

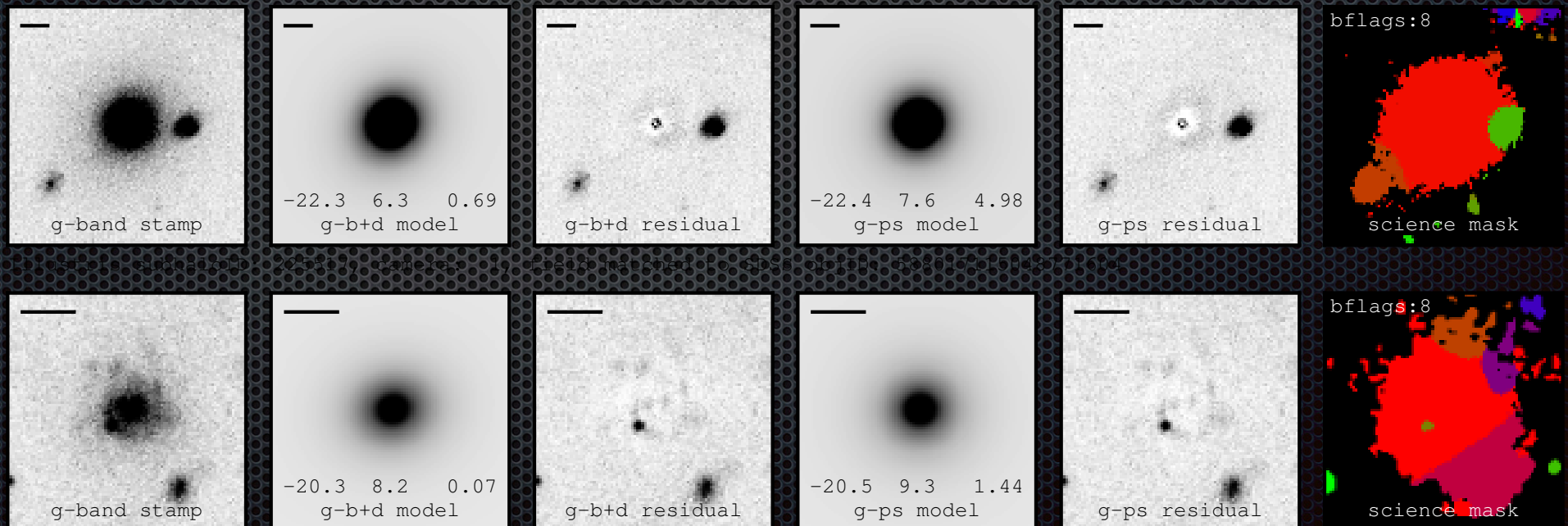
Observational Properties of Simulated Galaxies

Computational Cosmology as seen through a Telescope

CAASTRO: The Changing Face of Galaxies

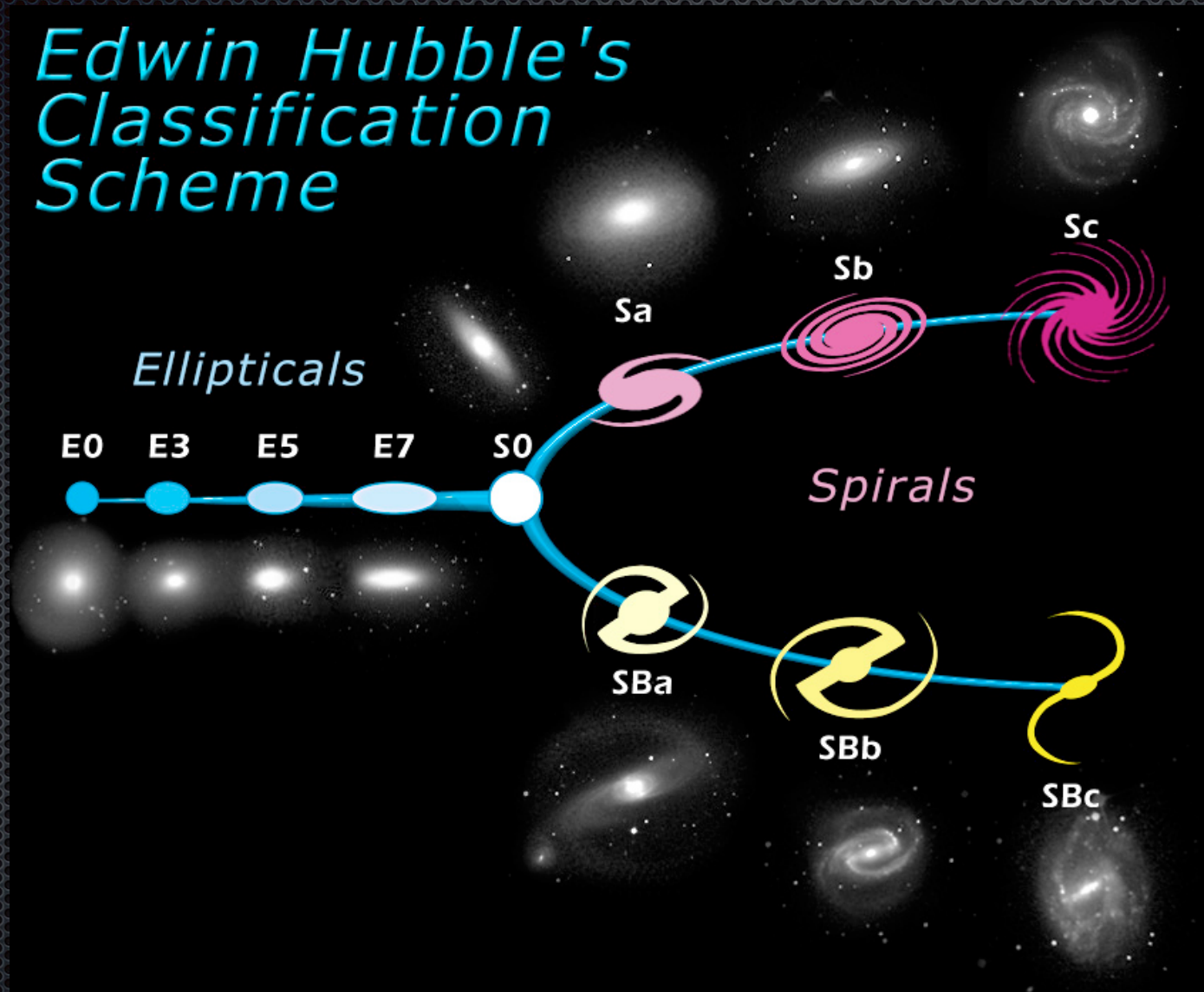
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Luc Simard,
Sara L. Ellison



