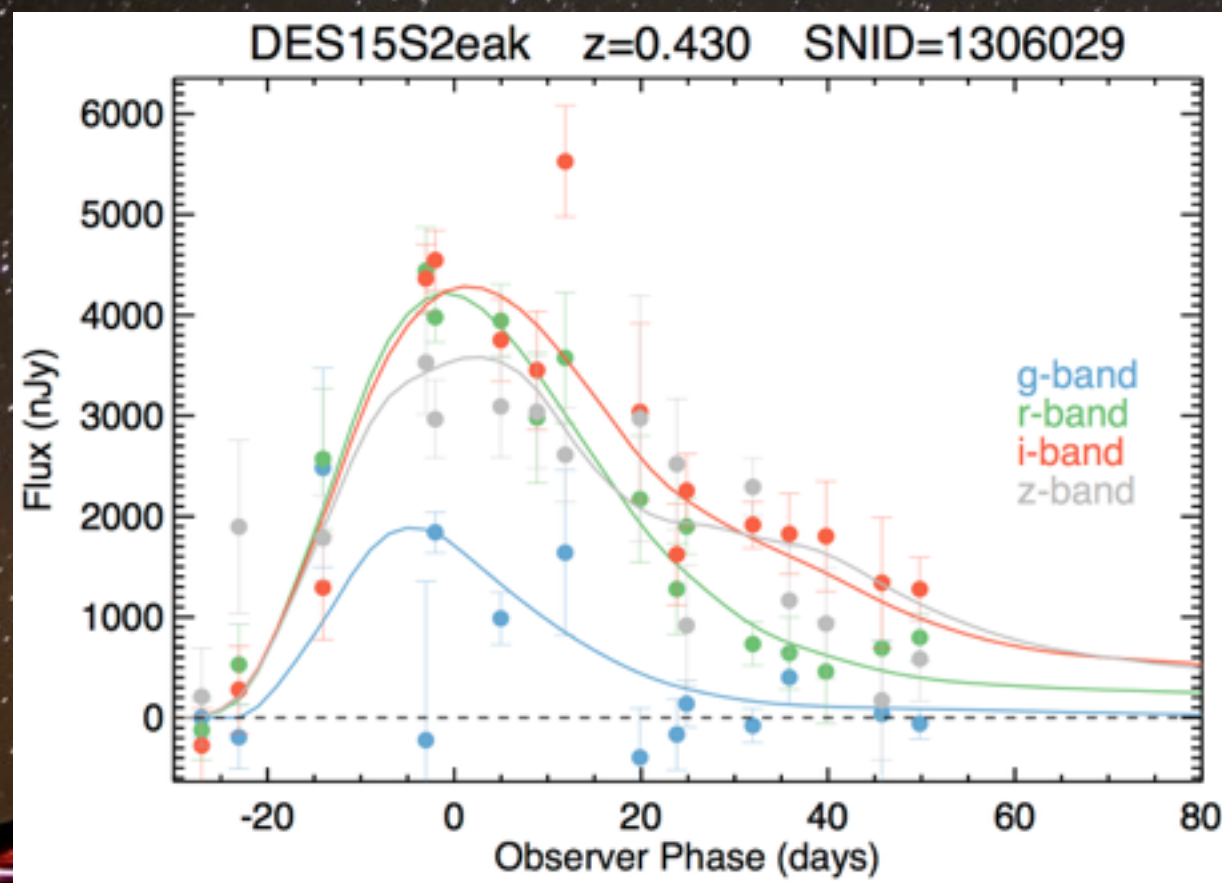
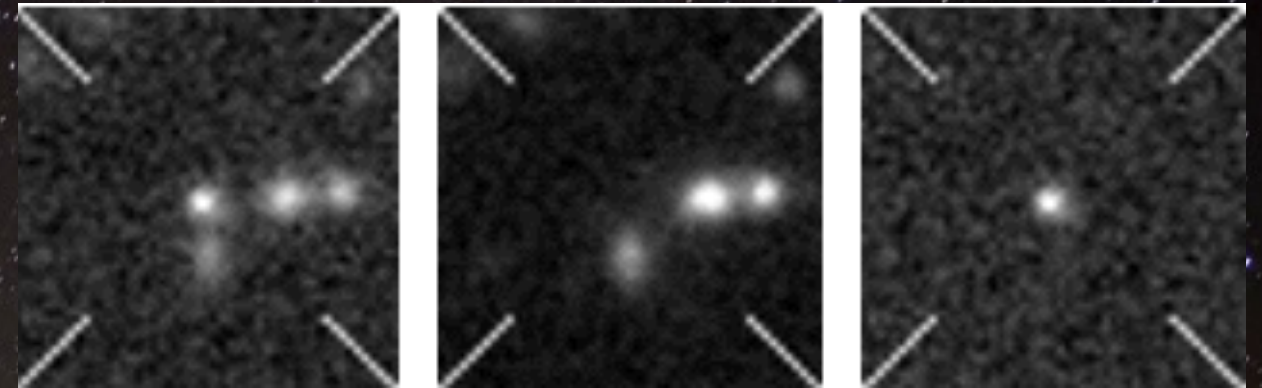
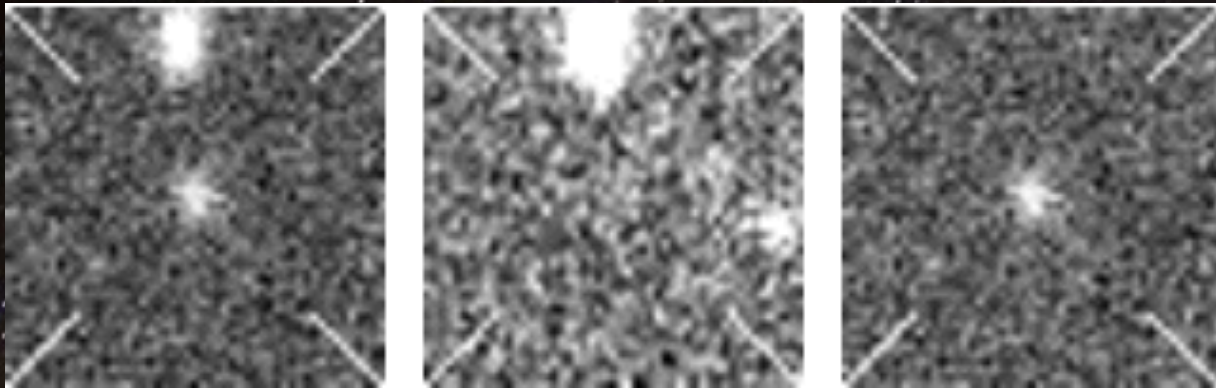
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LOTS OF SUPERNOVA, TOO FEW (BUT STILL LOADS OF) TELESCOPES

