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Disturbed Things Come in small packages





Thanks to: Scott Croom Julia Bryant Joss Bland-Hawthorn Kinematics can uncover processes shaping galaxy evolution

- Kinematic classification of SAMI galaxies
- Physical basis for scatter off the Tully-Fisher relation
- What determines asymmetry?



- SAMI Galaxy Survey data cubes
- Hα emission line kinematic maps fit by LZIFU (Ho et al., 2014) – subsample of 816 galaxies
- MY SAMPLE: 532 galaxies after Hα S/N cuts



Krajnovic et al., 2007

KINEMETRY

Acymmetry (n>1)

POWER IN HIGHER MODES INDICATES KINEMATIC ASYMMETRY



Classification using kinemetry



Scatter off the Tully-Fisher Relation (can be physically explained)



Asymmetry is related to stellar mass



Why are asymmetry and mass anticorrelated?

Interactions?

Are low mass galaxies satellites of large galaxies?

Are low mass galaxies qualitatively different?
High gas fraction → turbulence, star formation...



More likely to be

Less likely

kinematically disturbed

What about distance to nearest GAMA: Liske et al., 2015 Brough et al., 2013



Disturbed things trend small

- Stellar mass is more strongly related to asymmetry than distance to nearest neighbour
 - Asymmetry can be caused by turbulence, bursty star formation
 - Interaction is not the primary cause of asymmetry
- Scatter off the Tully-Fisher relation can be explained
 - Scatter is often caused by offset between photometric and kinematic PA

Differences in PA \rightarrow rotation velocity