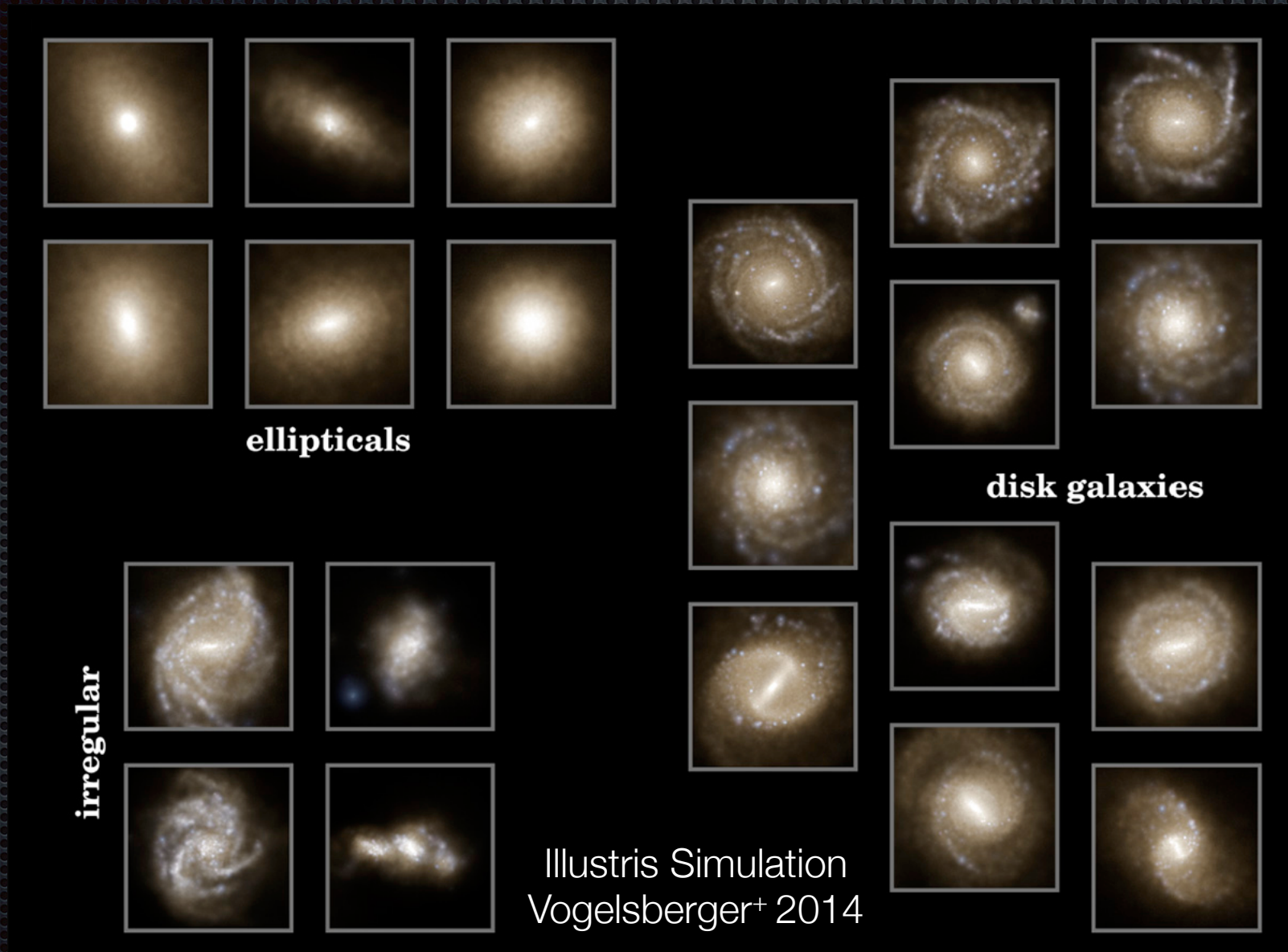
✦ Introduction

Galaxies demonstrate diverse morphological structures

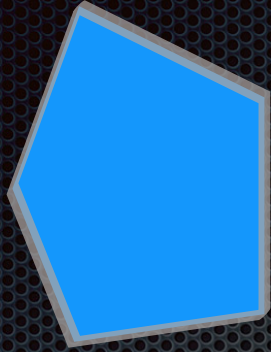


✦ Introduction Hydrodynamical Simulations

Goal: reproduce the observed properties & morphologies of galaxies