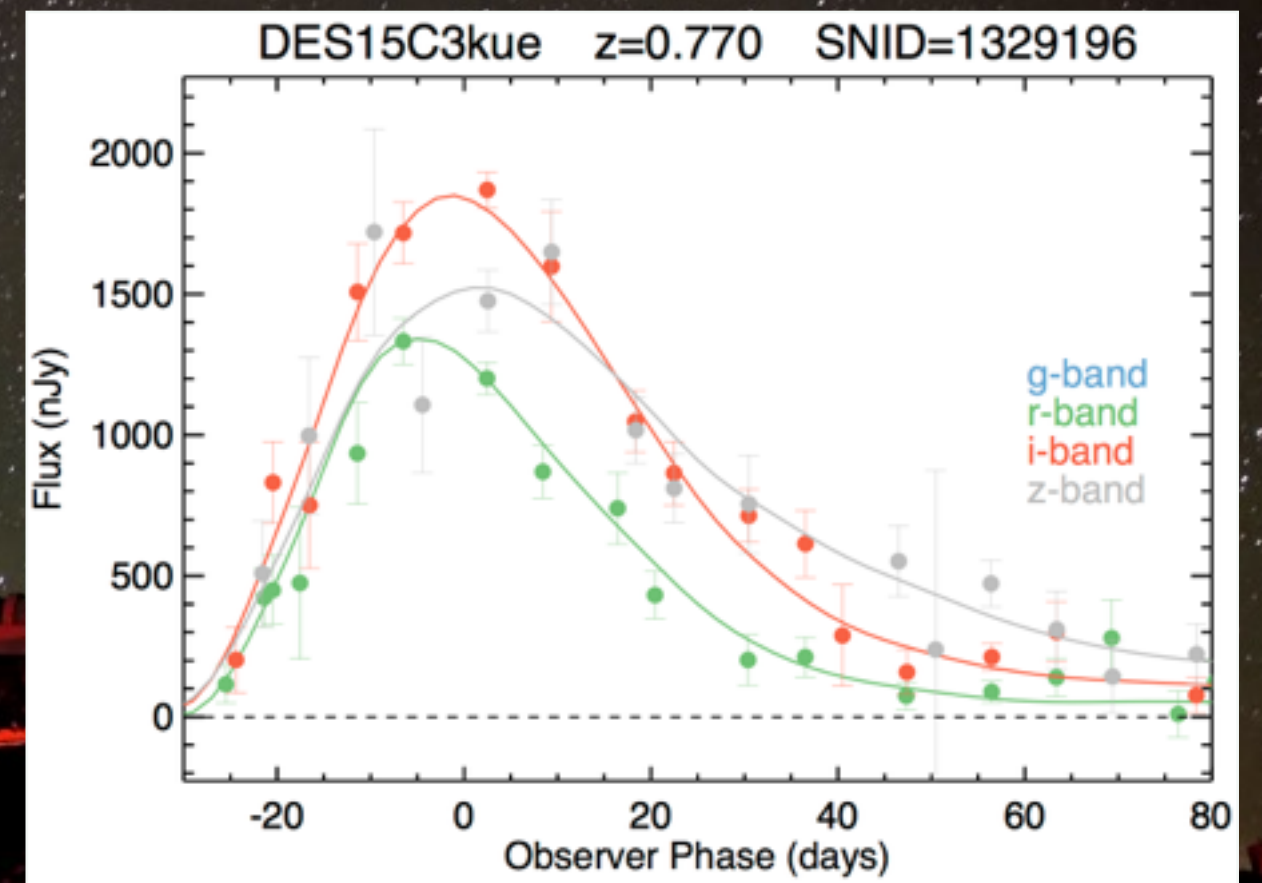
	Instrument	Time	S	Targets	DESY4+
AAT	AAOmega/2dF	48h	~Complete	$z < 0.3$	52n over Y4+Y5
VLT	X-Shooter	14h	Faint hosts	& high-z	7n + 7n (2016/7:A/B)
Magellan	LDSS-3, IMACS	37h	Representative	& mid-z	6n (2016B)
SALT	RSS	79h	~Complete	$z < 0.3$	17h (to Oct31)
Gemini	GMOS-N/S	22h	Representative	& high-z	95.1h (2016B) LPP: 99h 2017B
MMT	BCS	6h	Representative	& mid-z	2n (2016B)
Keck	LRIS, DEIMOS	6n	Weirdos	& high-z	2n (2016B)
GTC	OSIRIS	55h	Weirdos!	19/18	12h (2016B)