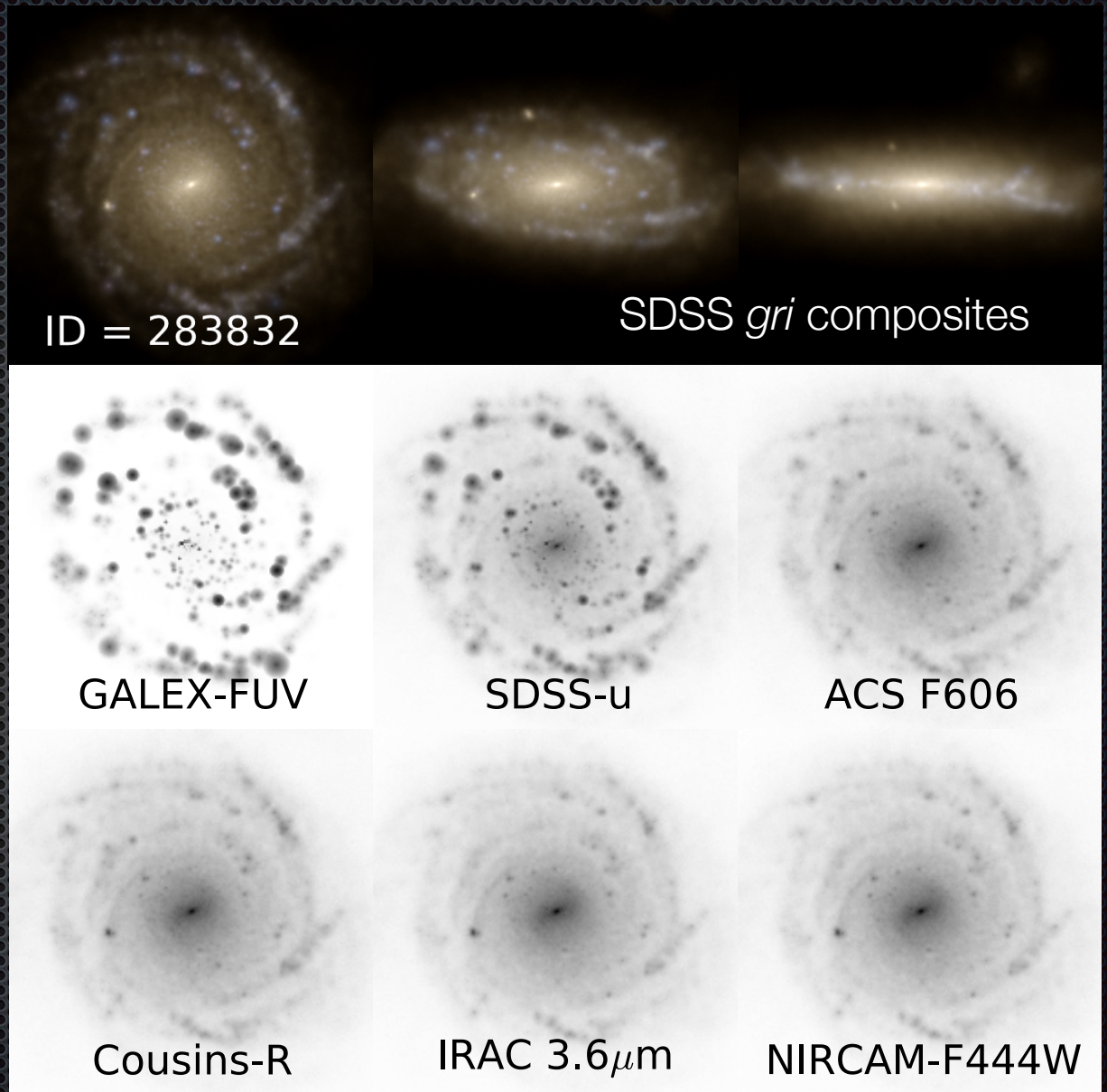
✦ Introduction Synthetic images



gas cell



star particle + SED



✦ Methods

Galaxies from $z=0$ of the Illustris Simulation with SDSS Realism

Key components:

subhaloID: 312287
camera: 1

r-band synthetic image

PSF convolution
+ Poisson noise

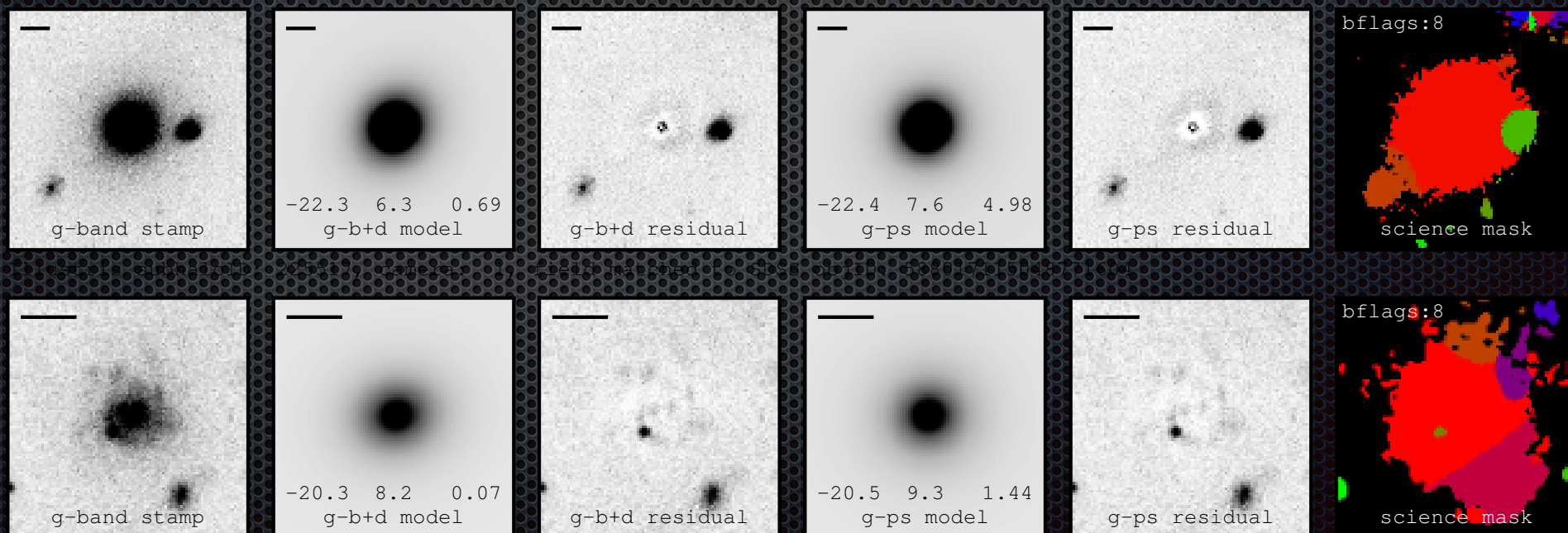
insertion into SDSS sky

**Goal: derive galaxy properties consistently
with observed galaxies**

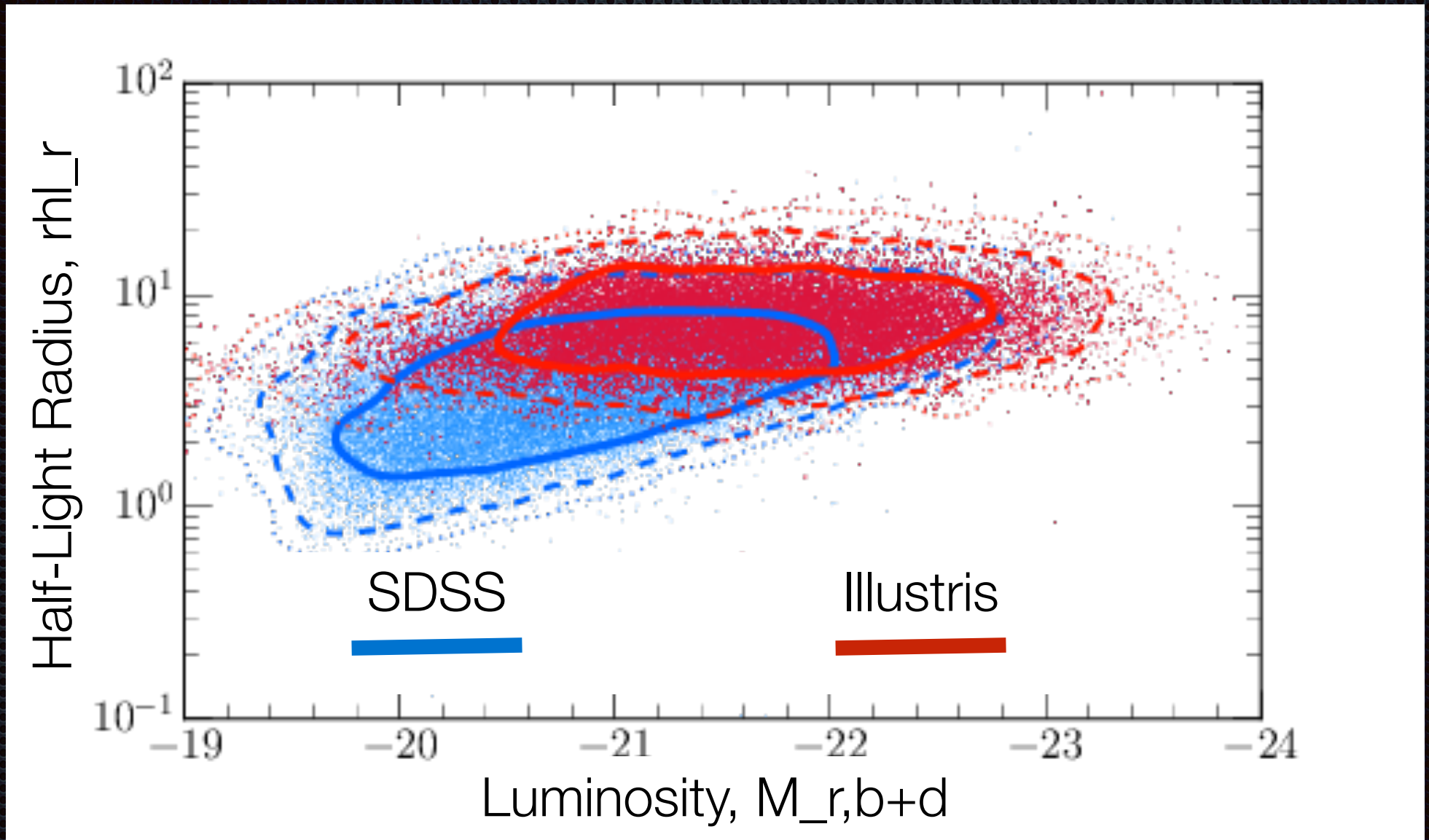
Bottrell et al. 2016, MNRAS submitted

(email connor.bottrell@gmail.com for early access to catalogs)

- Public catalogs of 2D parametric decompositions of Illustris galaxies with full SDSS realism (bulge+disc and pure sersic)
- Processed with the *same* decomposition pipeline as catalogs by [Simard et al. 2011](#)