- Spectroscopy needed, at least, to test photometric classification...

DES-SN @ AGE 3



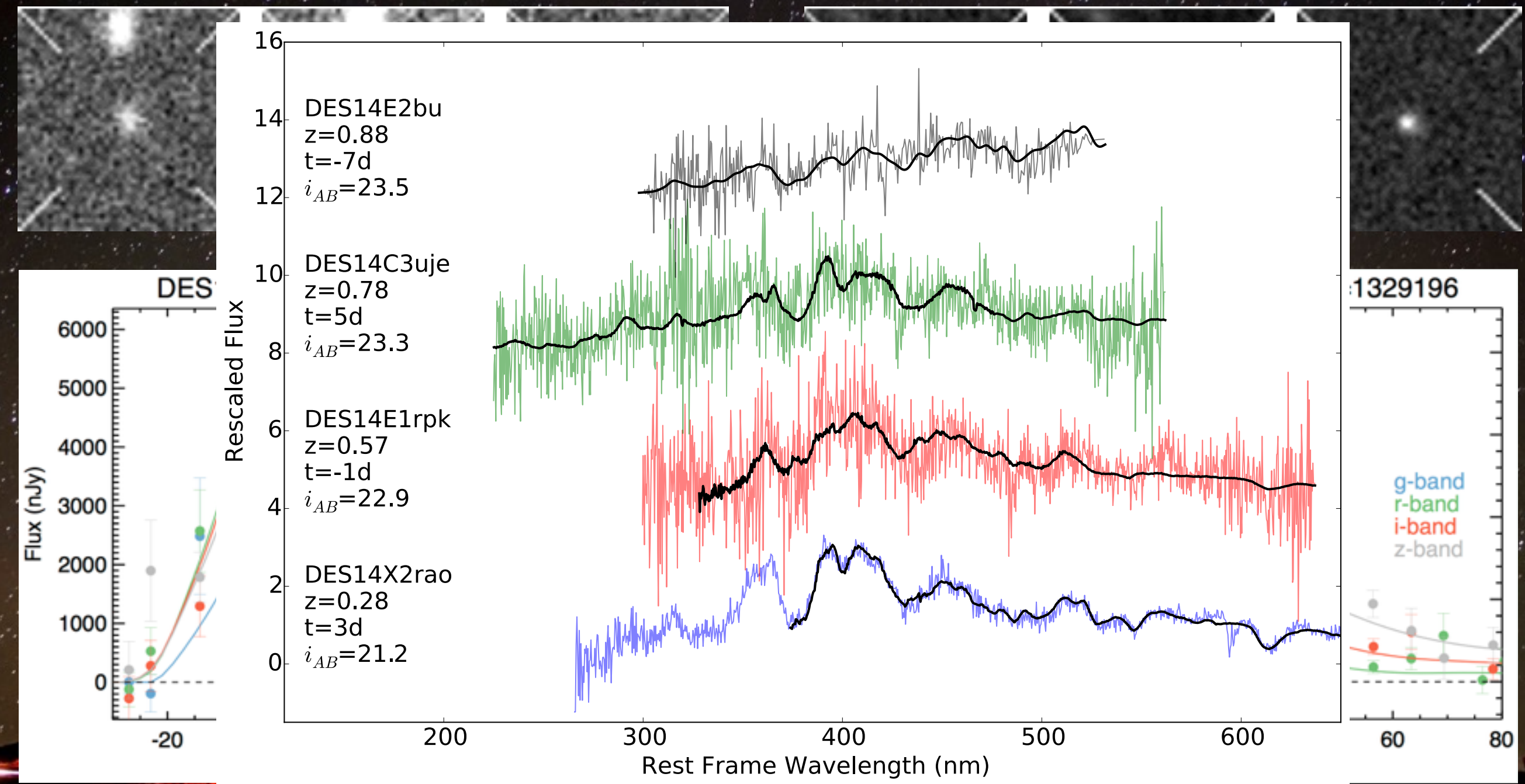
lowish-z



high-z

- *We can find supernova...*

DES-SN @ AGE 3