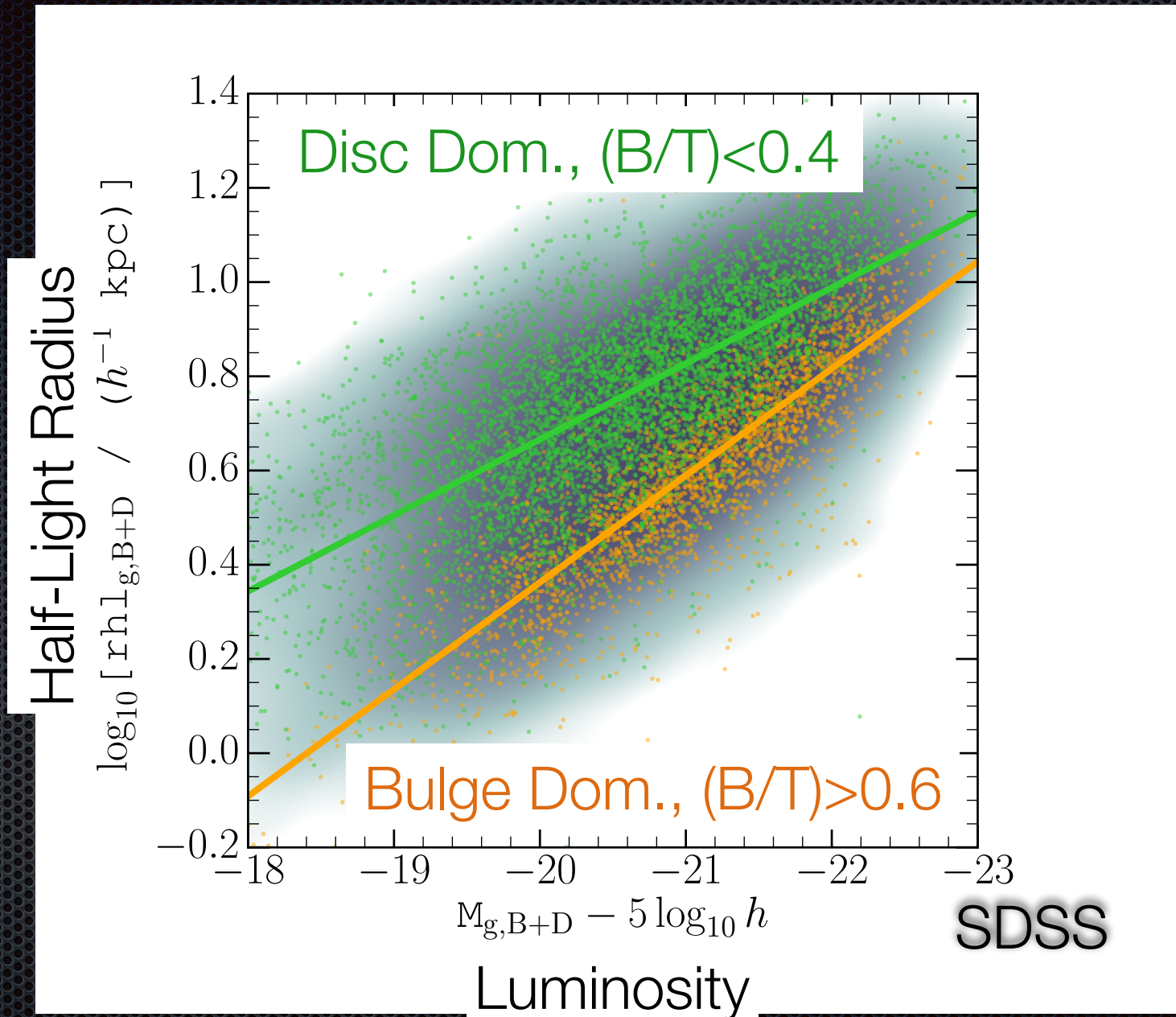


✦ Size Luminosity Relation



Illustris galaxies (red): larger, brighter, shallower slope for same masses

Size Luminosity Relations of Discs/Spheroids

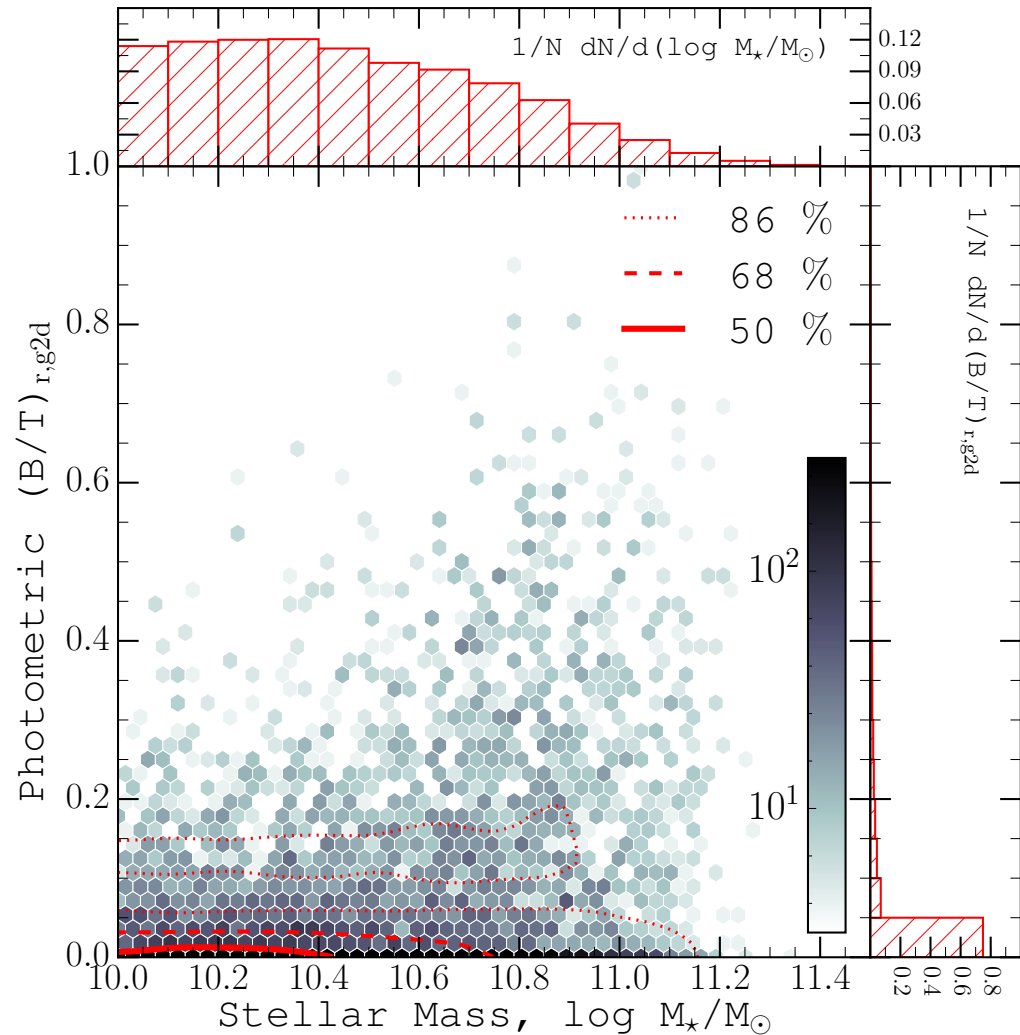
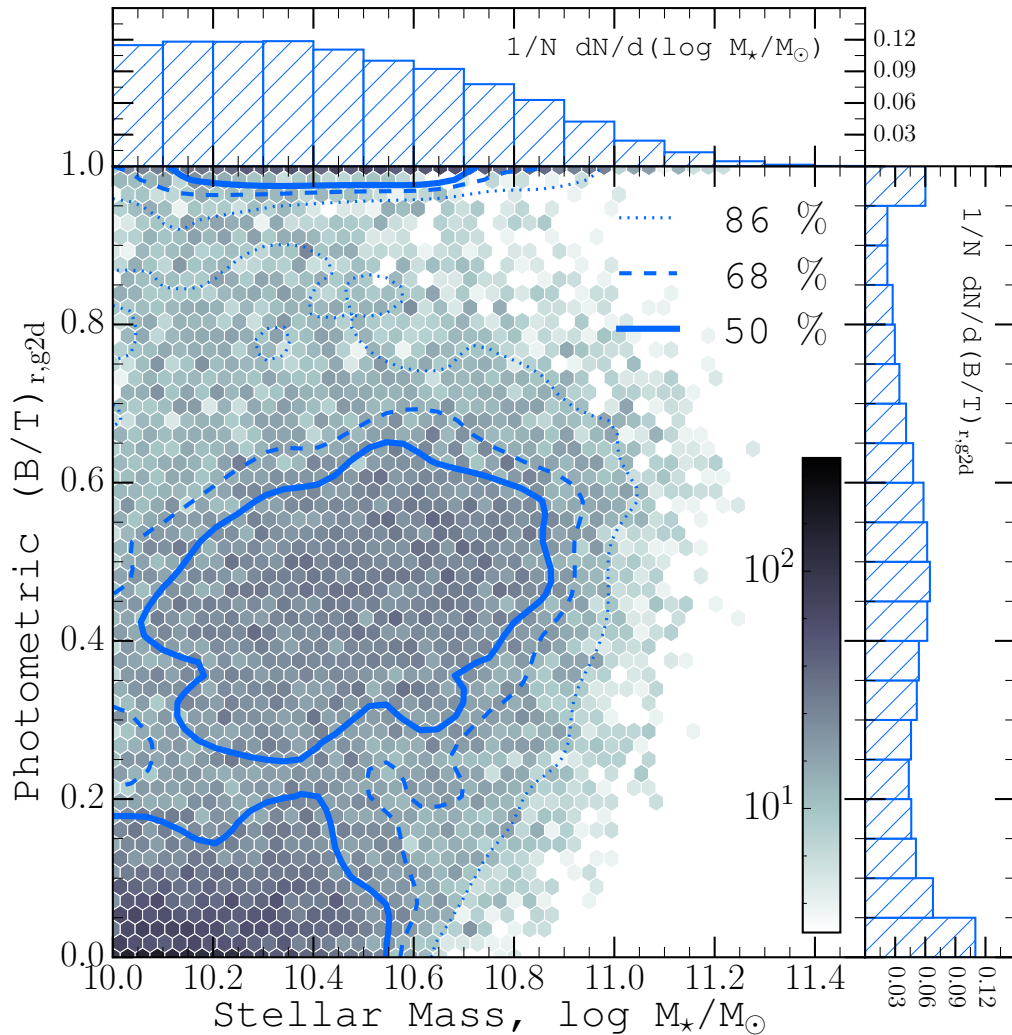