lowish-z

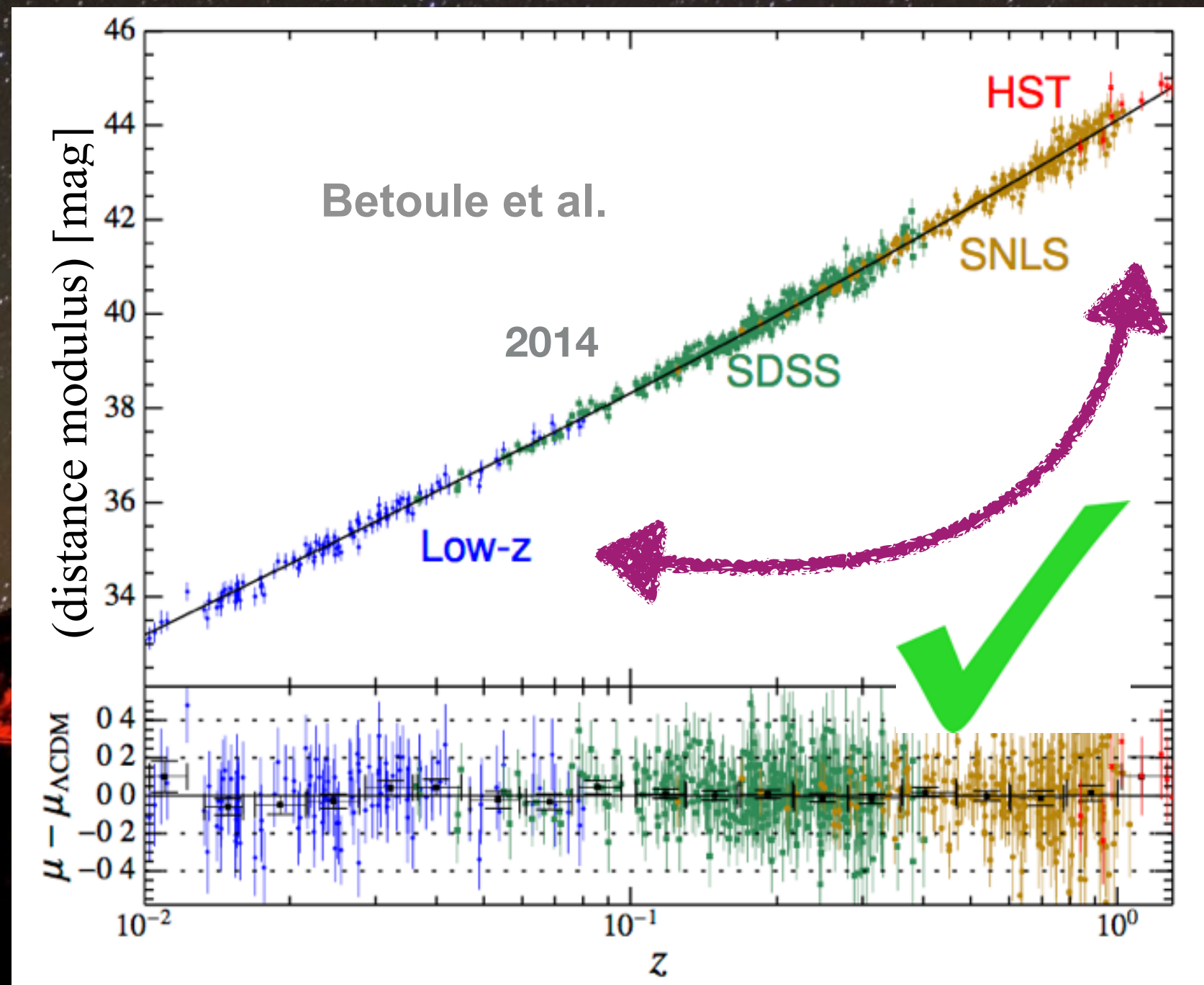
high-z

- We can find supernova... and confirm that they are Ia's!

DES-SN @ AGE 3 - IN SUMMARY

TO RE-CAP:

DES: "ALL OF THIS AND MORE WITH ONE SURVEY"

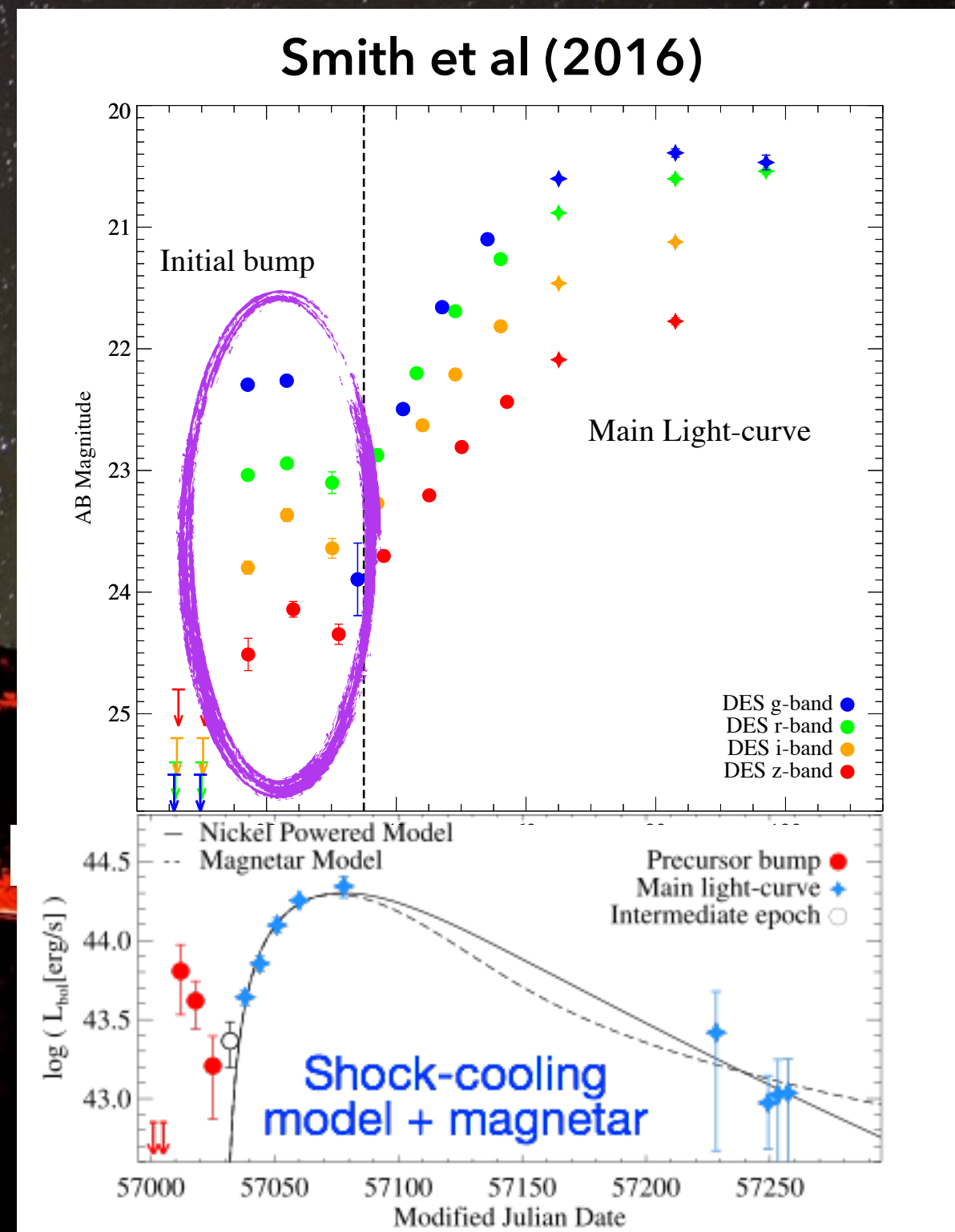


DES-SN @ AGE 3 - AN ASIDE

TRANSIENT SURVEYS = COOL STUFF: SUPER-LUMINOUS SN

V. RARE & IN NEED OF UNDERSTANDING

- 13 total; 6 @ $z > 1$ & new physics!



PLENTY OF STEPS BEFORE WE REACH “W”



TRANSIENT DETECTION:

KESSLER ET AL (DIFFIM) + GOLDSTEIN ET AL (ML)



CALIBRATION:

BIGGEST SYSTEMATIC (CURRENTLY): LASKER ET AL



PHOTOMETRY:

“SCENE MODELLING”: BROUT ET AL (IN PREP)



HOST IDENTIFICATION / TARGETING:

GUPTA ET AL (IDENT) + YUAN ET AL (OZDES)



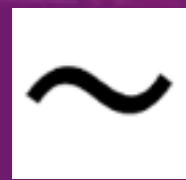
CLASSIFICATION:

PSNID: SAKO ET AL &/OR KESSLER+SCOLNIC



SPECTROSCOPY / BIASES:

D’ANDREA ET AL



HOST GALAXY PROPERTIES:

SMITH ET AL (IN PREP)