Size-luminosity relation is sensitive to morphology

✦ Bulge to Total Fractions

SDSS

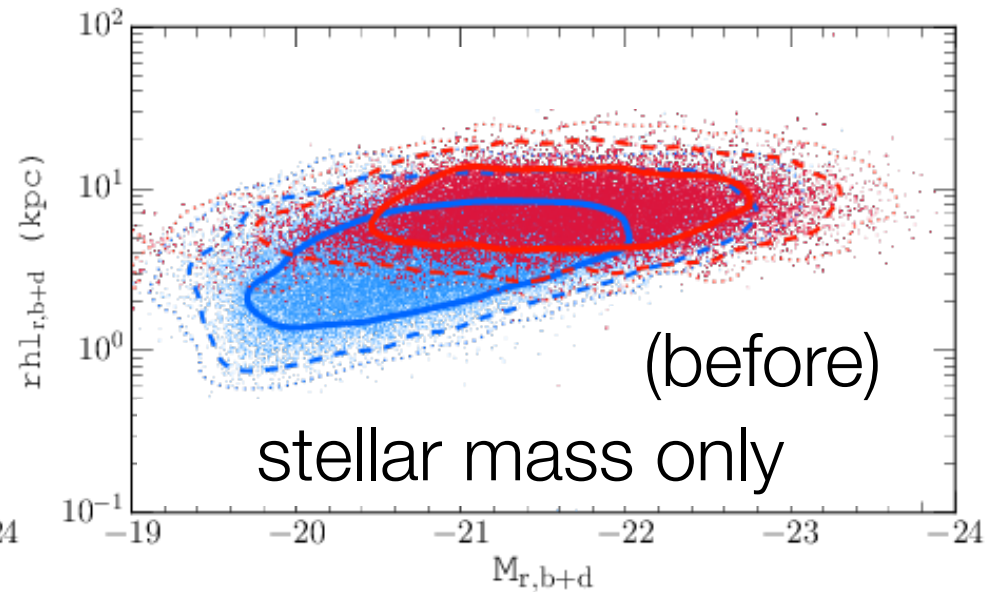
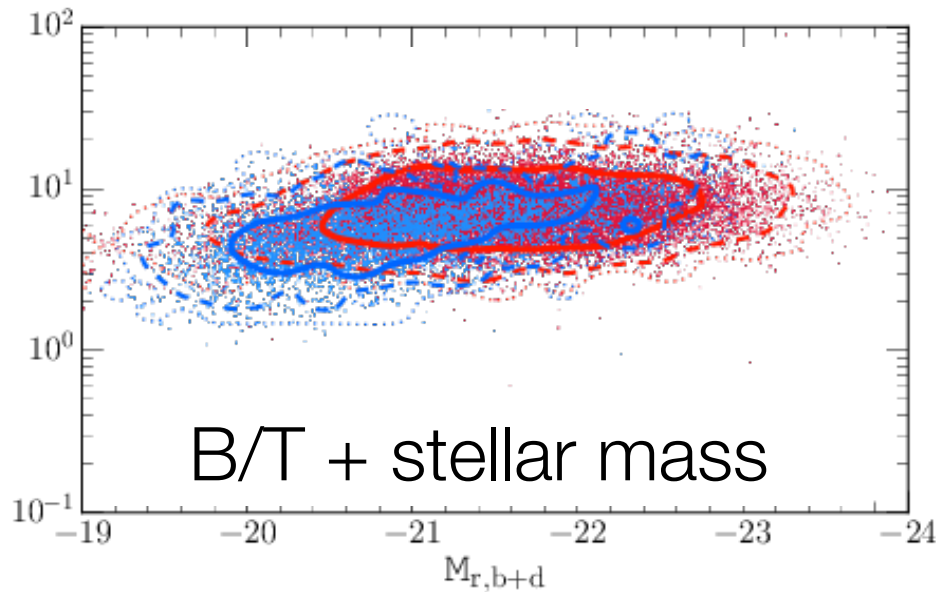
Illustris



In samples matched by mass and (low) redshift, Illustris is bereft of bulges where real galaxies have diverse morphologies

✦ Morphology Matched Size Luminosity Relation

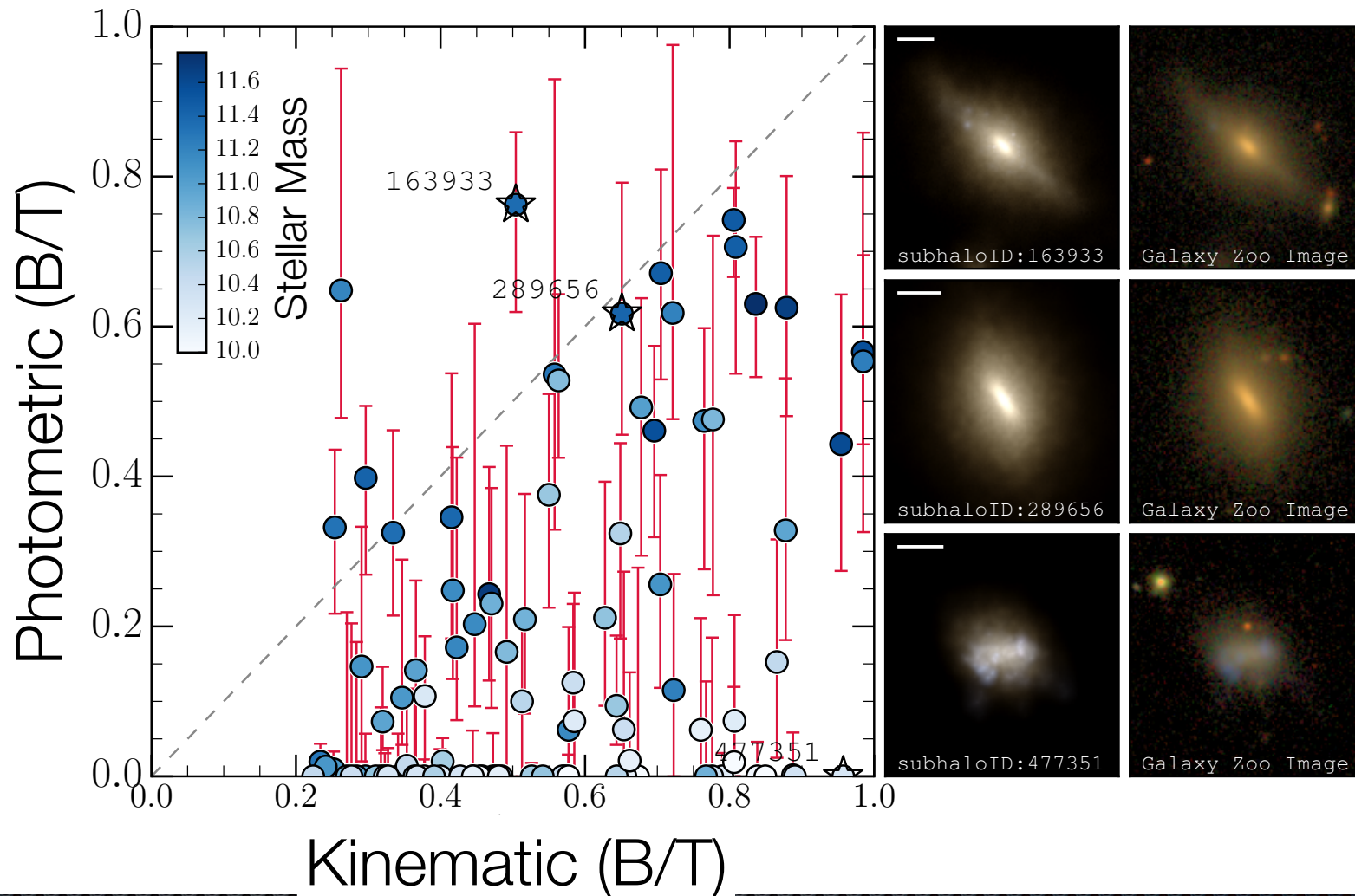
Half-Light Radius (kpc)