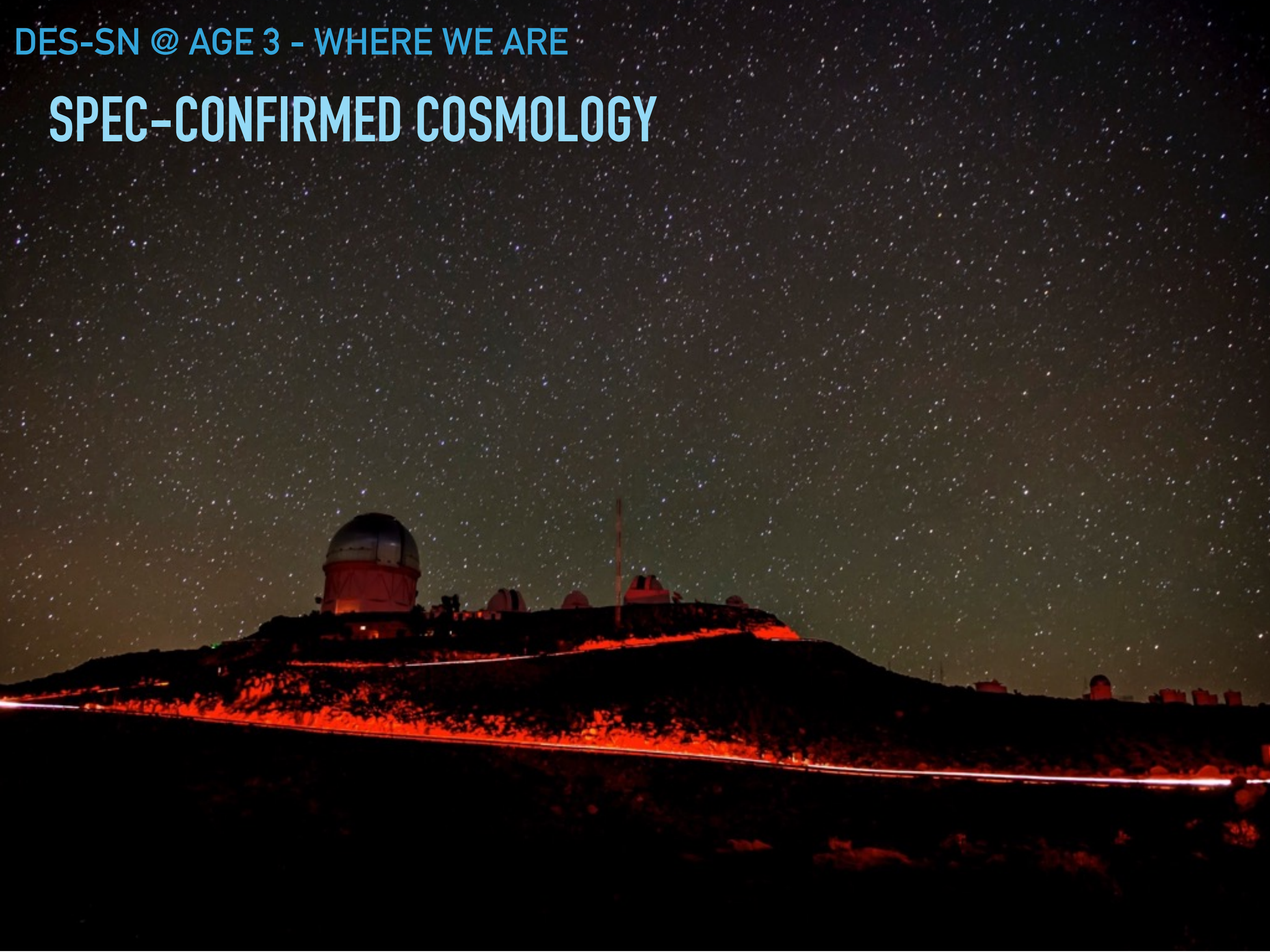
COSMOLOGICAL ESTIMATION:

BAYESIAN-NESS: HINTON ET AL (IN PREP) + OTHERS

- *Some outstanding issues, but things are moving forward....*

DES-SN @ AGE 3 - WHERE WE ARE

SPEC-CONFIRMED COSMOLOGY



DES-SN @ AGE 3 - WHERE WE ARE

DO WE KNOW HOW WE GOT HERE?

- *Cosmology to come shortly!*

- *Data broadly matches simulation and “evolves” with redshift as expected - this still needs to be corrected for!!!*

DES-SN @ AGE 3 - WHERE'S NEXT: PHOTOMETRIC CLASSIFICATION

DES Y1-3: >1500 SNIA'S; 900+ CURRENTLY WITH SPEC-Z

Currently modelling biases, selection functions & contamination

DES-SN @ AGE 3 - MY SUMMARY

- ▶ DES-SN is great!
 - ▶ 3 years down, 2 to go!
 - ▶ 250+ spectroscopically-confirmed SNe Ia
 - ▶ 13 SLSNe, inc. highest redshift ever!
 - ▶ 1500+ photometrically-classified SNe Ia
 - ▶ Twice as many as largest published sample!
 - ▶ we just need redshifts & time ;)
 - ▶ New techniques for calibration, photometry, classification & cosmology
 - ▶ Stay tuned!

DES-SN @ AGE 3 - FEEL FREE TO CLAP!



CHEERS MATE!