Luminosity, $M_{r,b+d}$

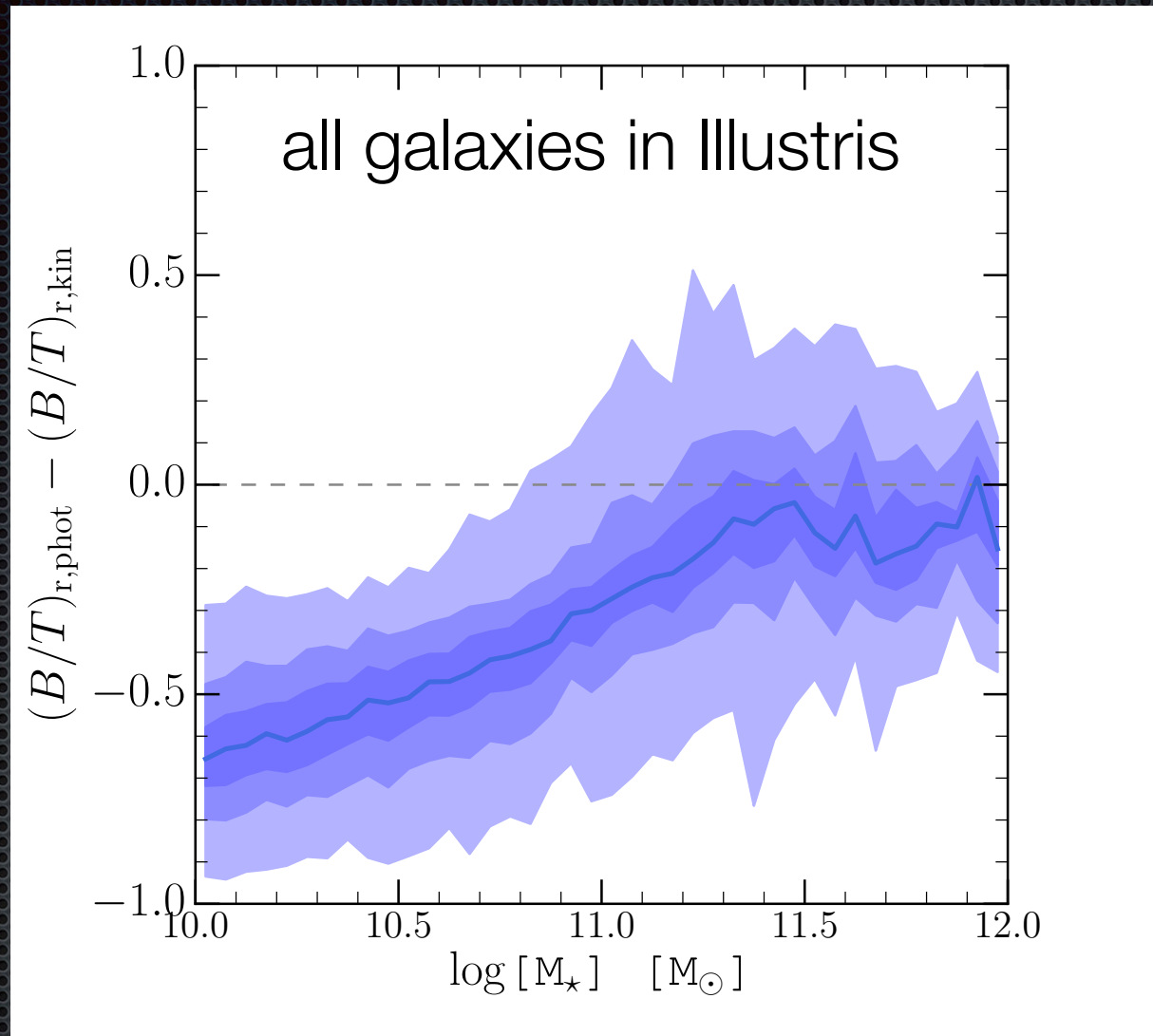
Morphology matching remedies broad differences in the SL relation, but is left with primarily the disc SL relation

Photometric and Kinematic Bulge Fractions



Photometric B/T systematically lower than kinematic B/T derived from stellar orbits. Lower mass galaxies most strongly affected.

✦ Photometric and Kinematic Bulge Fractions



Photometric B/T systematically lower than kinematic B/T derived from stellar orbits. Lower mass galaxies most strongly affected.

✦ Summary

- ✦ New **image-based comparison method** for observed galaxies and hydrodynamical simulations
 - ✦ **Public Decomposition Catalogs** and characterization of biases' affects on structural measurements

Bottrell et al. 2016, MNRAS submitted. E-mail connor.bottrell@gmail.com for early access to catalogs

- ✦ Difference in Size-luminosity relation driven by a **deficit of bulge-dominated galaxies in Illustris** for galaxies with $\log M^* < 11$ [M_{sun}]
- ✦ Discrepancy between kinematic and photometric B/T

Thank you